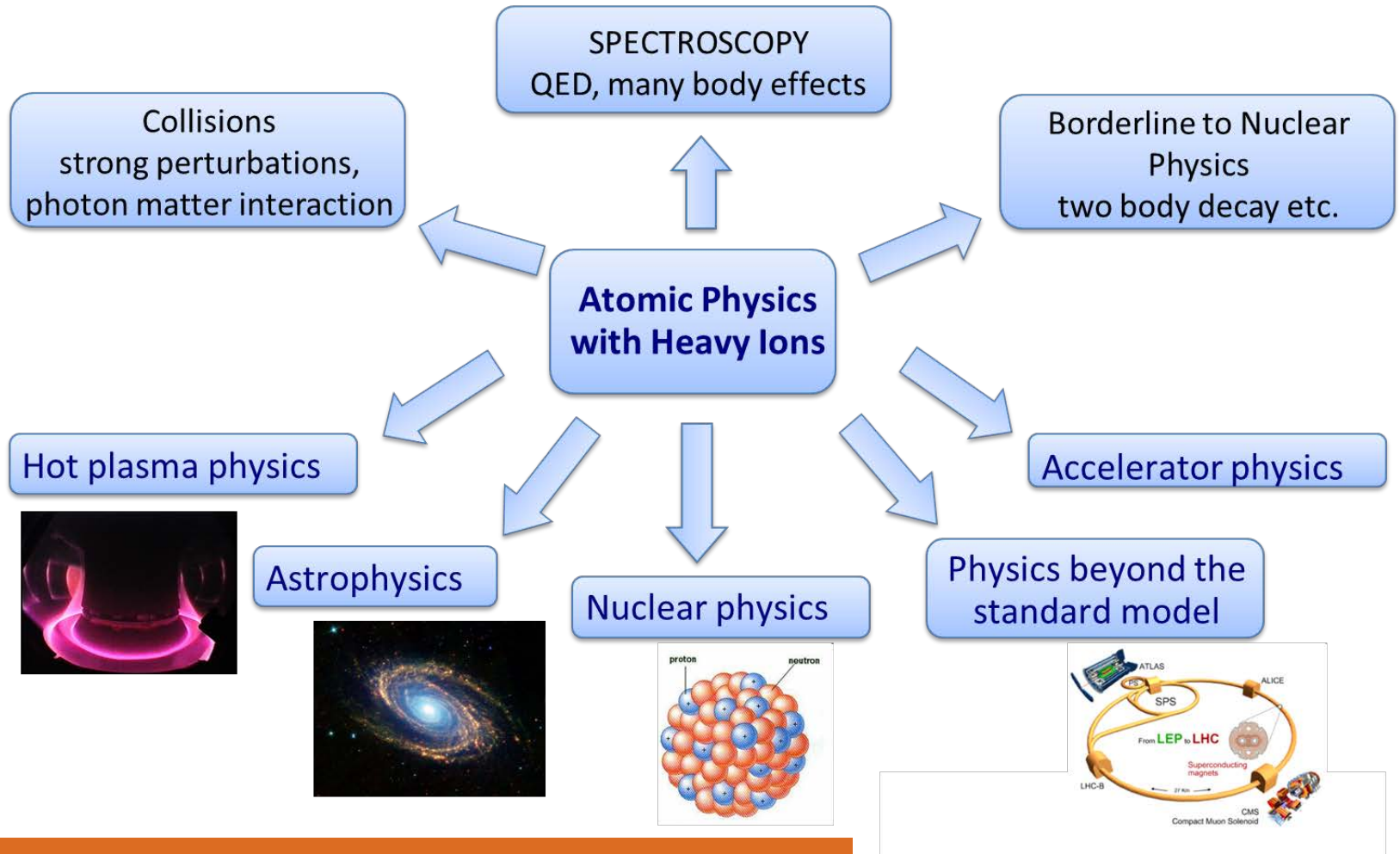


Our Field of Research: Atomic Physics



Highly Charged Ions / Strong EM Fields



Galaxies



AGNs

Coronae

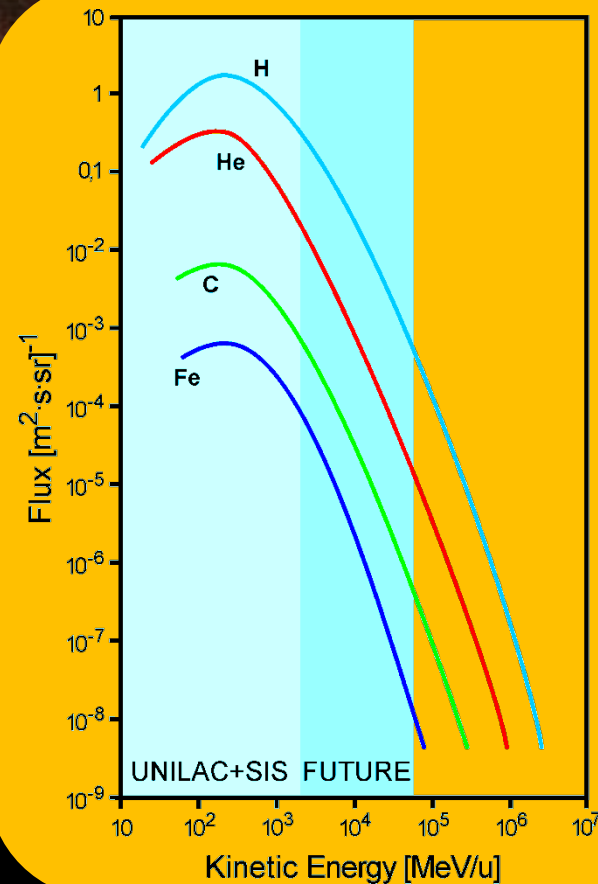
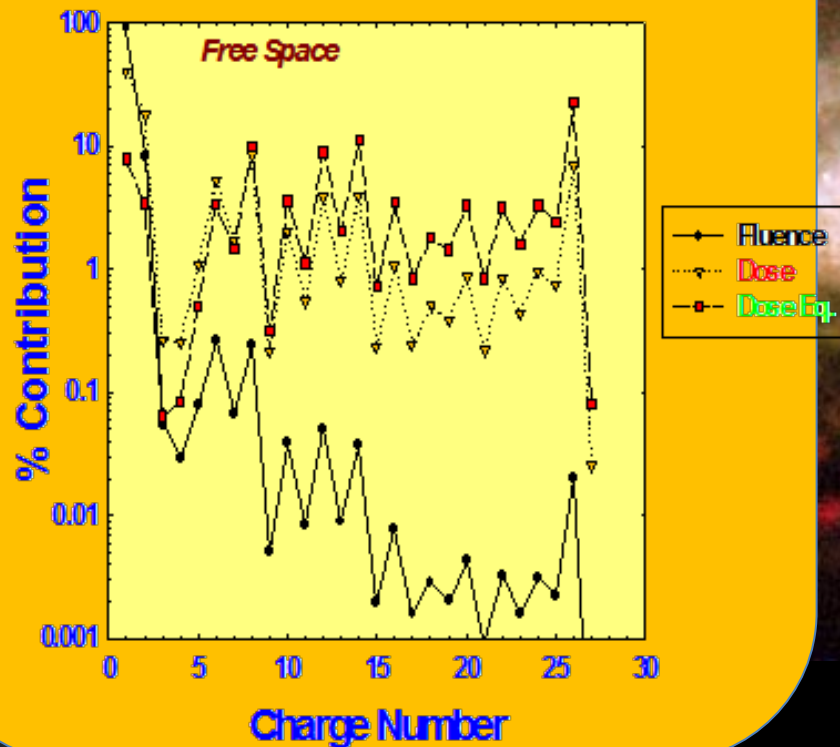


Comets

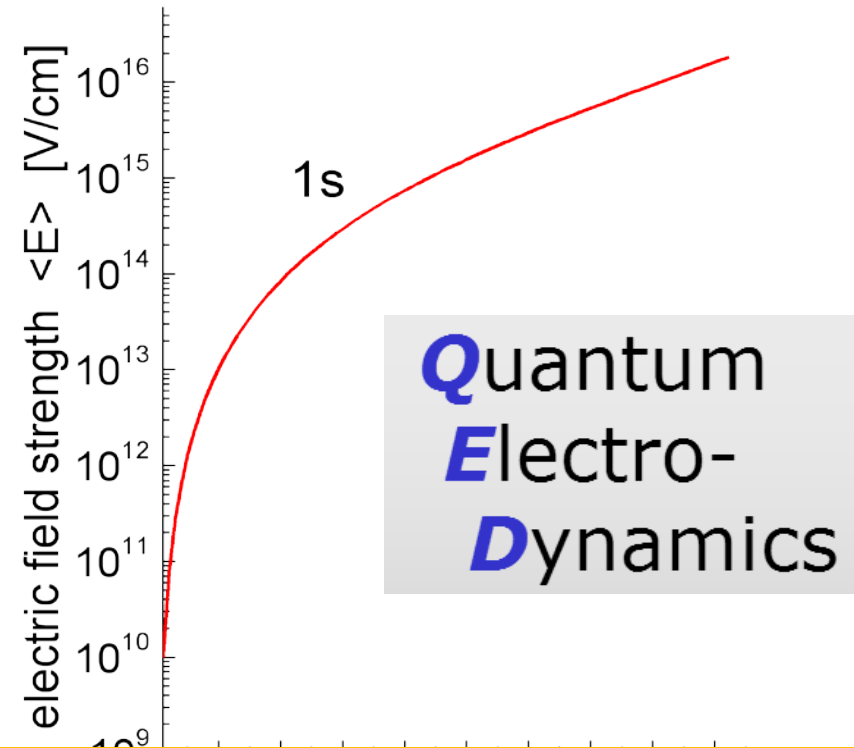
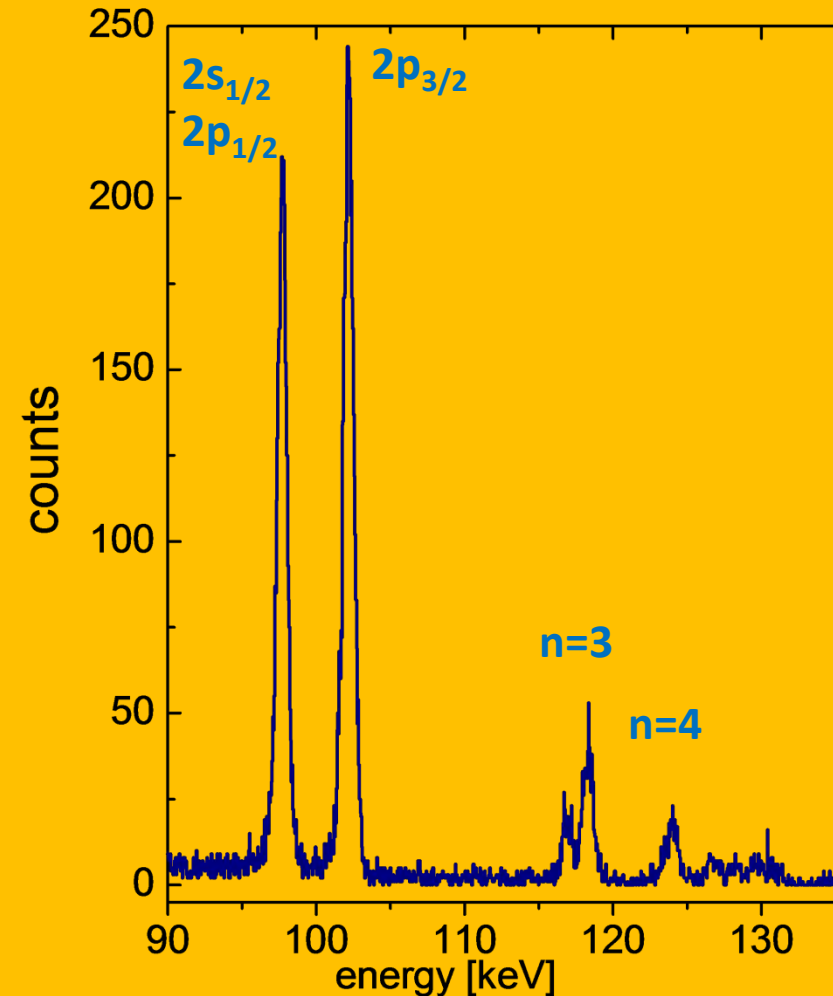
- Ionization and Particle Production Phenomena
- Radiative Processes

Relativistic Energies: Galactic Cosmic Radiation (GCR)

GCR Charge Contributions



Structural properties of heavy ions

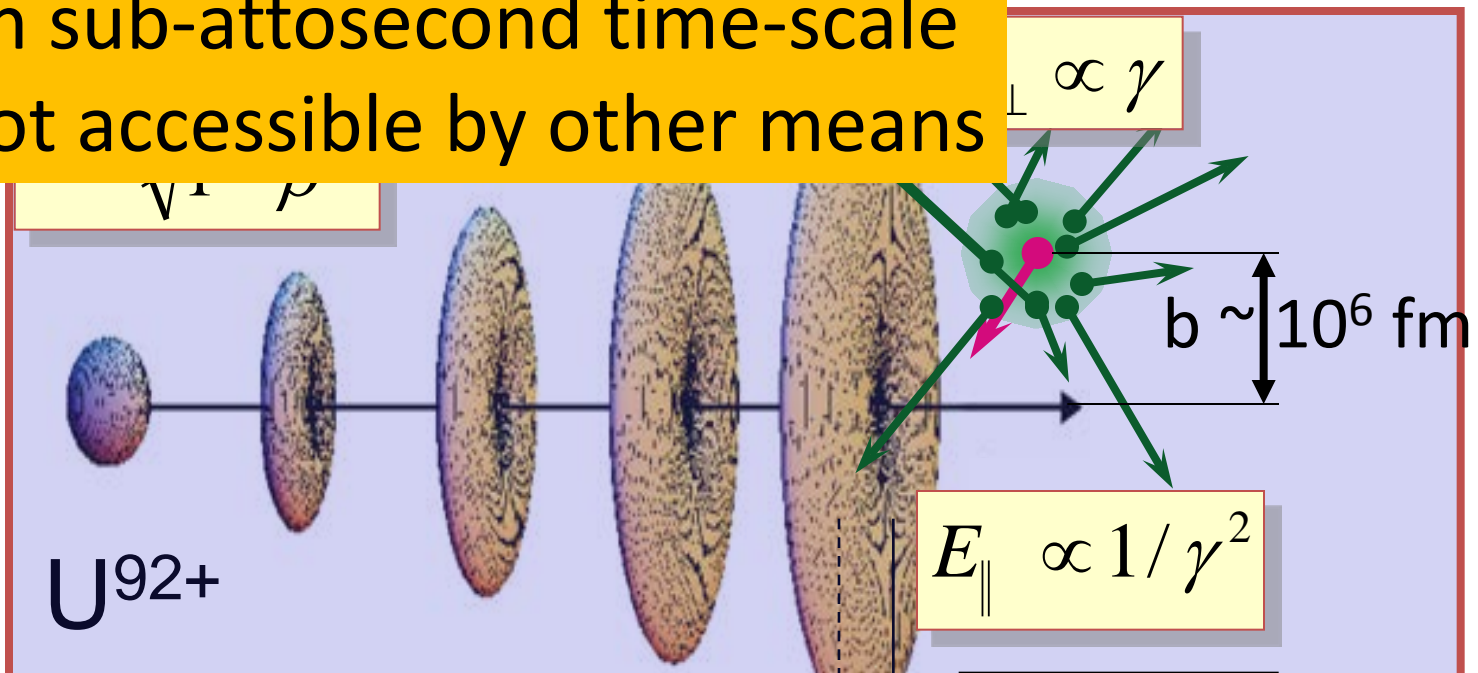


QED, the most precise theory
the non-perturbative regime
- not well-known up to now!

Extreme Dynamic Fields

Explore correlated electron dynamics

- on sub-attosecond time-scale
- not accessible by other means

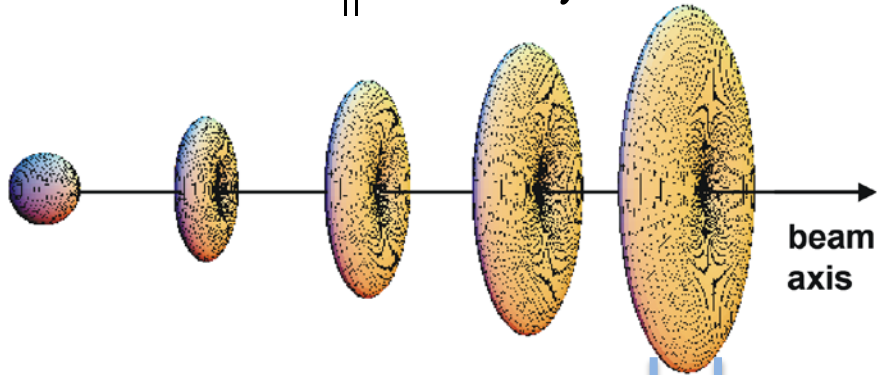


Explore relativistic quantum dynamics

- particle production
- non-perturbative regime
- coupling to the radiation field

Electromagnetic Phenomena under Extreme & Unusual Conditions

$$E_{\parallel} \propto 1/\gamma^2$$

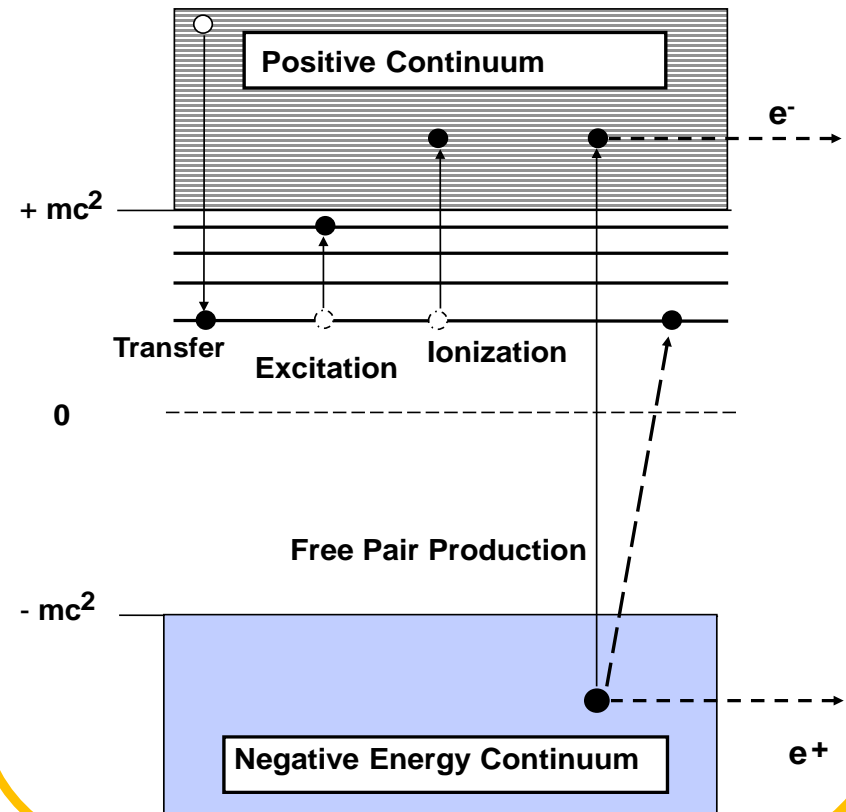


beam axis

$$E_{\perp} \propto \gamma$$

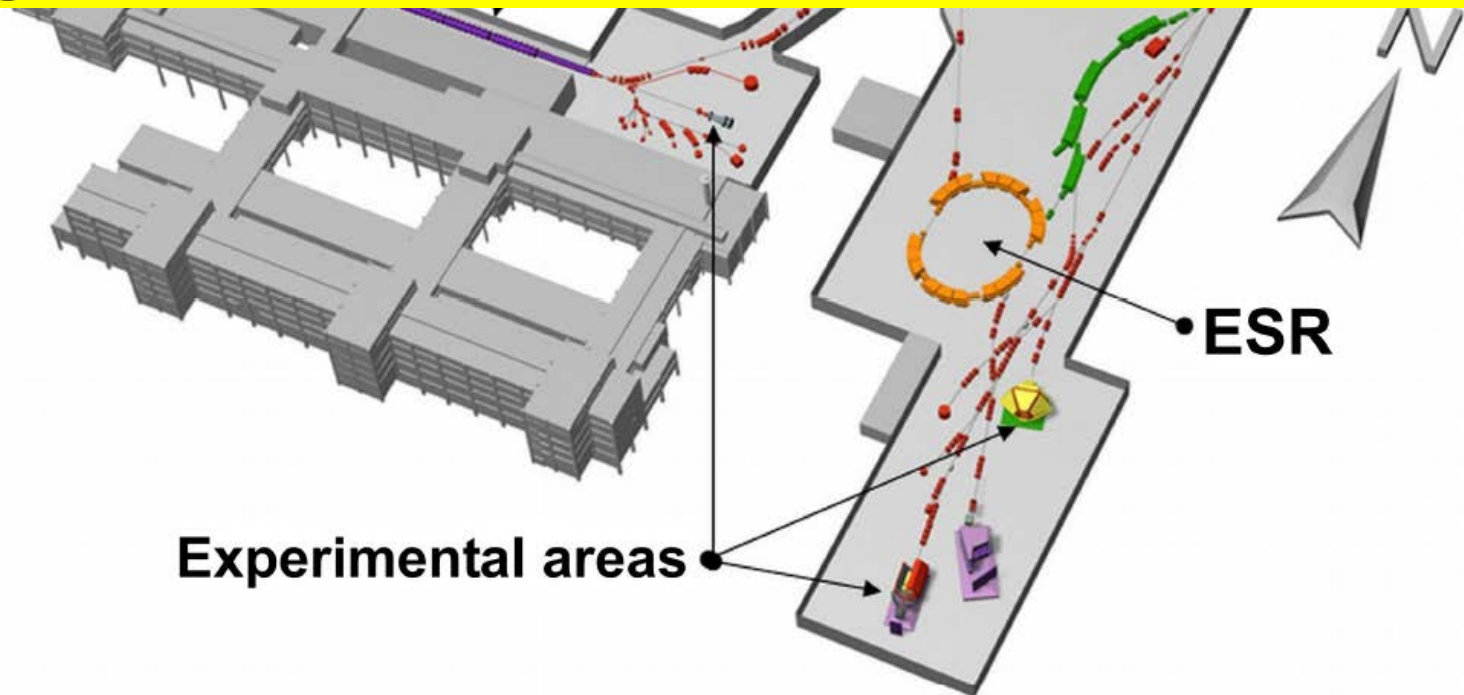


Collision times in the sub-attosecond regime
(10^{-22} s $<$ t $<$ 10^{-18} s)



Heavy Ion Accelerator System

- Every element in arbitrary charge state up to bare uranium are available for experiments
- Energies: from rest up to 1 GeV/u



- 10000 times increased intensity for rare isotopes
- energies from rest up to 35 GeV/u
- antiprotons



For atomic physics @ FAIR, x-ray and laser technology will be of outmost importance



September 2013



I. Extreme Dynamic Fields

II. Extreme Static Fields

III. Fundamental Physics

Our Tools

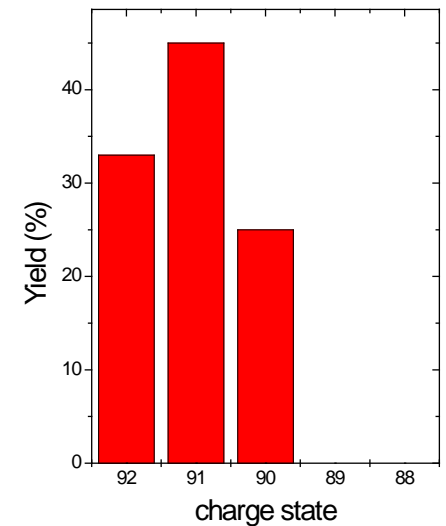
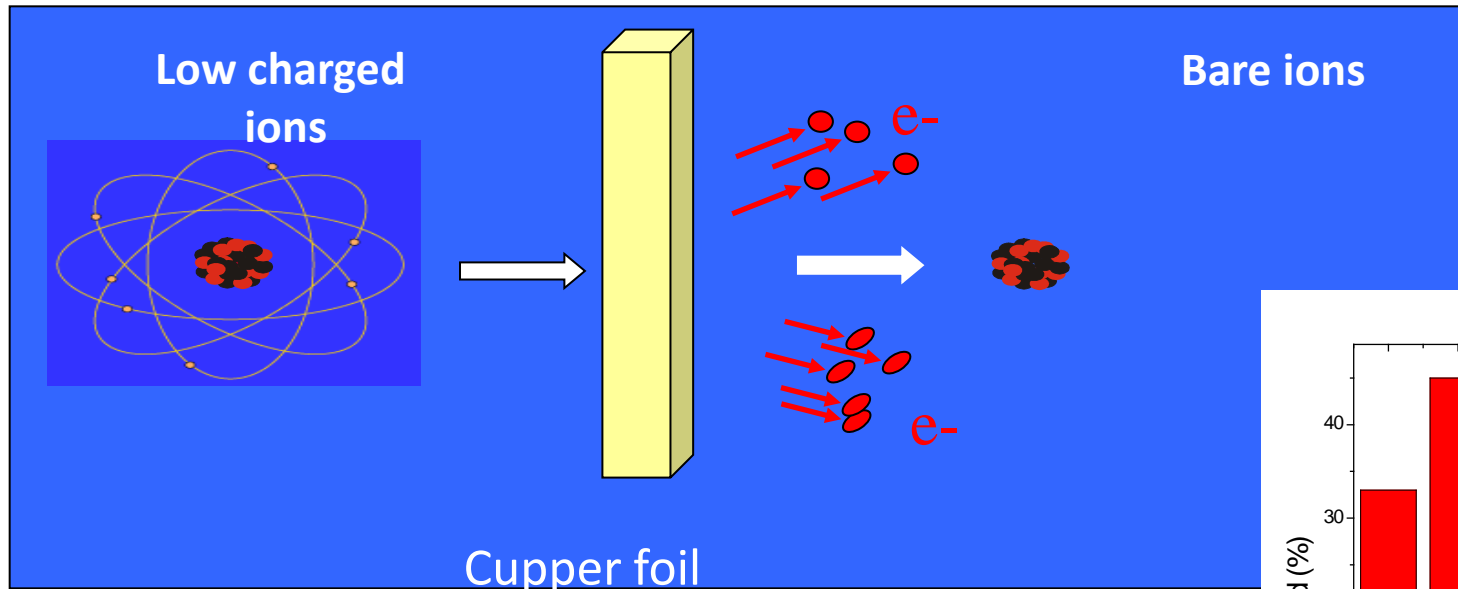
Why relativistic velocities?

Bohr criteria: *Largest ionization cross section at $v \approx v_K$*

Uranium

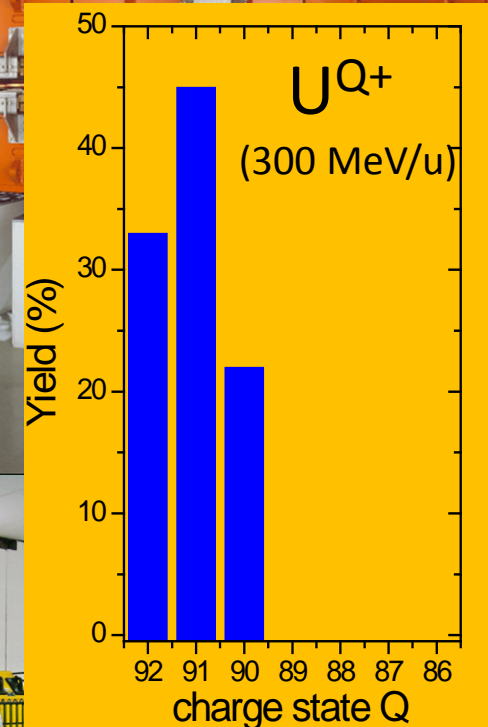
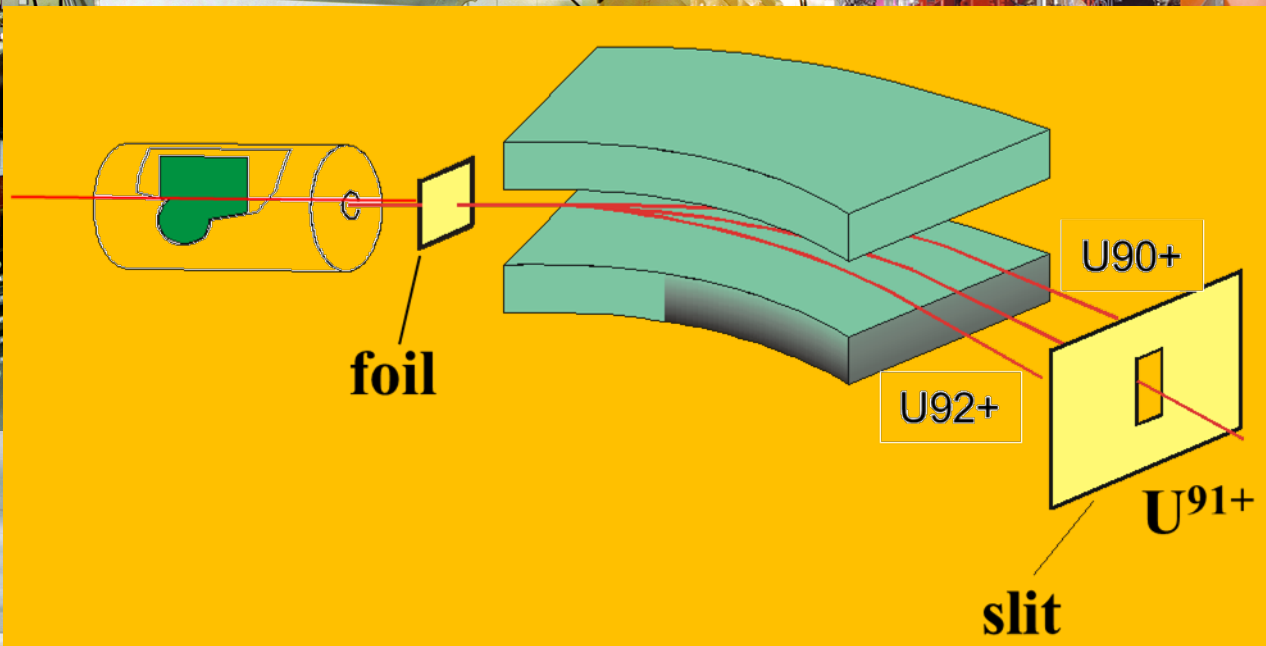
$$v_K/c \approx 0.67$$

Beam energy ???

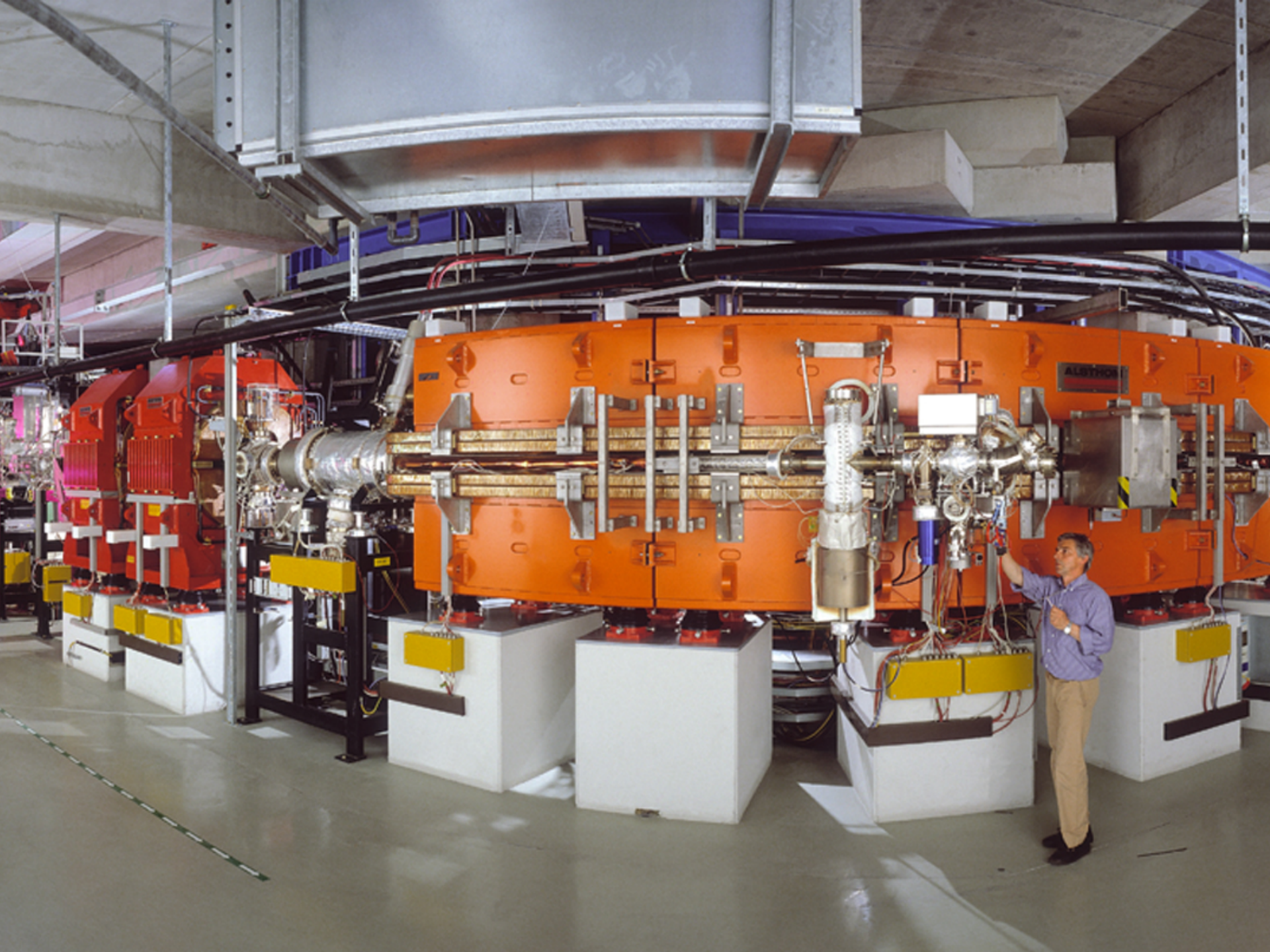


Storage Rings /Synchrotrons/ Charge State Separators

ESR/GSI



Every element in arbitrary charge state up to the heaviest bare elements are available for experiments



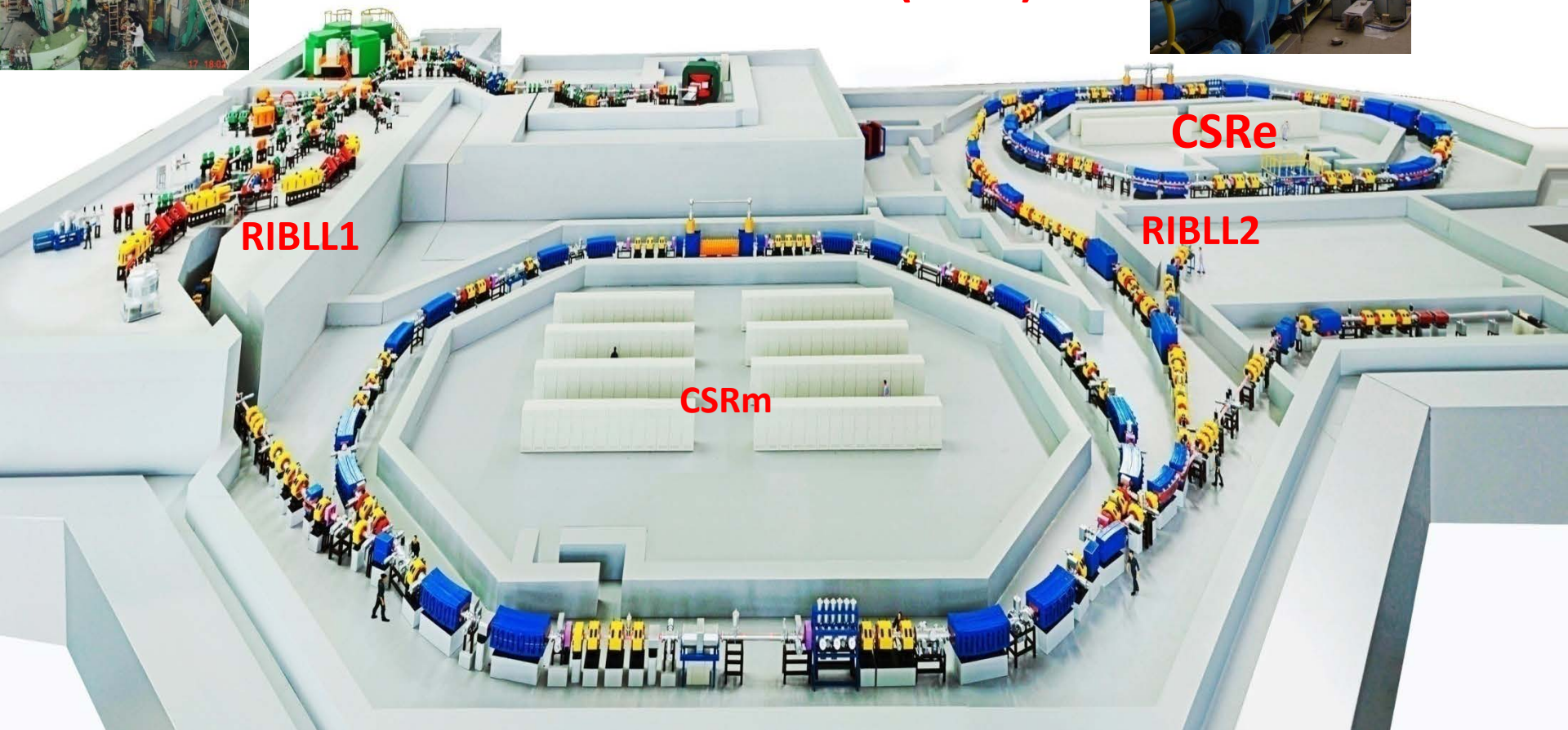
ALBTRON

Heavy Ion Research Facility in Lanzhou (HIRFL)

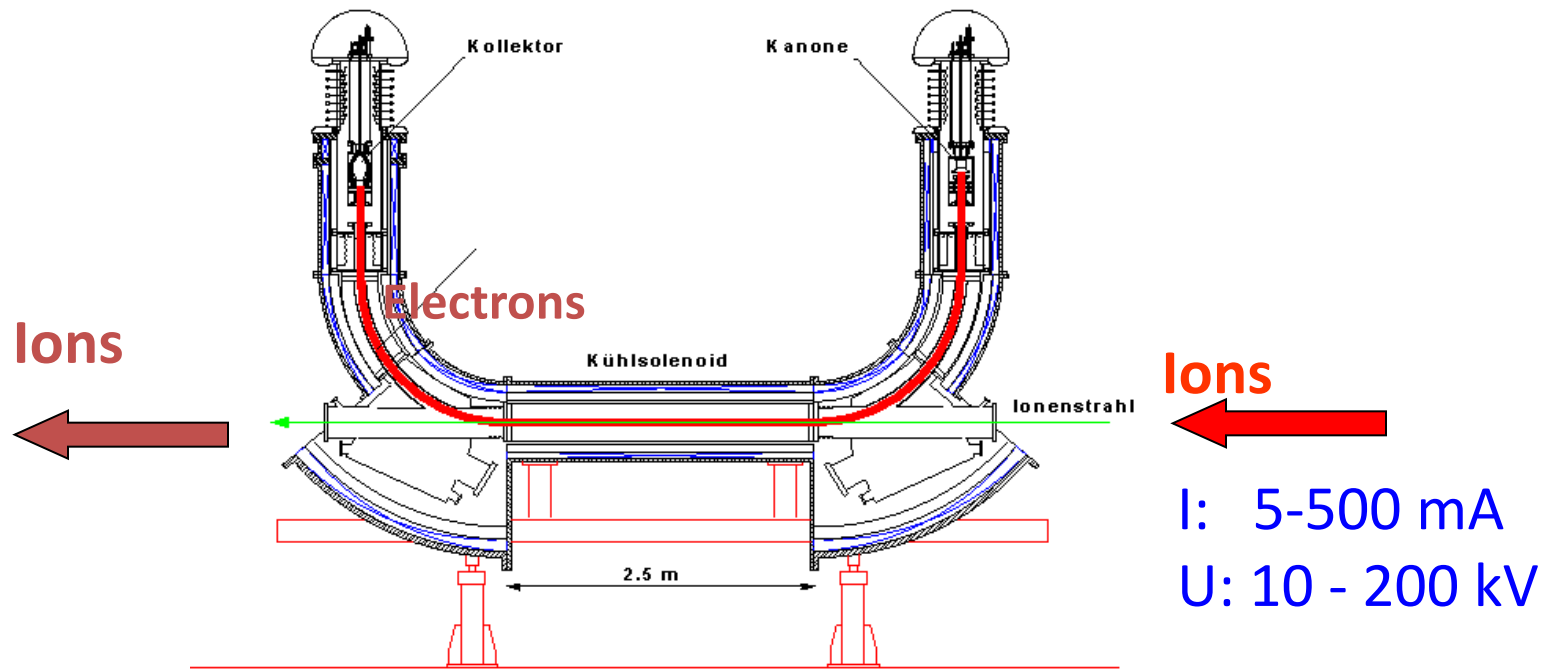


SSC(K=450)

SFC (K=69)



Electron Cooling



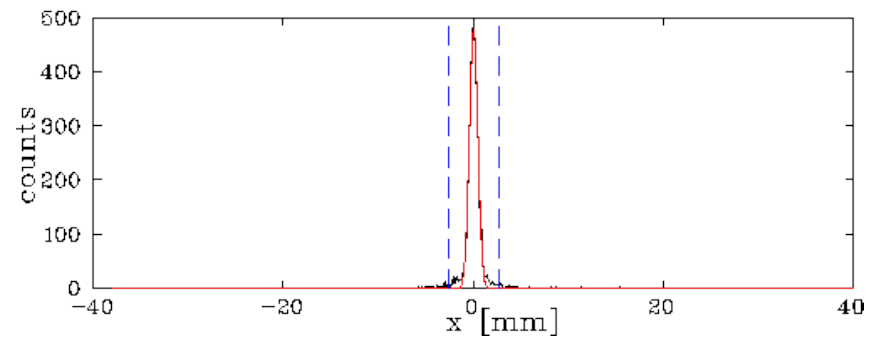
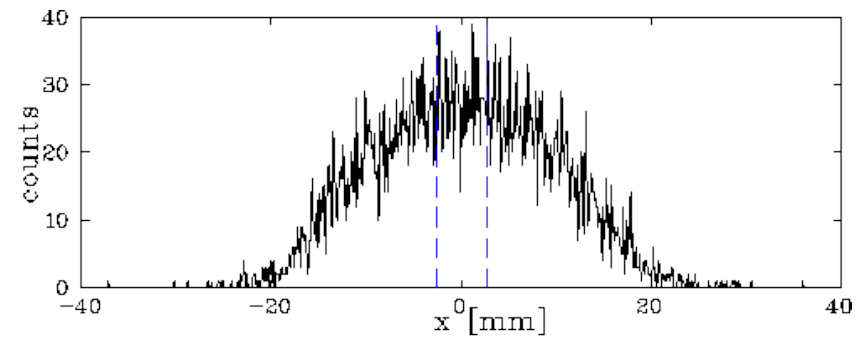
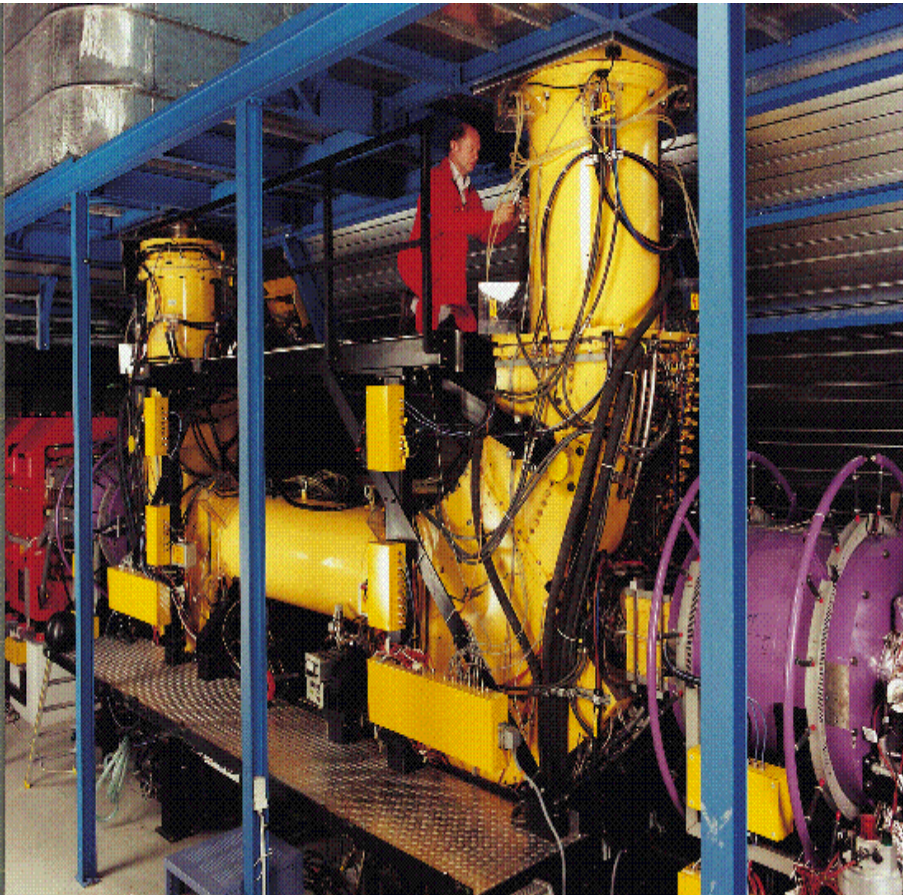
Ions interact 10^6 1/s with a collinear beam of cold electrons

Properties of the cold ions

Momentum spread $\Delta p/p : 10^{-4} - 10^{-5}$
Diameter 2 mm

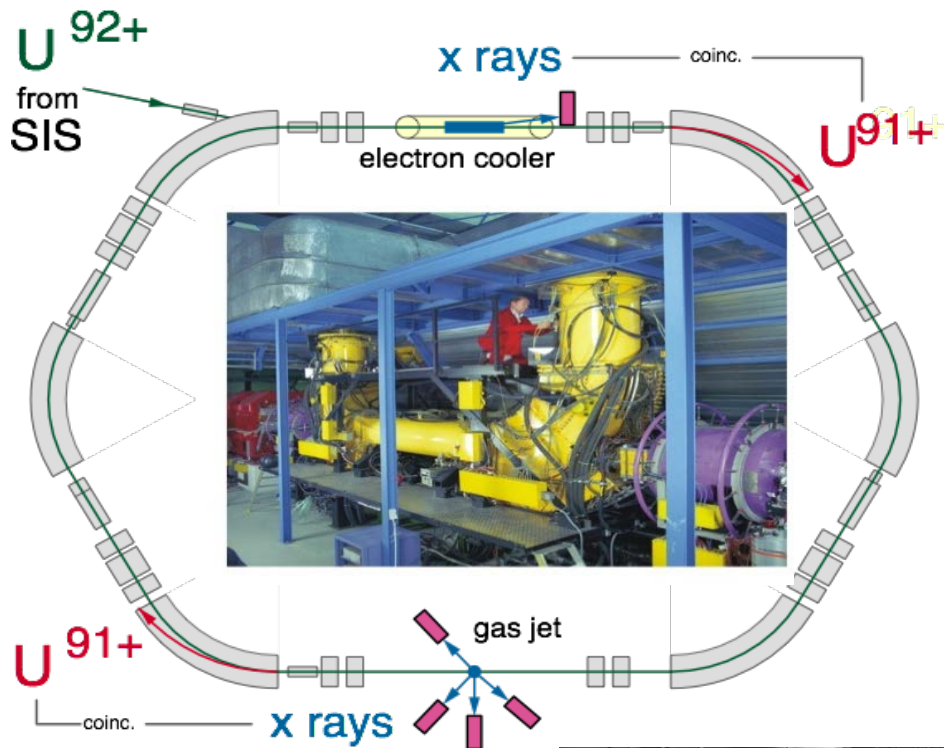
'Cooling': narrowing velocity, size and divergence of the stored ions

Electron cooling: Budker, 1967 Novosibirsk

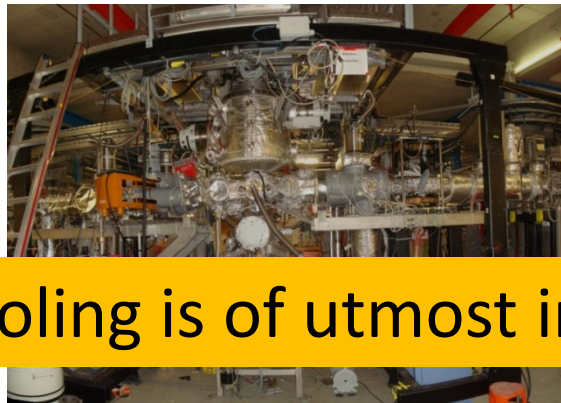
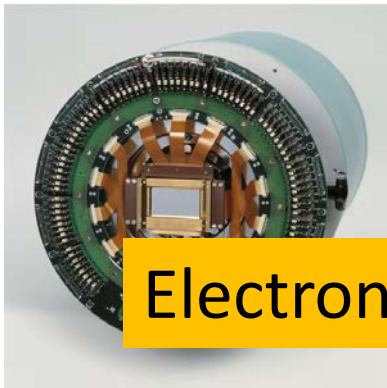
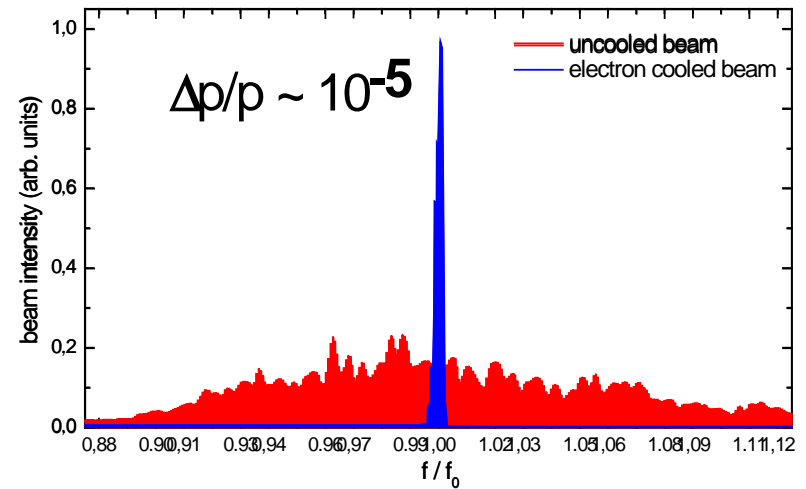


momentum exchange
with 'cold', collinear e- beam. The ions
get the **sharp velocity** of the electrons,
small size and divergence

Storage Rings



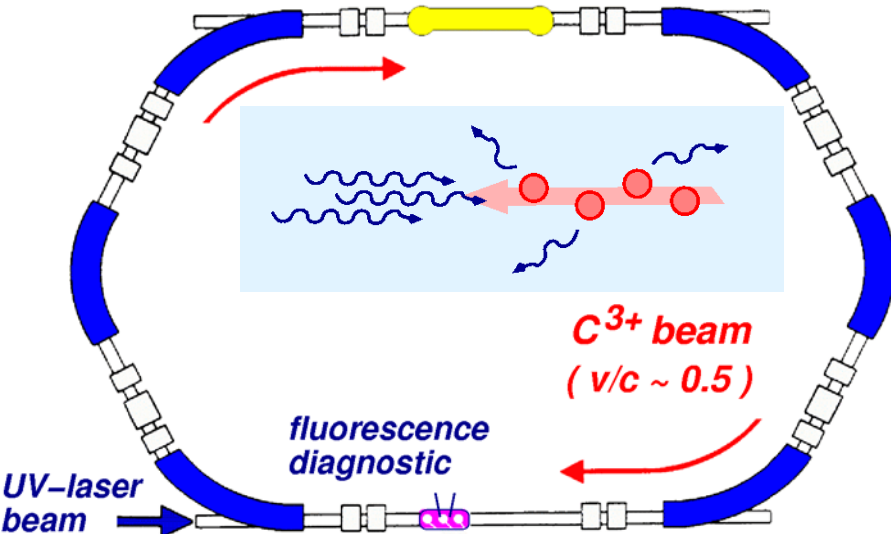
electron cooling



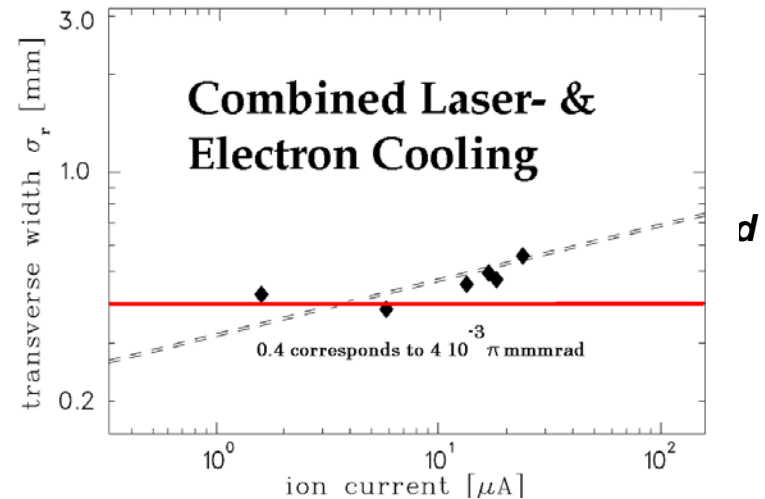
Electron Cooling is of utmost importance

Laser cooling of C^{3+} beams

momentum dependent (Doppler tuned)
 laser deceleration + **bunching**
 (**restoring force**) \Rightarrow cooling

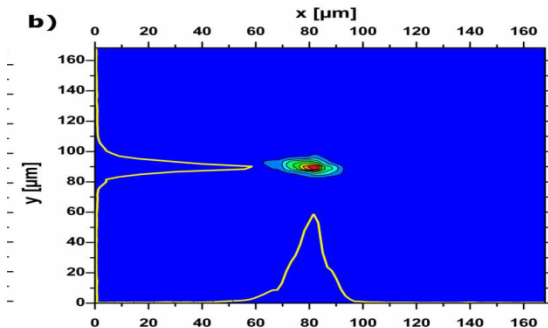


Demonstration of laser cooling of C^{3+} ions at 122 MeV/u in the ESR for application at SIS 100/300



bunch length reduced by a factor 2
 beam diameter reduced by a factor 4
 momentum spread reduced by a factor 10

U. Schramm et al.,



Experimental Conditions at the HESR

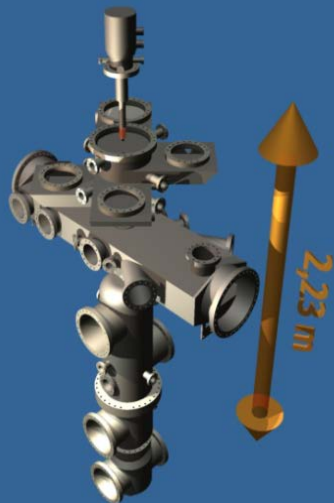
location of target stations



**Worldwide premiere:
Precision experiments using
cooled relativistic ion beams**

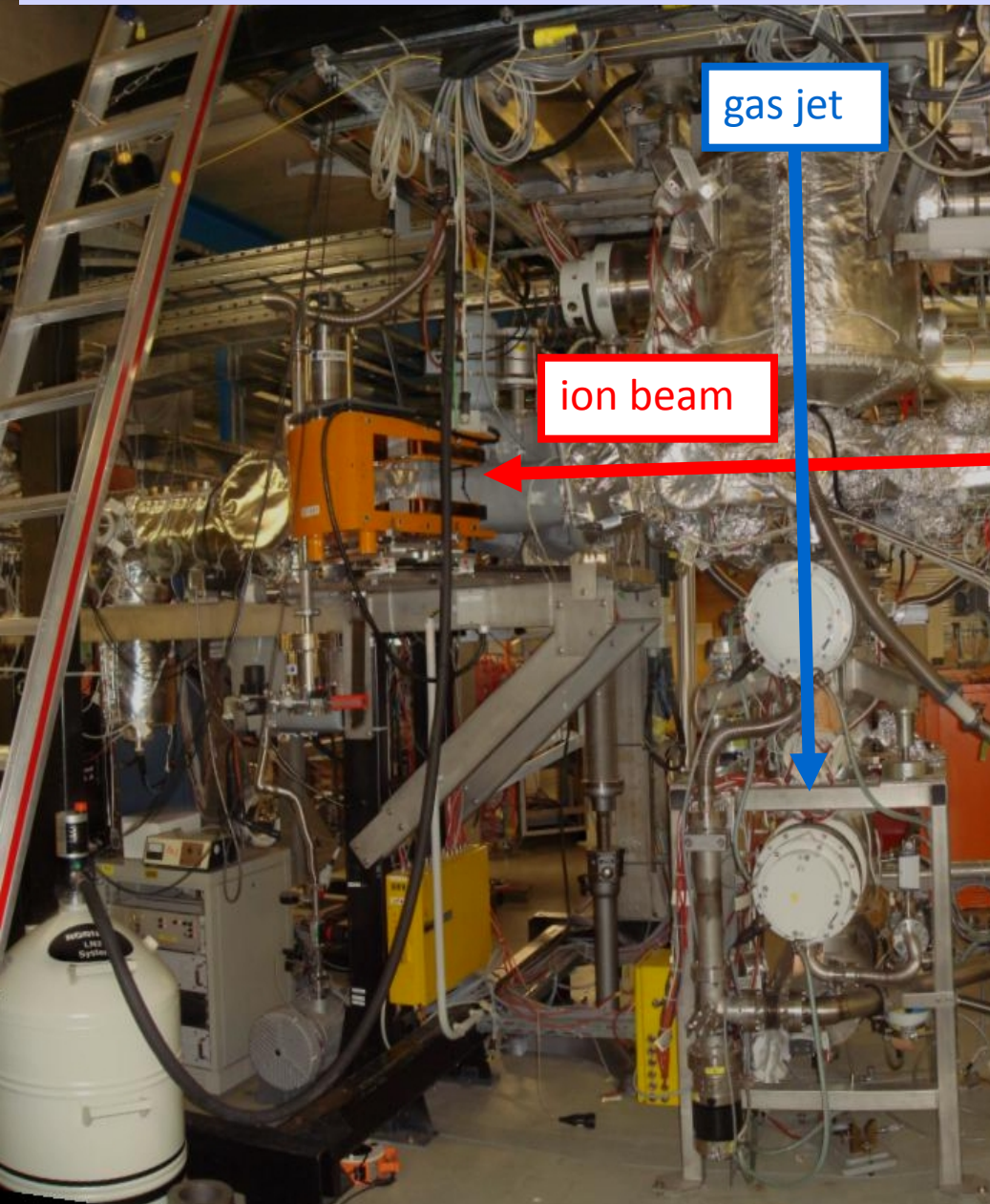
- species: p, pbar, HCl, RIB
- circumference 574 m
- injection energy 740 MeV
- $B\rho = 50 \text{ Tm}$
- for U^{92+} : 4.937 GeV/u
- $\gamma_{\text{MAX}}=6.30$; $\beta_{\text{MAX}}=0.987$
- momentum (energy) range

- Stochastic cooling & electron cooling
- electron-, gasjet-, fiber-targets (!)
- Particle detectors
- Ion stacking
- Luminosity (number of stored ions)
- Beam diameter/charge separation
- Acceleration and deceleration
- Coupling of laser to the ion beam line
- Building / Space for setups



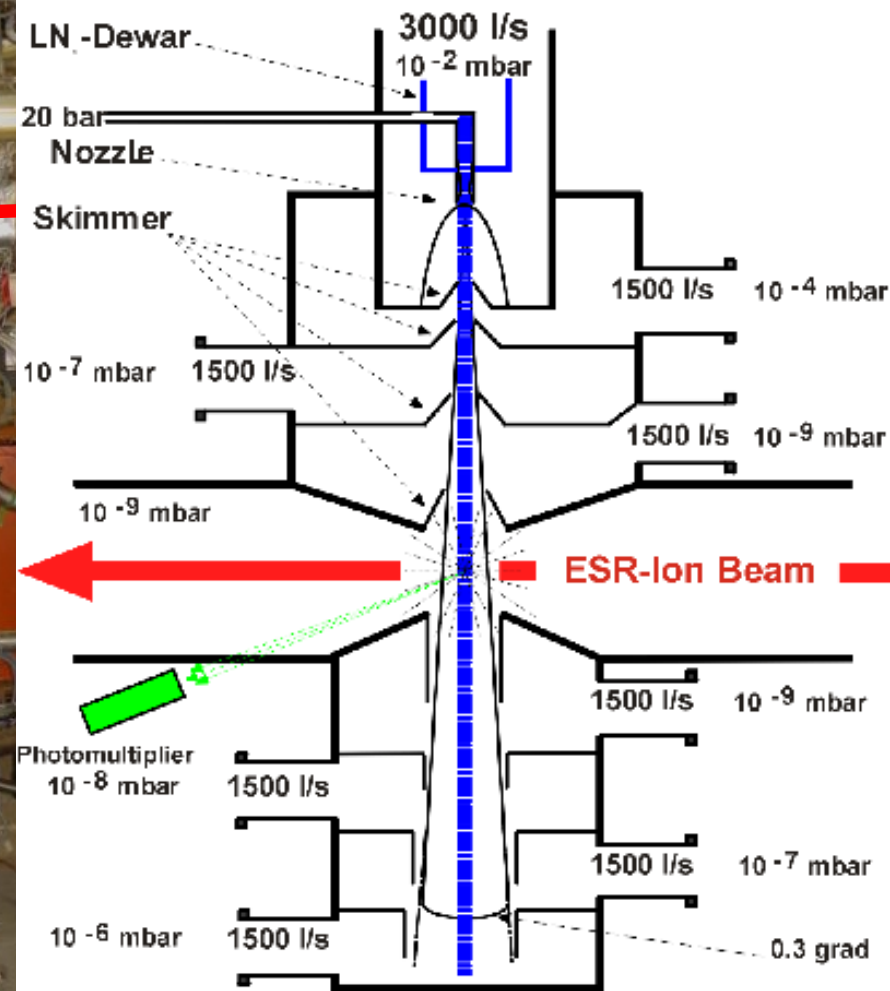
the
SR

Targets and spectrometer ...



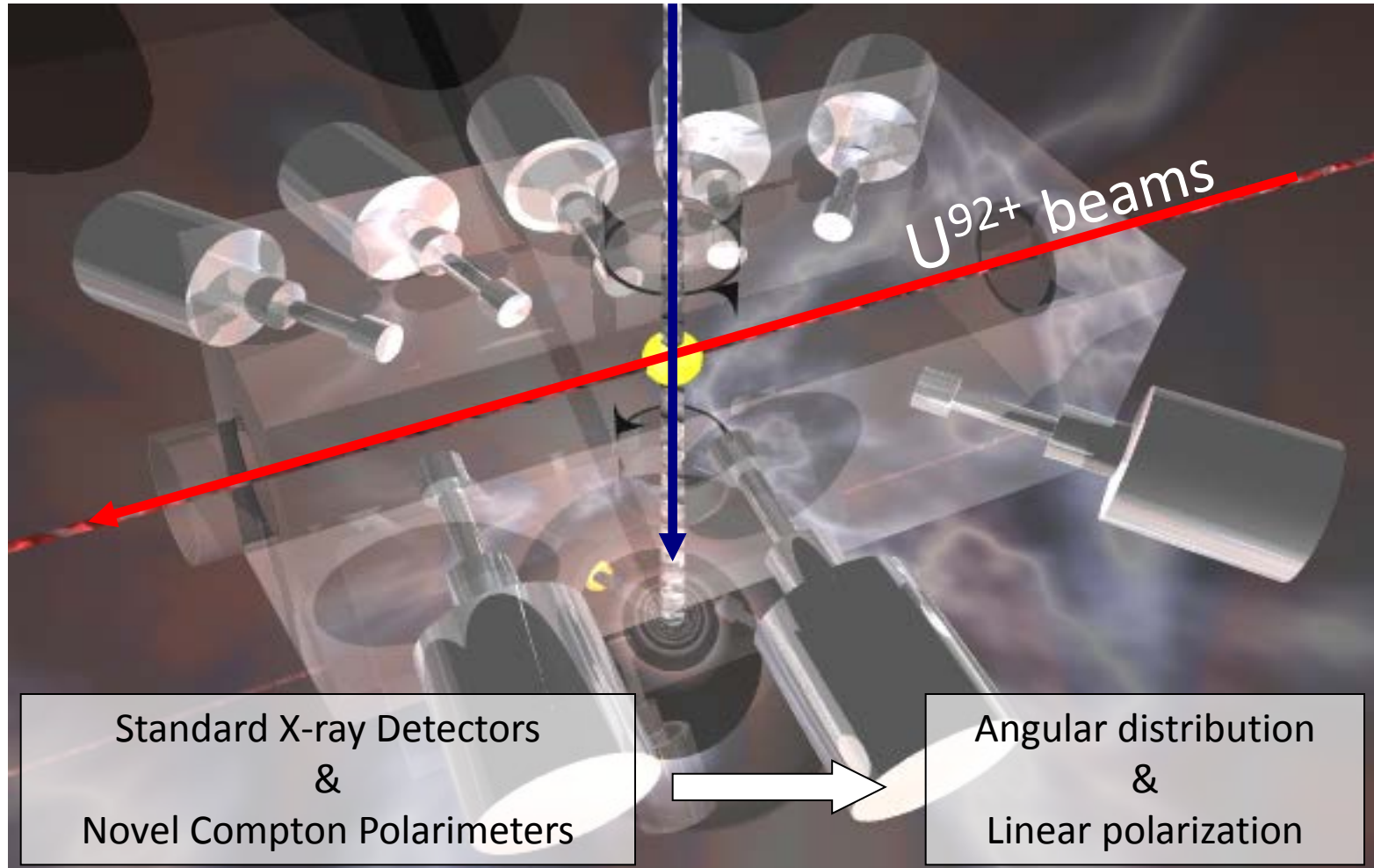
gas jet

ion beam



Setup at an gas target for x-ray detection (spectroscopy, angular distributions, polarimetry)

H₂, He, N₂, Ne, ... Xe gas target

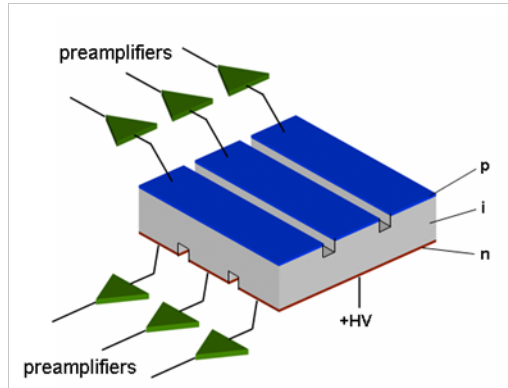
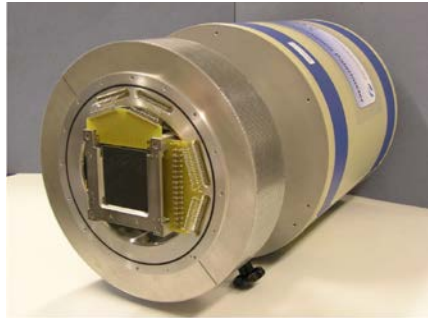


Standard X-ray Detectors
&
Novel Compton Polarimeters

Angular distribution
&
Linear polarization

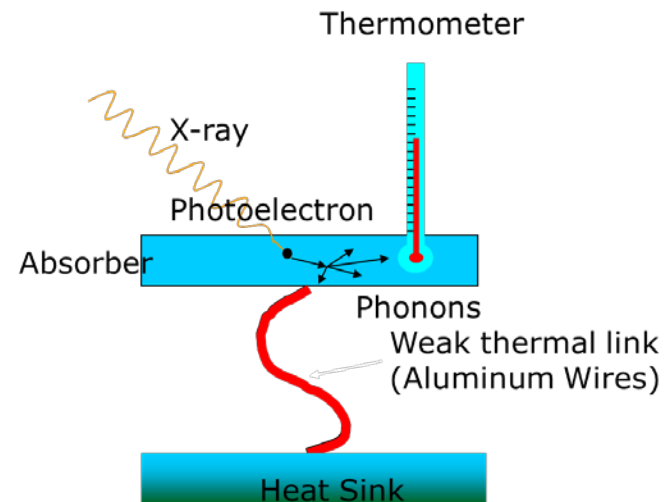
Novel detectors

Micro-strip semiconductor detectors

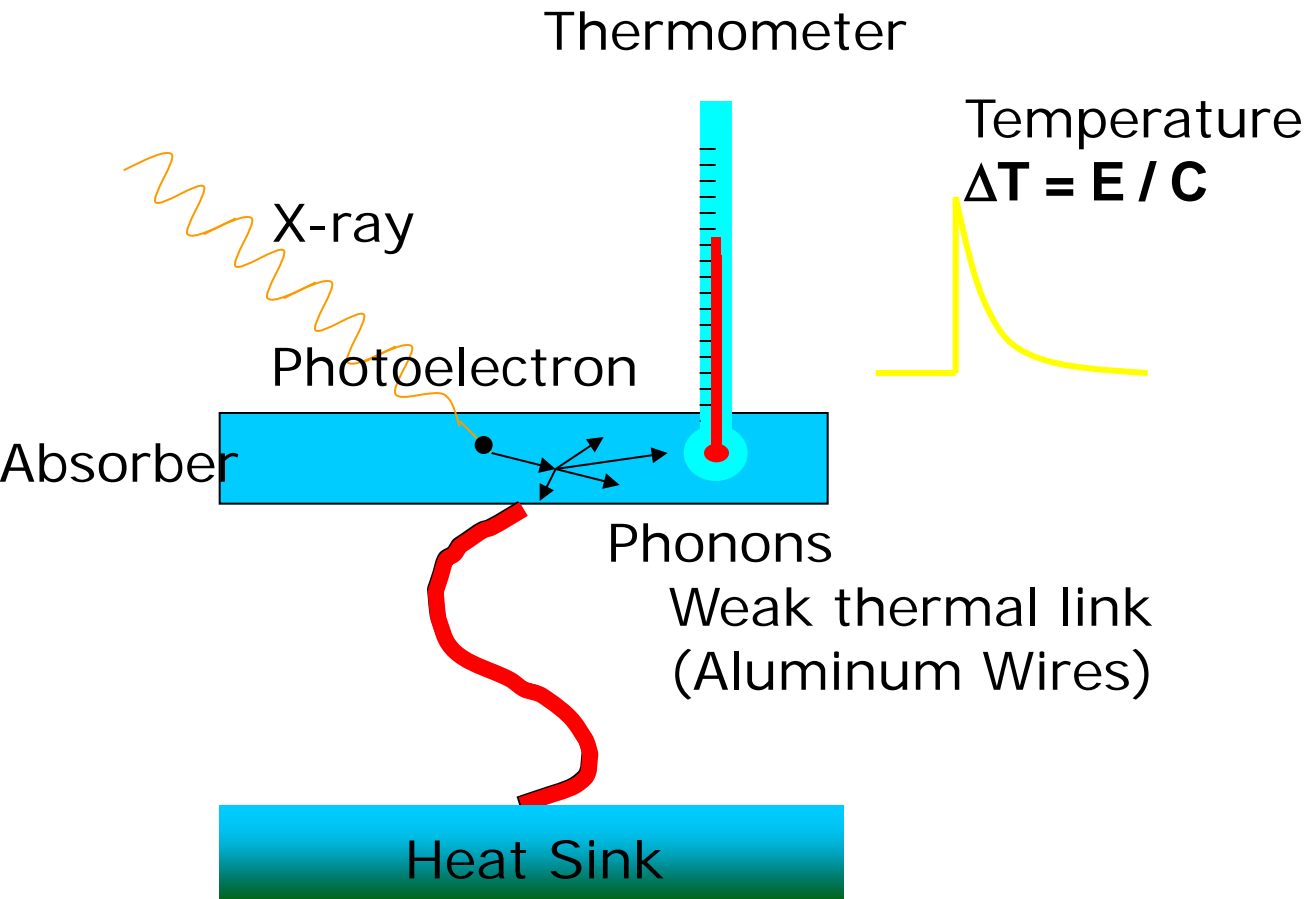


- ☑ Si(Li) or Ge(i)
- ☑ energy resolution
- ☑ timing
- ☑ 2D (3D) position sensitivity
- ☑ multi-hit capability

Micro-Calorimeter



Micro-Calorimeter



Heat capacity: $C = c \cdot m$
 $C \sim T^3$

Specific
heat capacity : c

Detector
mass: m

*Detector
operates
at about
50 mK*

Micro-calorimeter detector: large wavelength acceptance, large quantum efficiency, and excellent energy resolution (4 keV@5eV => 35 keV@30 eV).

maXs-200: detector arrays for hard x-rays

First array prototype

1x8 x-ray absorbers

2 mm x 0.5 mm

200 μm thick electroplated Au

80% QE at 100 keV

$\Delta E_{\text{FWHM}} < 50 \text{ eV}$

4 electronic channels

modular extension to longer arrays

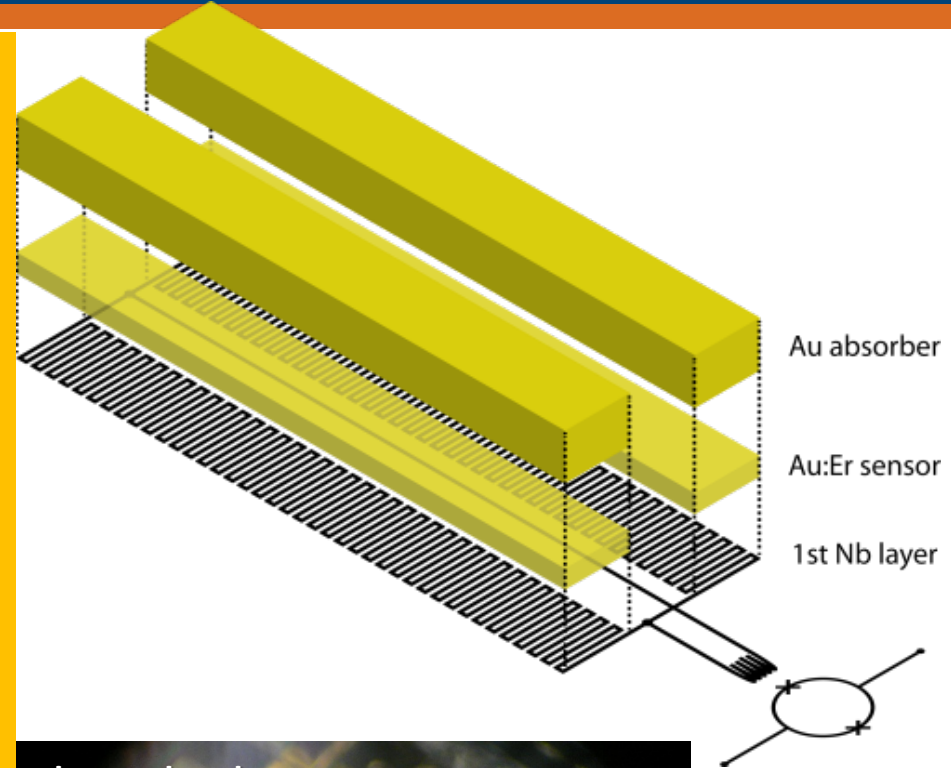
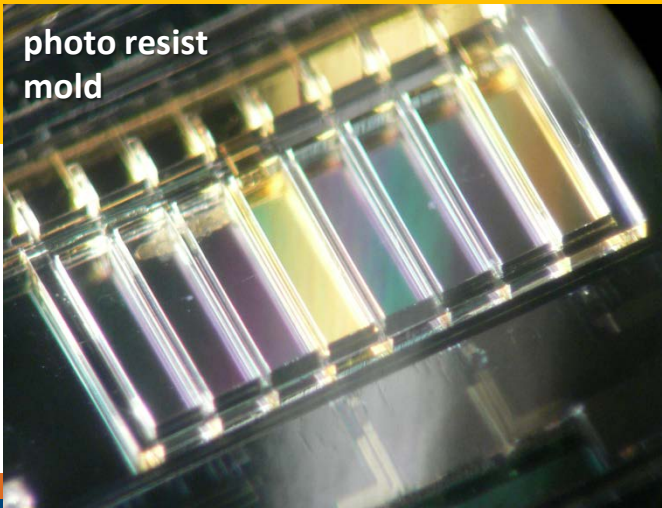
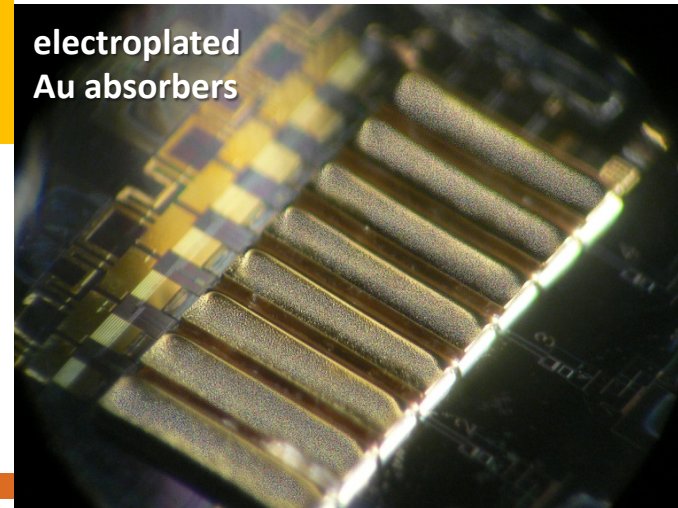


photo resist
mold

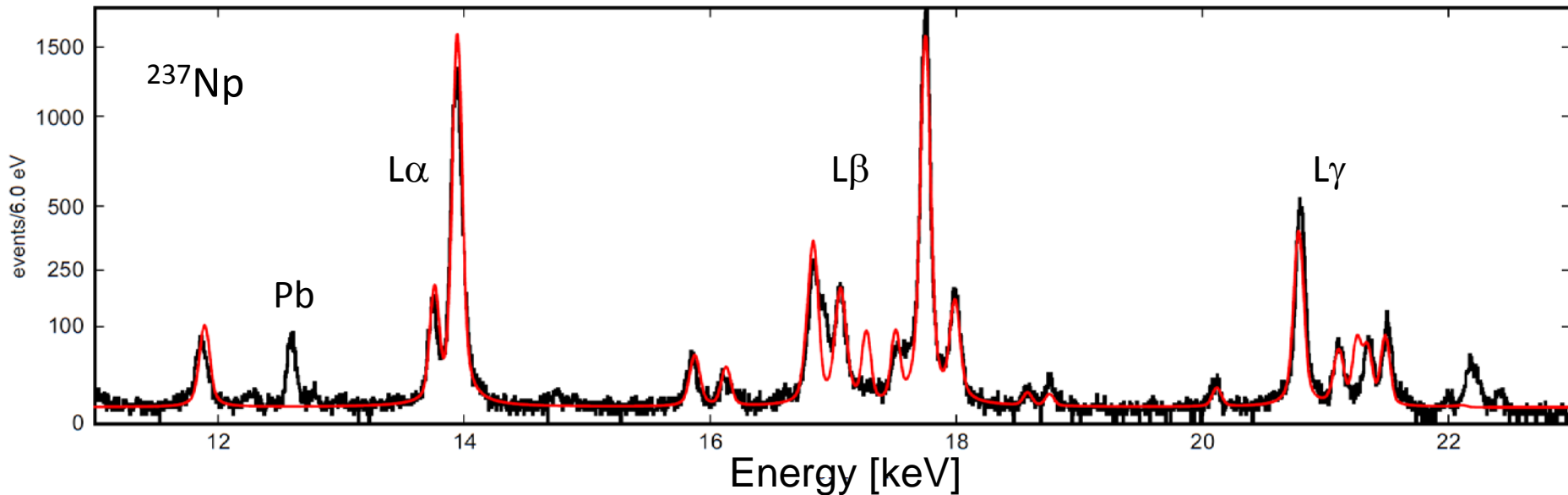


electroplated
Au absorbers



maXs-200: detector arrays for hard x-rays

First characterization with an ^{241}Am -source



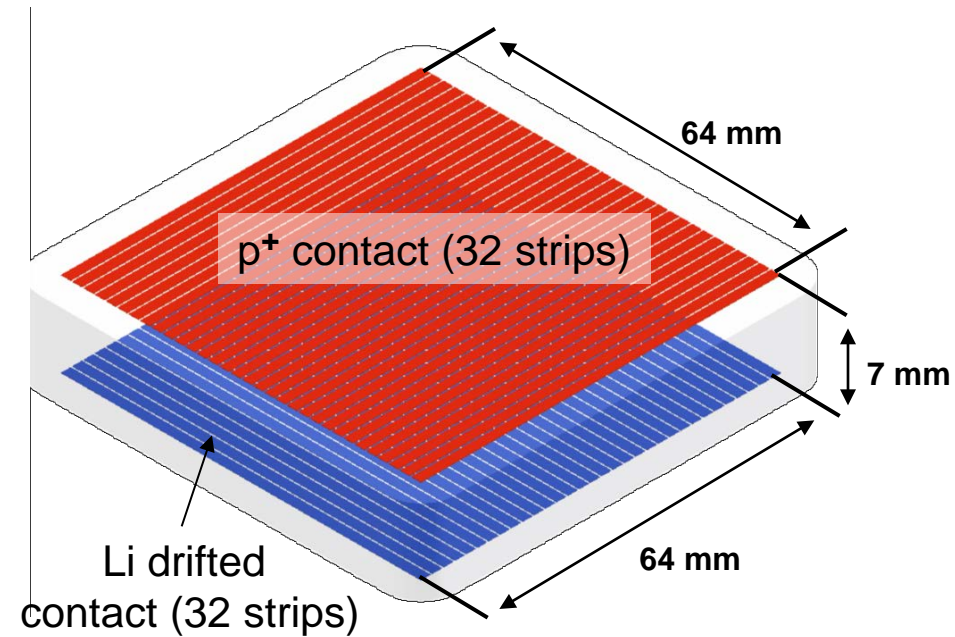
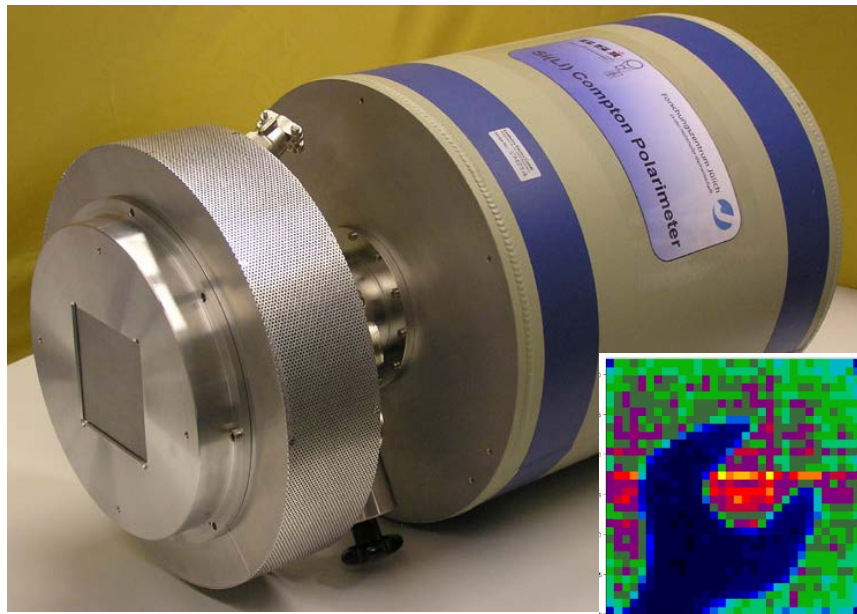
$$\Delta E_{\text{FWHM}} = 40 \text{ eV @ } 0\text{-}10 \text{ keV}$$

$$\Delta E_{\text{FWHM}} = 60 \text{ eV @ } 60 \text{ keV}$$

Slight degradation towards higher energies due to

- **Poor temperature stability** in this first experiment
- **Possible marginal position dependence**,
to be fixed by stems between absorber and sensor

2D Si(Li) X-ray Detector as a Compton Polarimeter



32x32 strips → 1024 pixels

64x64 mm → 4096 mm² active area

readout rate: a few kHz

Energy (1.5 keV FWHM) + Timing (100 ns) + 2D Position (2 mm) + Multihit Capability

Dedicated to efficient and precise polarization studies from 70 keV to a few 100 keV

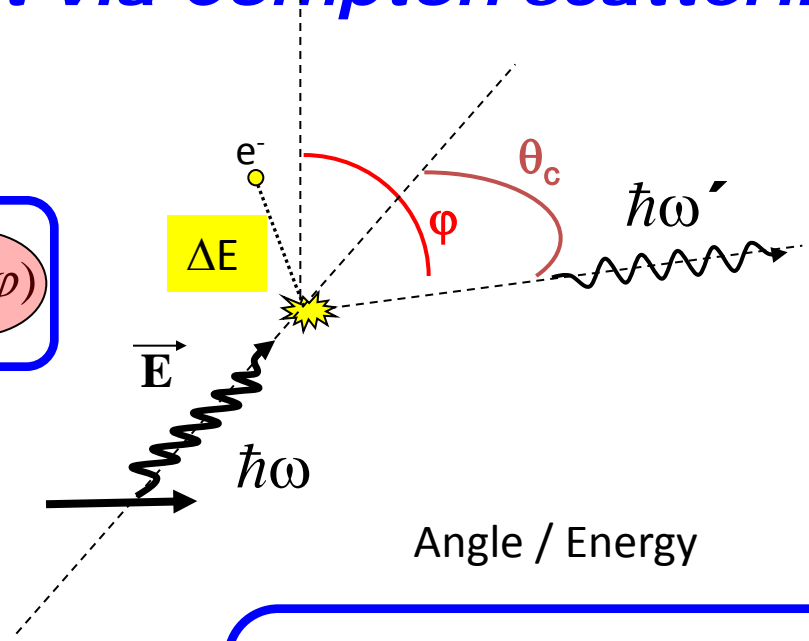
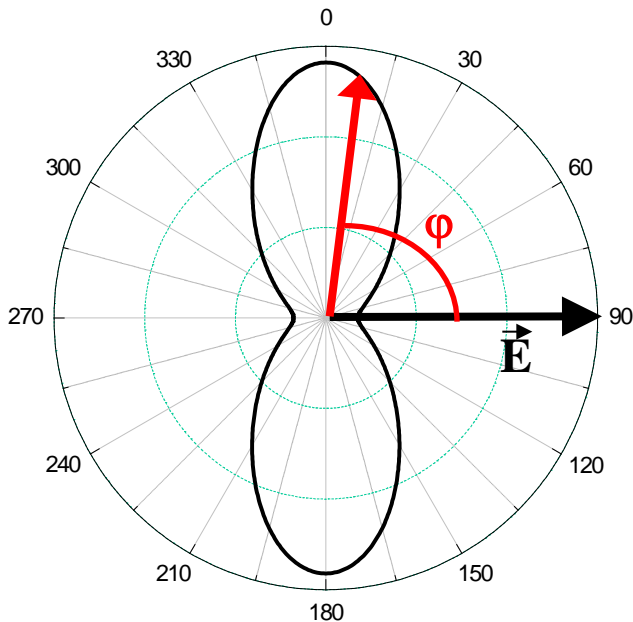
How to Measure Polarization for Hard X-Rays

Polarization Measurement via Compton scattering

Linearly polarized radiation

Klein-Nishina equation

$$\frac{d\sigma}{d\Omega} = \frac{1}{2} r_0^2 \left(\frac{\hbar\omega'}{\hbar\omega}\right)^2 \left(\frac{\hbar\omega'}{\hbar\omega} + \frac{\hbar\omega}{\hbar\omega'} - 2 \sin^2 \theta_c \cos^2 \varphi\right)$$



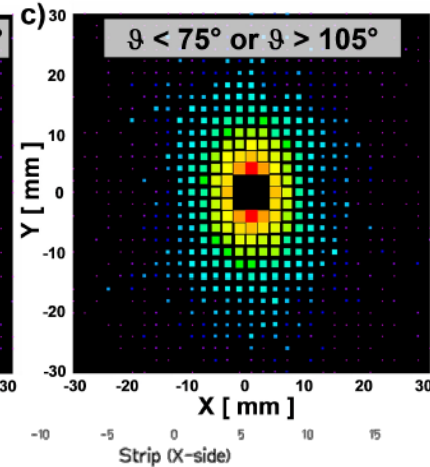
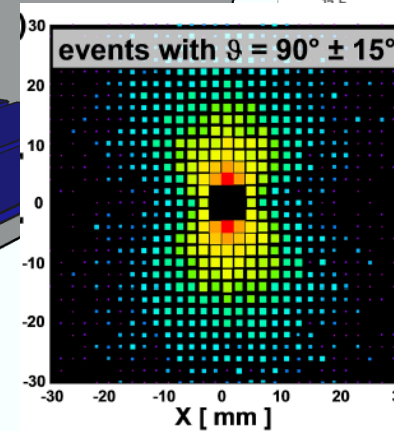
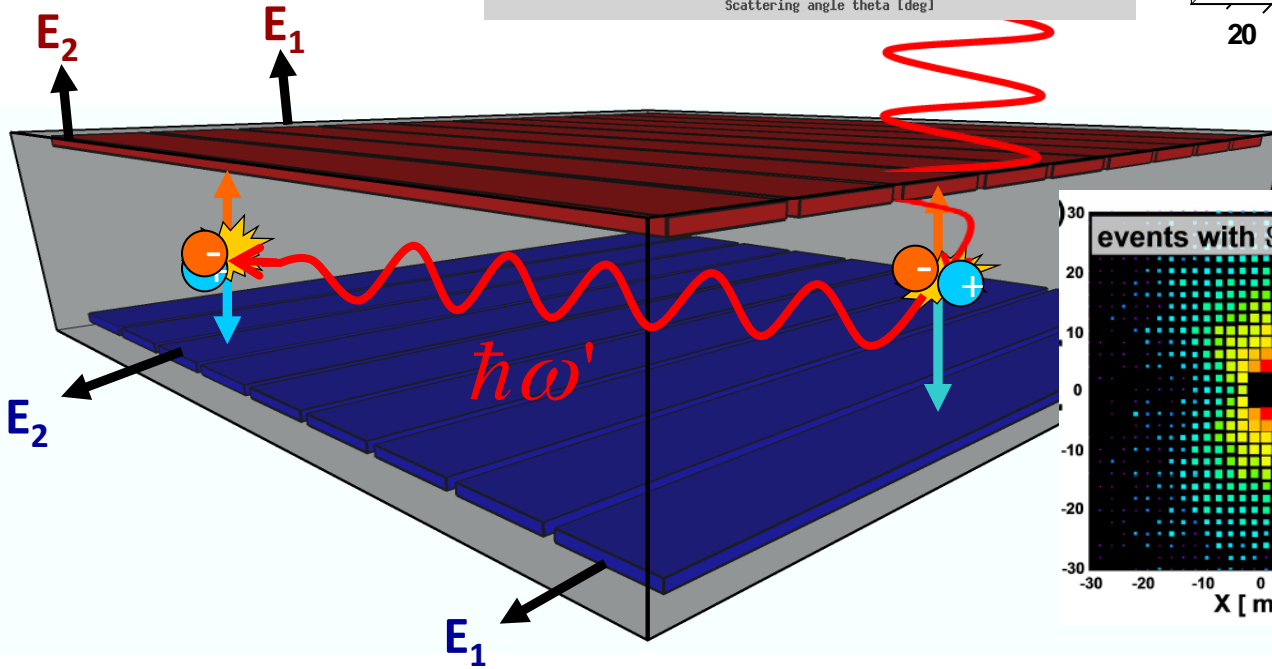
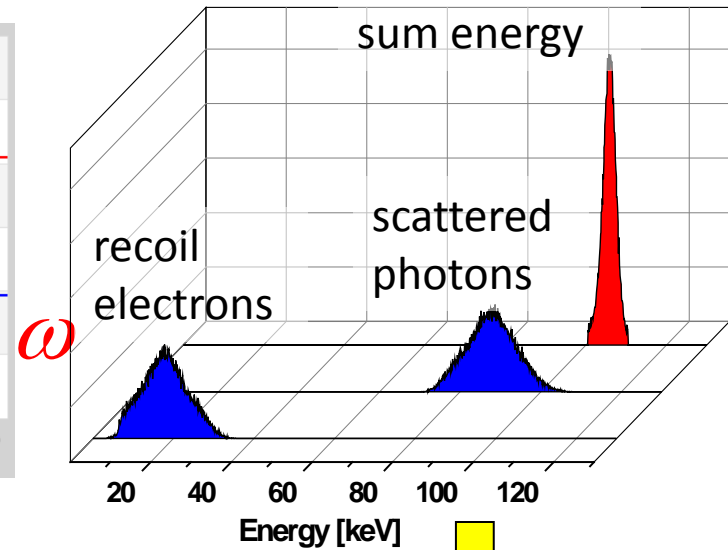
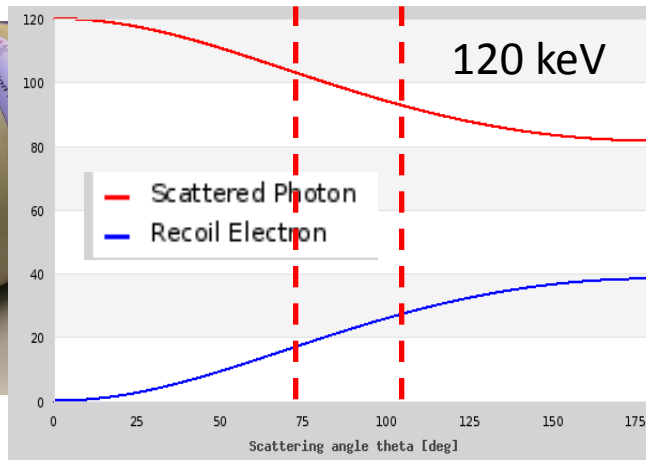
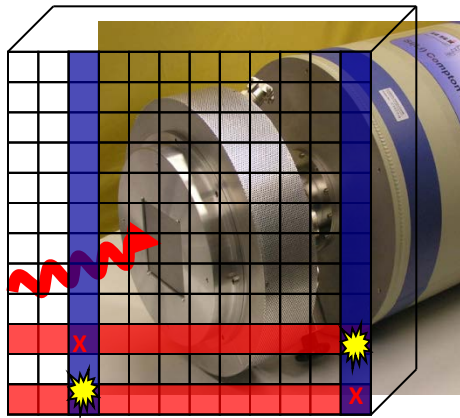
Angle / Energy

$$\hbar\omega' = \frac{\hbar\omega}{1 + \frac{\hbar\omega}{m_e c^2} (1 - \cos \theta_c)}$$

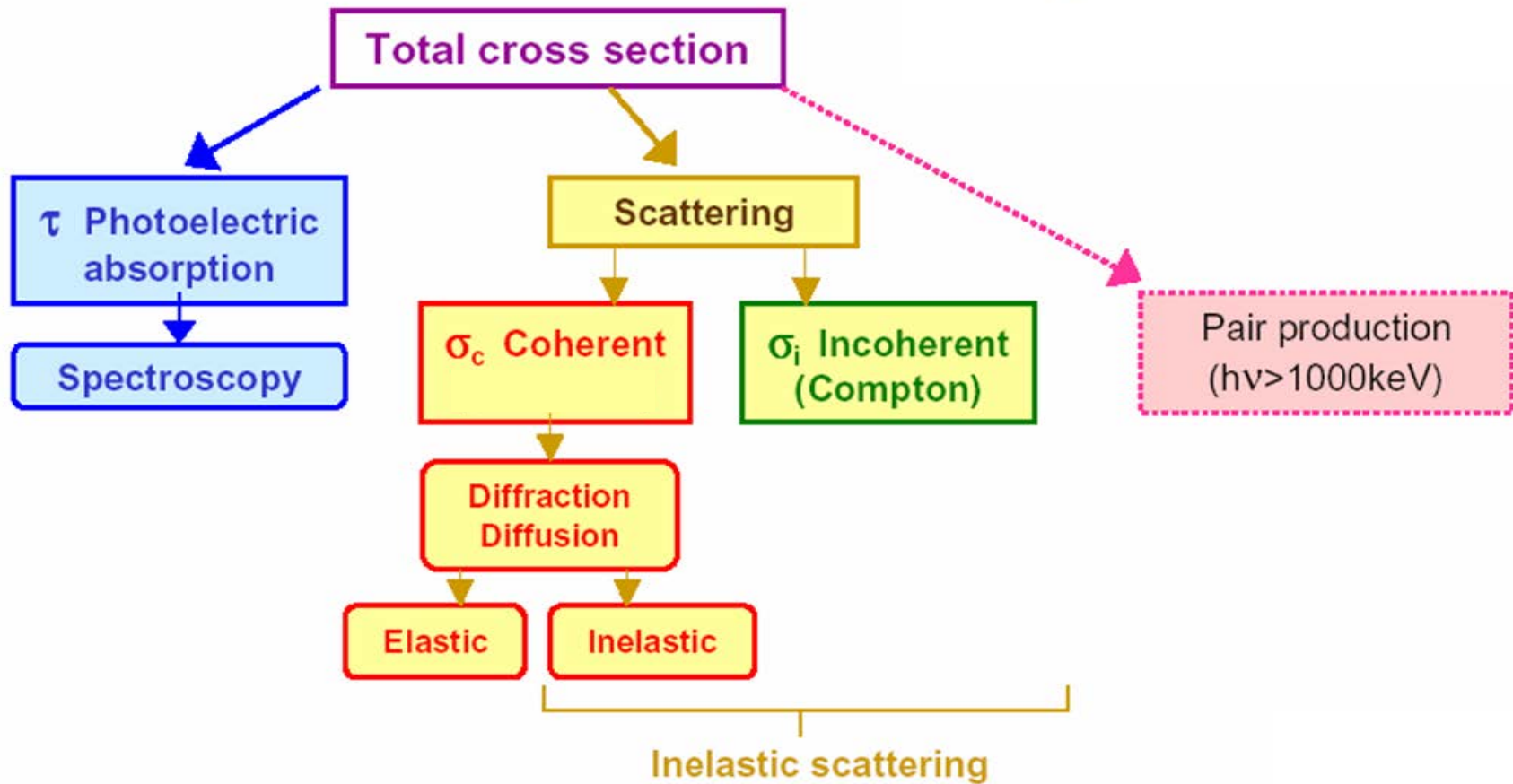
$$\hbar\omega = \hbar\omega' + \Delta E$$

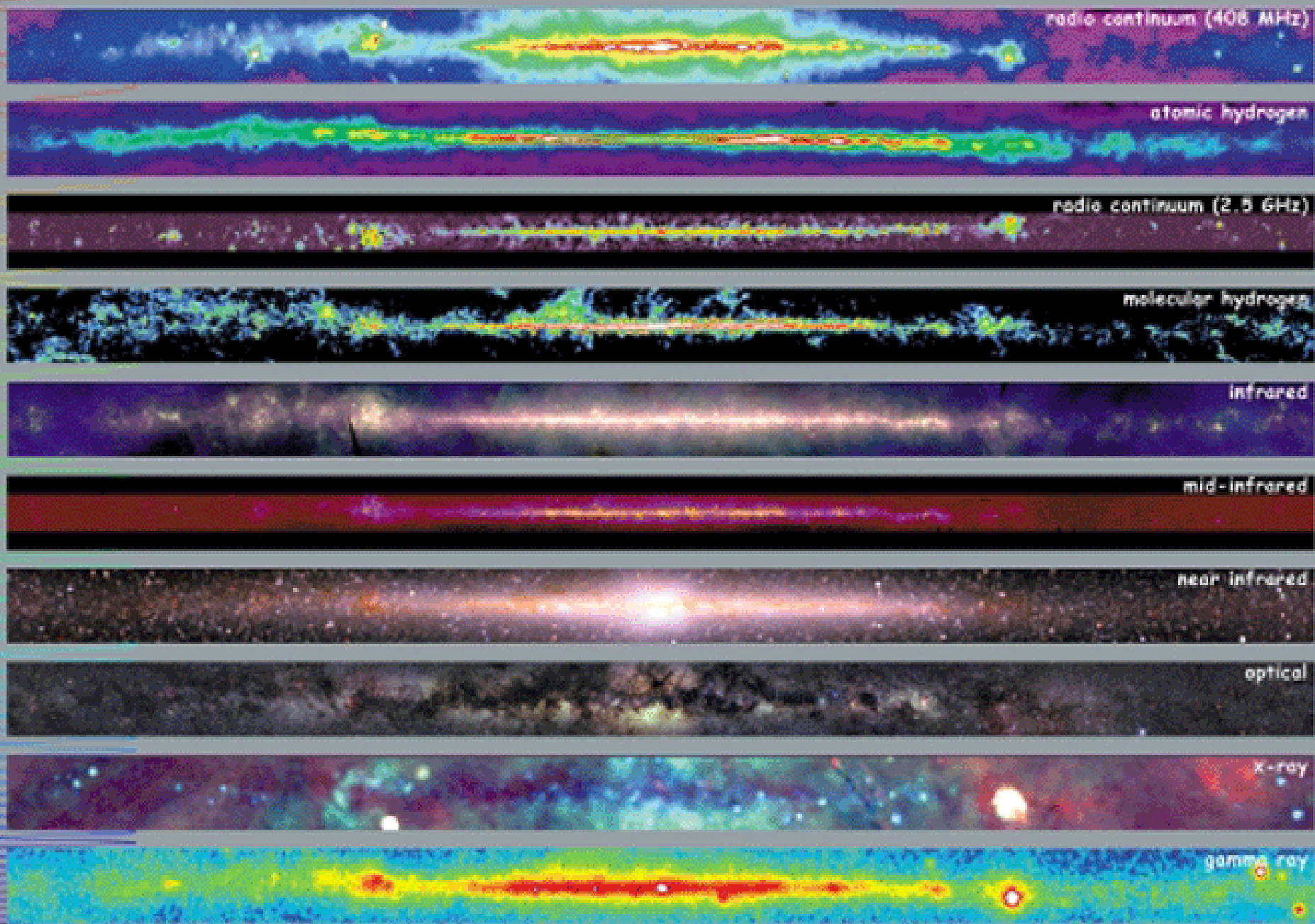
ΔE : electron recoil energy

Reconstruction of the Compton events



Interaction of radiation and matter

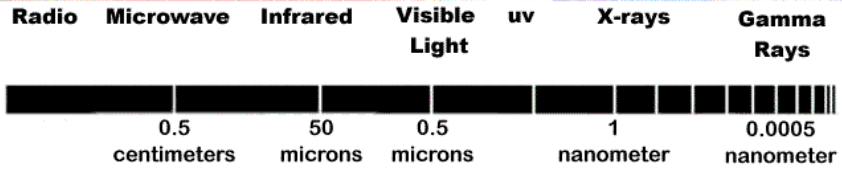
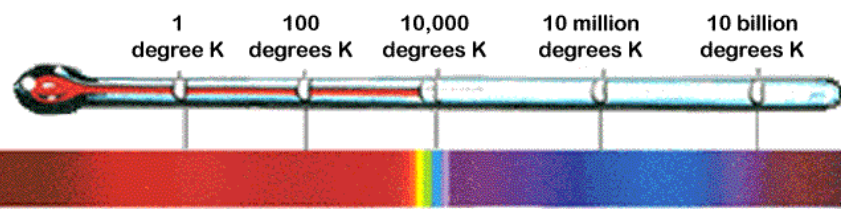




<http://adc.gsfc.nasa.gov/mw>



Multiwavelength Milky Way



WAVELENGTH

Note: degrees Kelvin (K) = degrees Celsius (C) + 273

Radio Infrared Visible UV X-Ray Gamma ray



	Radio	Infrared	Visible	UV	X-Ray	Gamma ray	E (eV)
λ (m)	1	1×10^{-3}	1×10^{-6}	1×10^{-9}	1×10^{-12}	1×10^{-15}	1×10^{-18}
ν (Hz)	2×10^8	2×10^{11}	2×10^{14}	2×10^{17}	2×10^{20}	2×10^{23}	2×10^{26}

$E: h c / \lambda$

$\lambda: c / \nu$

$\nu: c / \lambda$

E: Energy

λ : Wavelength

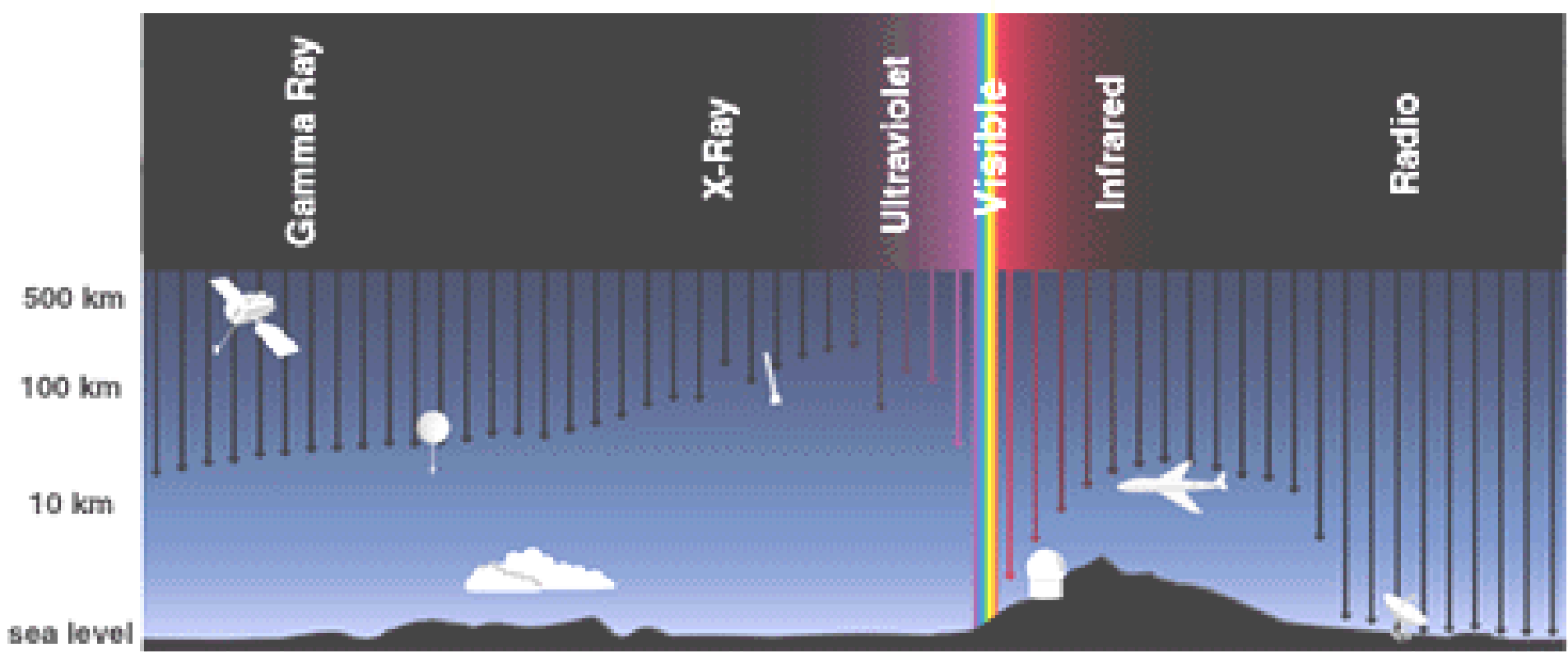
ν : Frequency

h: Planck constant

$= 4 \times 10^{-15} \text{ eV s}$

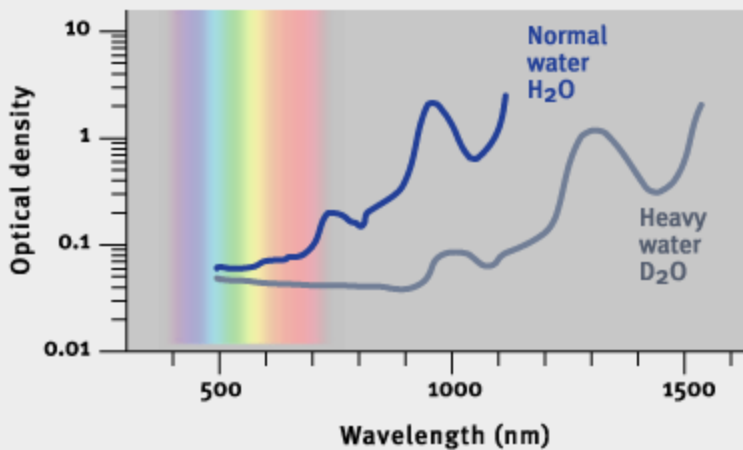
c: speed of light

$= 3 \times 10^8 \text{ m/s}$



<http://imagers.gsfc.nasa.gov/ems/atmosphere.gif>

Comparison of absorption of H₂O and D₂O



**atmospheric
interactions**

reflection
refraction
scattering
absorption

Lecture 1

Dirac Theory

DER SPIEGEL

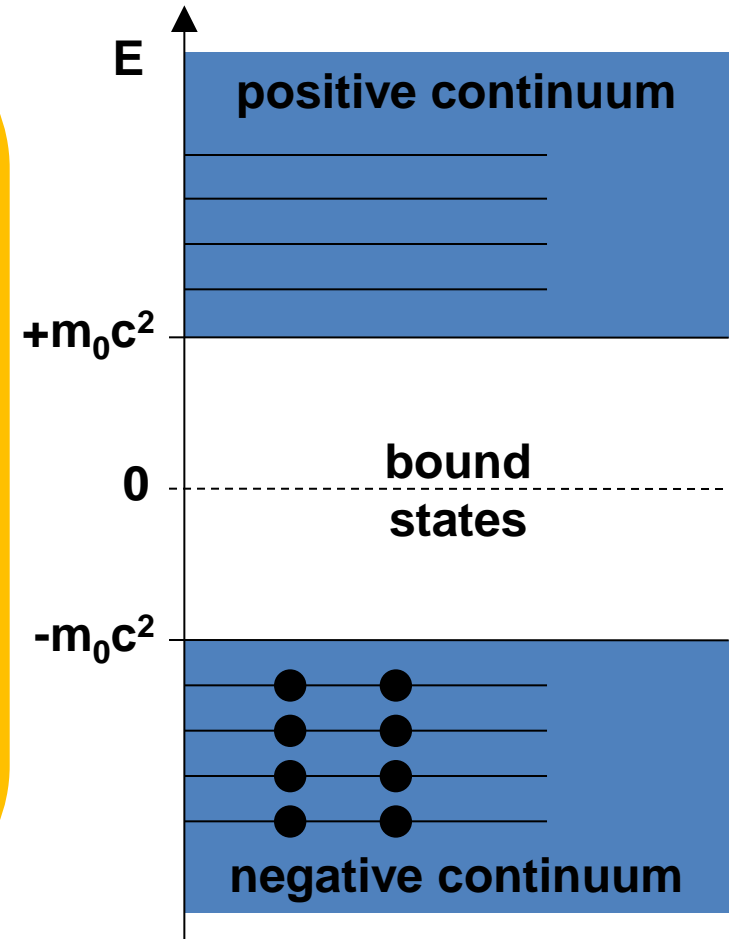
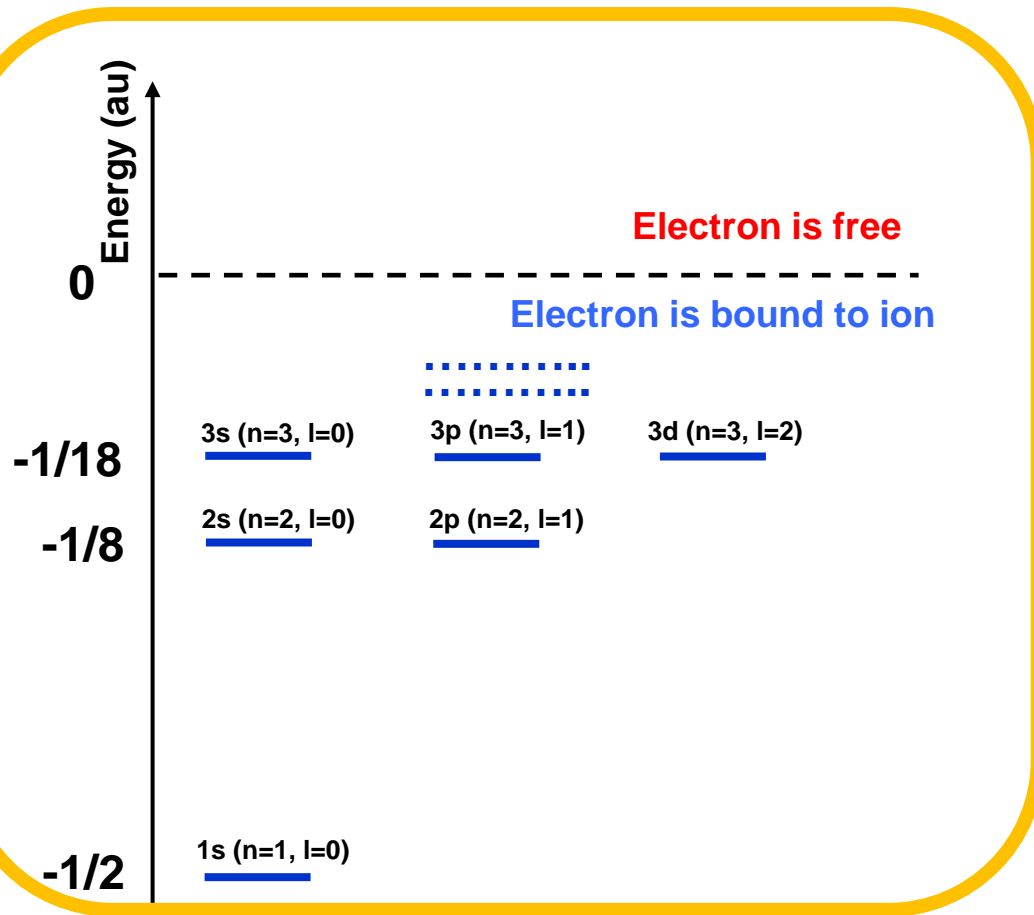
№ 28/14 7.12.
Deutschland 4,2014



DAS TOR ZU EINER ANDEREN WELT

Physiker
entschlüsseln
das Geheimnis der
Anti-Materie

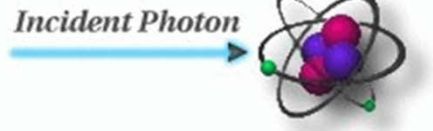
The Theorie of Matter and Anti-Matter



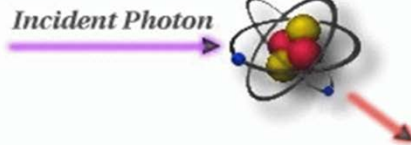
Lecture 4 Interaction of Photons with Matter

Interaction of photons and matter x-ray and gamma-regime

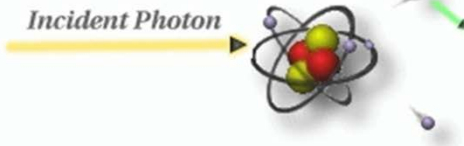
photo-effect / photo-absorption



Compton-scattering

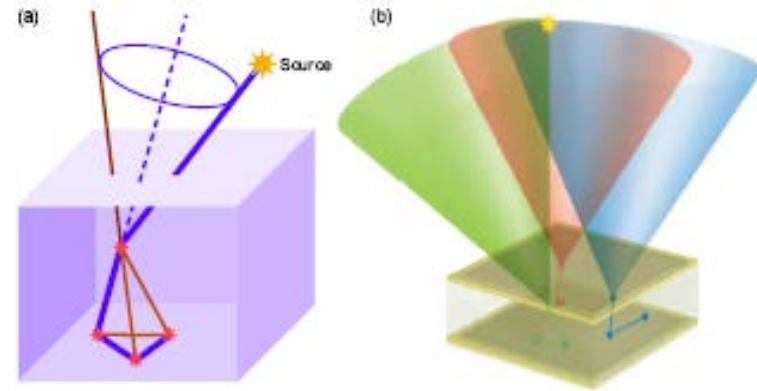


Pair production

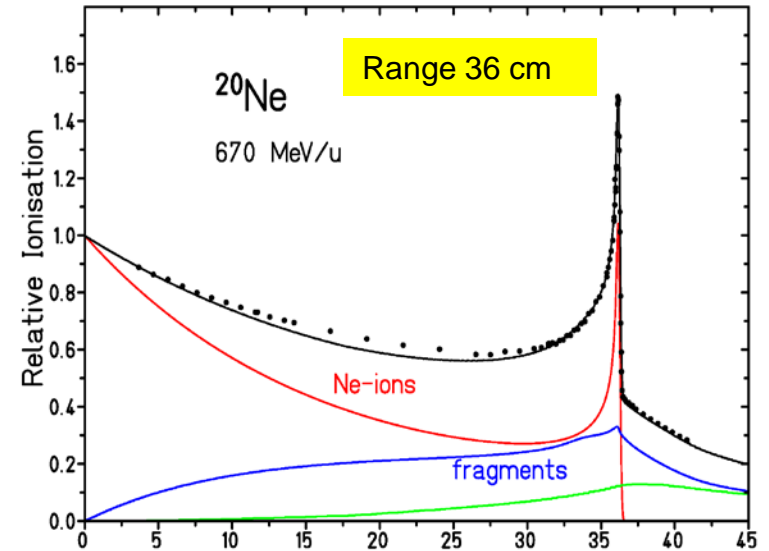
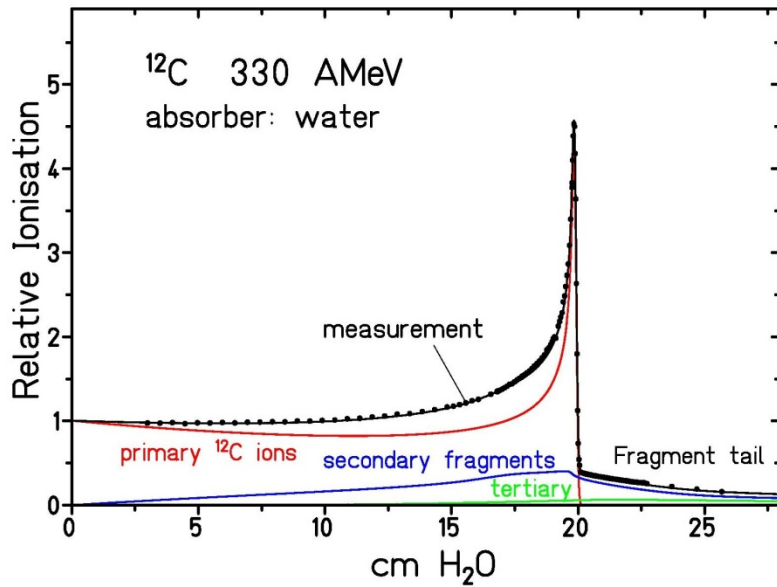


$$\hbar\omega' = \frac{\hbar\omega}{1 + \frac{\hbar\omega}{mc^2}(1 - \cos\theta)}$$

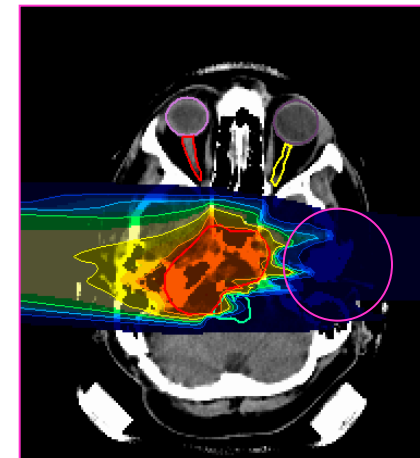
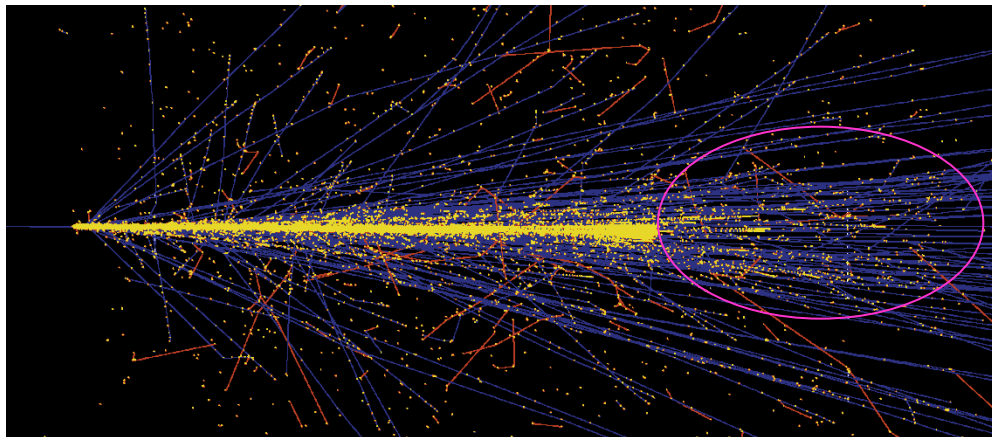
Compton Camera



Lecture 5 Interaction of Charged Particles with Matter

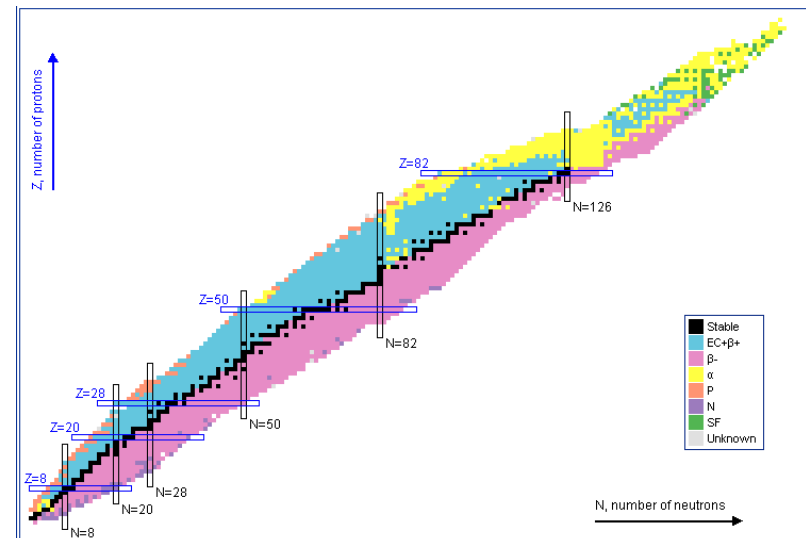
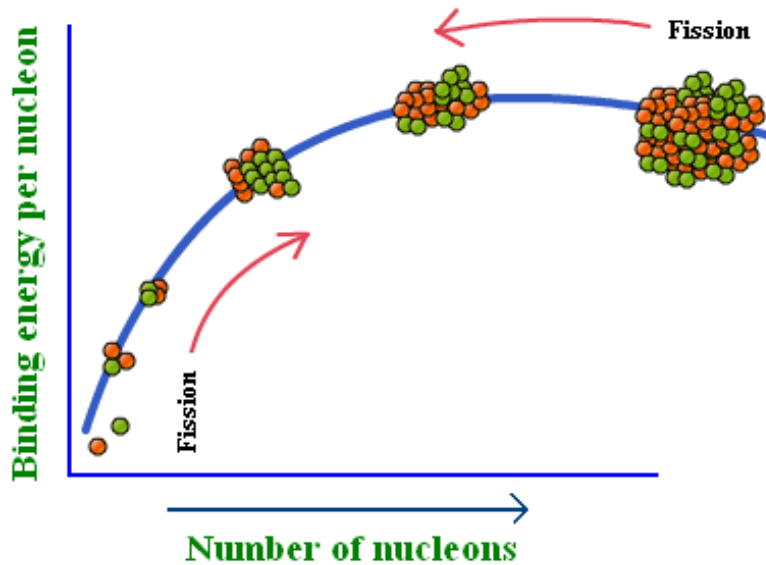
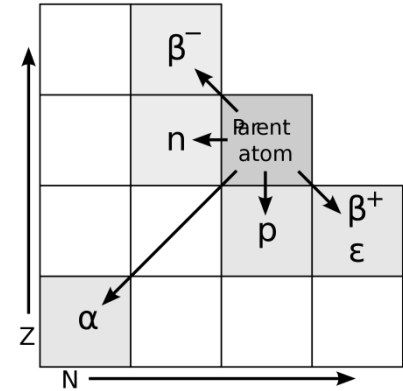
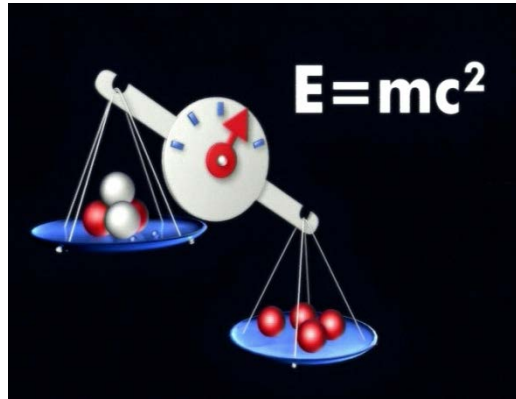


High-energy carbon beam stopping in water

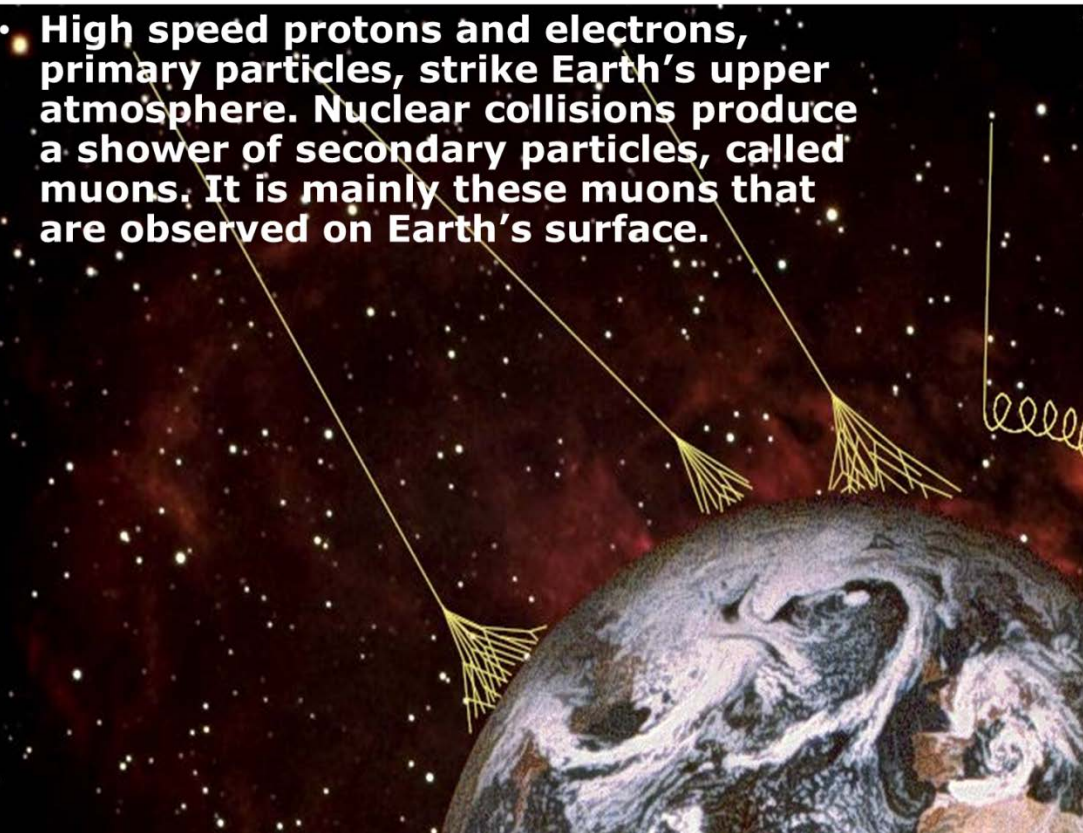


Lecture 7

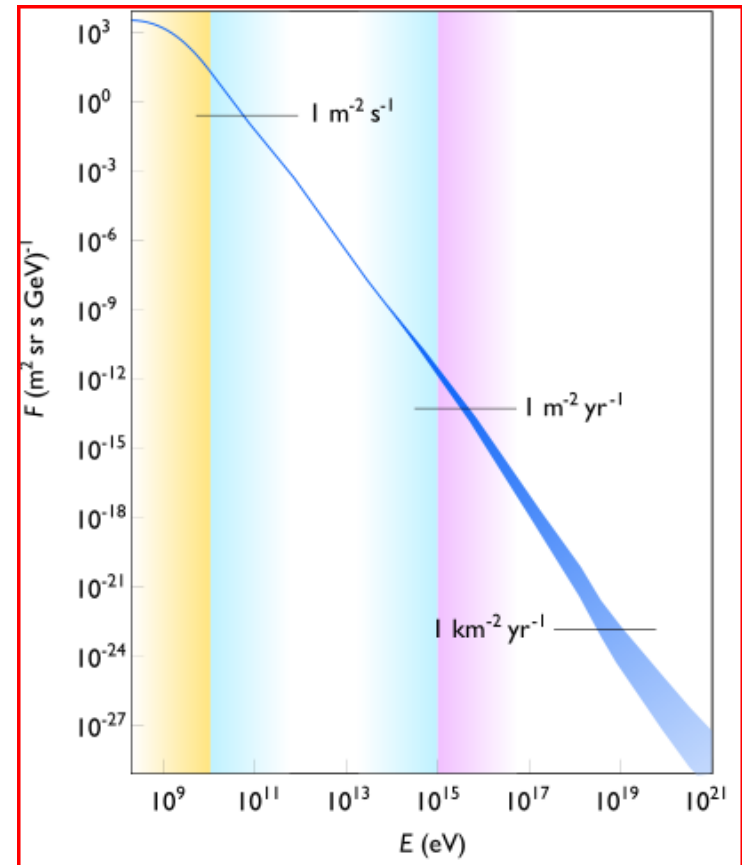
Nuclei and Their Decay Modes



- High speed protons and electrons, primary particles, strike Earth's upper atmosphere. Nuclear collisions produce a shower of secondary particles, called muons. It is mainly these muons that are observed on Earth's surface.

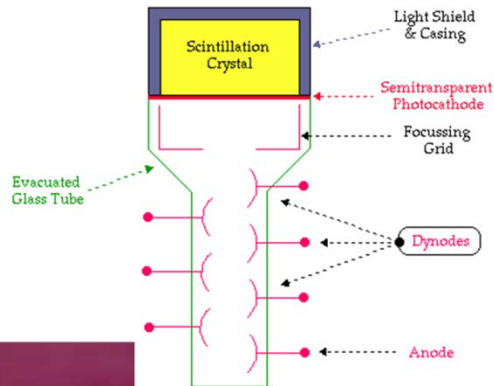


Flux: number of particles per unit area per second.

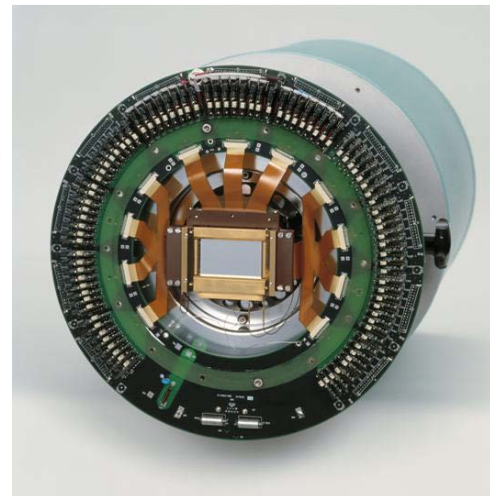
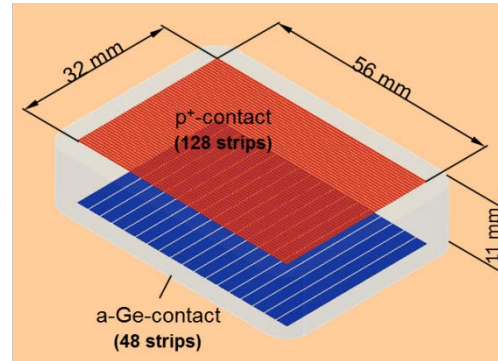


Lecture 9 Photon-, x-ray-, gamma-detectors

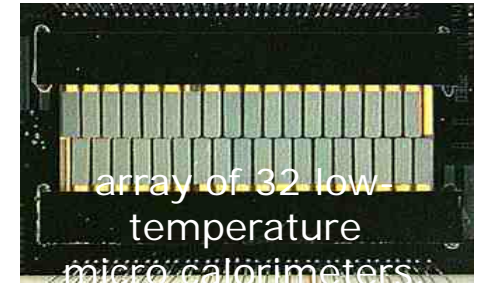
Inorganic Scintillators



Semi-conductors

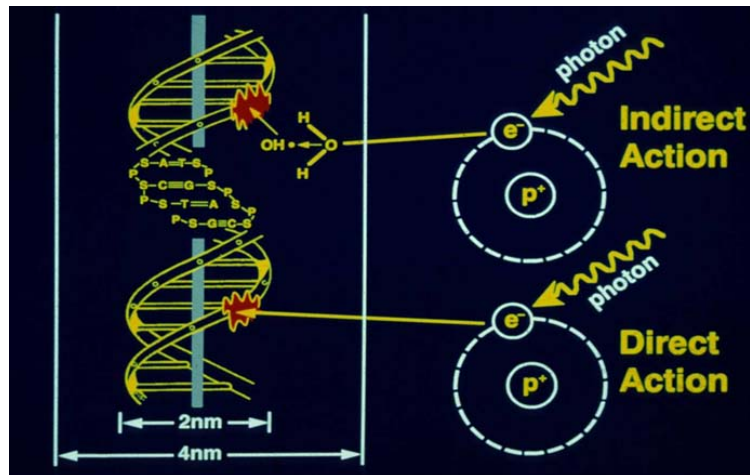
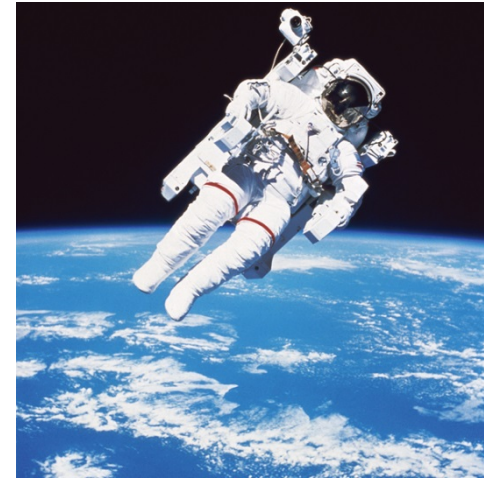


micro-calorimeters



Lecture 11 Radiation and their Biological Effectiveness

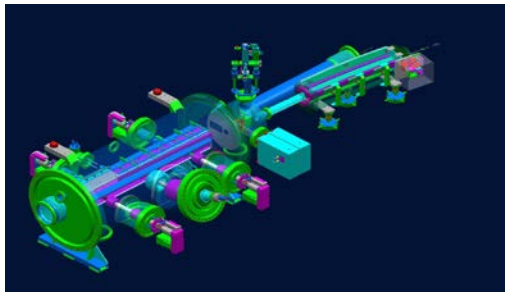
- Radiobiology
 - Acute (deterministic) effects
 - Late (stochastic) effects
- Heavy ions
 - Space radiation
- Radiotherapy
 - Conventional X-ray therapy
 - Particle therapy



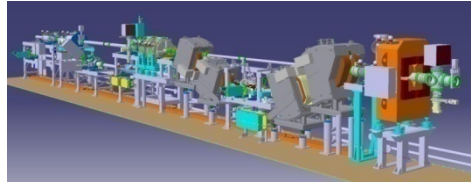
Lecture 13/14 Accelerators

Ion/electron sources, linacs, synchrotrons

$$\vec{F}_L = q(\vec{v} \times \vec{B})$$



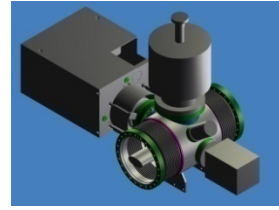
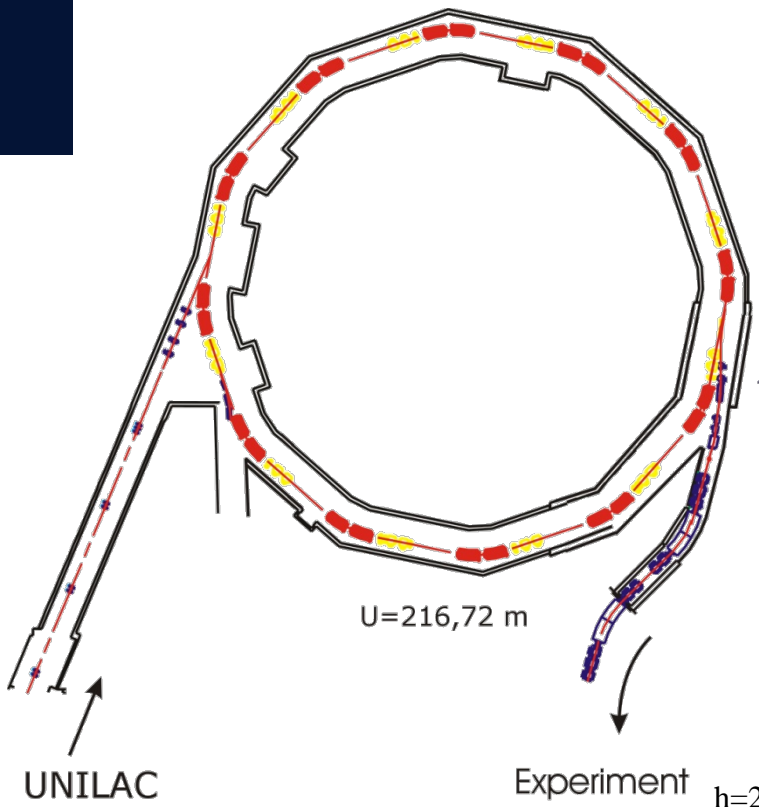
Injection system for low charged state heavy ions



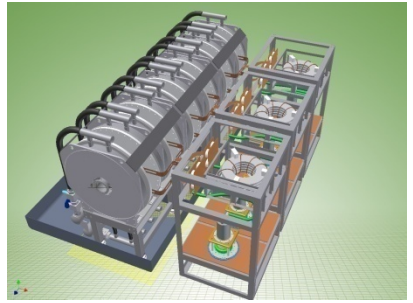
Charge separator for higher intensity and high quality beams



Power grid connection



Scrapers and NEG coating for pressure stabilization



h=2 acceleration cavity for faster ramping