# Strengths of the holographic approach at high baryon densities

#### EMMI Nuclear and Quark Matter Seminar, GSI, Darmstadt July 25th, 2018



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## Experiments





## Teaser



#### **Problems at strong coupling**

- non-perturbative regime
- lattice sign problem
- far from equilibrium
- near critical point





#### Goals

- compare to experimental data
- understand physical mechanisms

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Strengths of the holographic approach at high baryon densities

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# Outline

1. Holography primer



#### 2.1 Phase transitions and critical points

- 2.2 Mesons at finite densities
- 2.3 Magnetic field anisotropic hydrodynamics
- 2.4 Chiral transport effects
- 2.5 Holographic heavy ion collision correlations

#### 3. Discussion

2. Results



# 1. Holography primer





#### **Basic idea**

Assume we have a hard problem that is difficult to solve in a given theory, for example the **standard model** 



model or effective description

(Hard) problem in "similar" theory holography (gauge/gravity correspondence)



Simple problem in a particular gravitational theory



## **Basic idea**

Assume we have a hard problem that is difficult to solve in a given theory, for example the **standard model** 

gravity dual to QCD
or standard model?
not known yet

model or effective description

(Hard) problem in "similar" theory holography (gauge/gravity correspondence)



Simple problem in a particular gravitational theory

Holography is good at predictions that are qualitative or universal.



# **Gauge/Gravity concepts**

The Gauge/Gravity correspondence is based on the holographic principle. ['t Hooft (1993)]  $S_{max}$ (volume)  $\propto$  surface area







# **Gauge/Gravity concepts**





## Gauge/Gravity Correspondence









## Gauge/Gravity Correspondence









## Gauge/Gravity Correspondence





## Example: Reissner-Nordström black hole

N=4 Super-Yang-Mills theory at nonzero temperature & charge

$$T = r_H^2 \frac{|f'(r_H)|}{4\pi}$$

metric: 
$$ds^2 = \frac{r^2}{L^2} \left( -fdt^2 + d\vec{x}^2 \right) + \frac{L^2}{r^2 f} dr^2$$
  
 $f(r) = 1 - \frac{mL^2}{r^4} + \frac{q^2 L^2}{r^6}$ 



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## Example: Reissner-Nordström black hole

metric & gauge field N=4 Super-Yang-Mills correspondence defining a RN black theory at nonzero brane (solve Einsteintemperature & charge Maxwell eq's) metric:  $ds^2 = \frac{r^2}{L^2} \left( -fdt^2 + d\vec{x}^2 \right) + \frac{L^2}{r^2 f} dr^2$ QFT temperature: ← →  $T = r_H^2 \frac{|f'(r_H)|}{4\pi}$  $f(r) = 1 - \frac{mL^2}{r^4} + \frac{q^2L^2}{r^6}$ conserved charge Q, gauge field: radial AdS  $A_t = \mu - \frac{Q}{Lr^2}$ thermodynamically coordinate dual to chemical potential: black  $\mu = \frac{\sqrt{3q}}{2r_{\rm ex}^2}$ 



## 2. Results

#### 2.1 Phase transitions and critical points

2.2 Mesons at finite densities

- 2.3 Magnetic field anisotropic hydrodynamics
- 2.4 Chiral transport effects

2.5 Holographic heavy ion collision - correlations



Phase diagram for holographic quarks at nonzero baryon potential [O'Bannon, Karch; (2007)]



#### Phase diagram for holographic quarks at nonzero isospin potential



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#### Phase diagram for holographic quarks at nonzero isospin potential



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Matthias Kaminski

## 2. Results

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#### 2.2 Results - Holographic mesons at finite densities

[Erdmenger, Kaminski, Rust; PRD (2007)]



#### 2.2 Results - Holographic mesons at finite densities

Vector meson triplet splitting:







## 2. Results

- 2.1 Phase transitions and critical points
- 2.2 Mesons at finite densities

#### 2.3 Magnetic field — anisotropic hydrodynamics

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## Universal magnetoresponse in QCD ...

[Endrödi, Kaminski, Schäfer, Wu, Yaffe; arXiv:1806.09632]



Lattice QCD with 2+1 flavors, dynamical quarks, physical massestransverse pressure: $p_{\rm T} = -\frac{L_{\rm T}}{V} \frac{\partial F_{\rm QCD}}{\partial L_{\rm T}}$  $F_{\rm QCD} \dots$  free energytransverse pressure: $p_{\rm T} = -\frac{L_{\rm L}}{V} \frac{\partial F_{\rm QCD}}{\partial L_{\rm T}}$  $L_{\rm T} \dots$  transverse system sizelongitudinal pressure: $p_{\rm L} = -\frac{L_{\rm L}}{V} \frac{\partial F_{\rm QCD}}{\partial L_{\rm L}}$  $L_{\rm L} \dots$  longitudinal system size

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#### ... and N=4 Super-Yang-Mills theory

[Endrödi, Kaminski, Schäfer, Wu, Yaffe; arXiv:1806.09632]





# Polarized matter in strong magnetic field

Generating functionals  $W \sim P$  (pressure) for thermodynamics  $B \sim \mathcal{O}(1)$ 

[Kovtun; JHEP (2016)]

$$T^{\mu\nu} = Pg^{\mu\nu} + (Ts + \mu\rho)u^{\mu}u^{\nu} + T^{\mu\nu}_{EM}$$

$$J^{\alpha} = \rho u^{\alpha} - \sum_{bound \ current} T^{\mu\nu}_{EM} = M^{\mu\alpha}g_{\alpha\beta}F^{\beta\nu} + u^{\mu}u^{\alpha} \left(M_{\alpha\beta}F^{\beta\nu} - F_{\alpha\beta}M^{\beta\nu}\right)$$
[Israel; Gen.Rel.Grav. (1978)]
Polarization tensor:
$$M_{\mu\nu} = p_{\mu}u_{\nu} - p_{\nu}u_{\mu} - \epsilon_{\mu\nu\rho\sigma}u^{\rho}m^{\sigma} \qquad M^{\mu\nu} = 2\frac{\partial P}{\partial F_{\mu\nu}}$$
Including vorticity:
$$W \sim M_{\omega}B \cdot \omega$$
[Kovtun, Hernandez; JHEP (2017)]



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Including vorticity:
$$W \sim M_{\omega}B \cdot \omega$$
[Koutun, Hernandez; JHEP (2017)]
Leads to (anisotropic) hydrodynamics
and new transport effects:
(incomplete list)
[Ammon, Grieninger, Kaminski,
Koirala, Leiber, Wu; to appear]
$$\frac{\eta_{\perp}}{\mu}$$
perpendicular shear viscosity
$$\frac{\eta_{\perp}}{\psi}$$
perpendicular Hall viscosity
$$\frac{\eta_{\perp}}{\psi}$$
perpendicular conductivity
$$\frac{\eta_{\perp}}{\psi}$$

$$\frac{\eta_{\perp}}{\psi$$



# (Anisotropic) hydrodynamics in strong B

 $B \sim \mathcal{O}(1)$ 

 (anisotropic) hydrodynamics as single framework allows for polarization, magnetization, external vorticity, *E*, *B* [Kovtun, Hernandez; JHEP (2017)]
 [Ammon, Grieninger, Kaminski, Koirala, Leiber, Wu; to appear]

new transport effects (and Kubo formulae)

opportunity: dynamical *E* and *B*; magnetohydrodynamics [Kovtun, Hernandez; JHEP (2017)]

opportunity: study equilibrium and near-equilibrium transport



## Reminder: holography far beyond hydrodynamics

Example: 3+1-dimensional N=4 Super-Yang-Mills theory; poles of shear correlation function





Strengths of the holographic approach at high baryon densities

[Starinets; JHEP (2002)]



#### 2. Results

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#### **2.4 Chiral transport effects**

2.5 Holographic heavy ion collision - correlations



#### String theory prediction of new transport in hydrodynamics

[Erdmenger, Haack, Kaminski, Yarom; JHEP (2008)] [Banerjee et al.; JHEP (2011)]





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**Condensed** matter experiments with Weyl-semi-metals confirm:

[Landsteiner; (2014)] [Li et al; (2014)] [Zhang et al; (2015)] [Cortijo, Ferreiros, Landsteiner, Vozmediano; (2015)]



#### Hydrodynamic proof of new current contributions

 $B \sim \mathcal{O}(\partial)$ 









## Hydrodynamic proof of new current contributions

 $B \sim \mathcal{O}(\partial)$ 

#### **New contributions** in electric current discovered:



Proof in hydrodynamics: [Son,Surowka PRL 2009]



[Erdmenger, Haack, Kaminski, Yarom; JHEP (2008)] [Banerjee et al.; JHEP (2011)]



Analytic result: chemical

potential important!

 $\xi = C\mu$ 

C ... anomaly coefficient

Gauge/Gravitational Holography - Strong Physics From Another Dimension Page

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## Zeroth order chiral magnetic effect -thermodynamic chiral currents





Strengths of the holographic approach at high baryon densities

 $B \sim C$ 

## **EFT result: strong B thermodynamics**

[Ammon, Kaminski et al.; JHEP (2017)] [Ammon, Leiber, Macedo; JHEP (2016)]

Strong B thermodynamics with anomaly in **thermodynamic frame**:



building on:

[Kovtun; JHEP (2016)] [Jensen, Loganayagam, Yarom; JHEP (2014)]



## Holographic result: thermodynamics

[Ammon, Kaminski et al.; JHEP (2017)]

Background solution: charged magnetic black branes

[D'Hoker, Kraus; JHEP (2009)] [Ammon, Leiber, Macedo; JHEP (2016)]

- external magnetic field
- charged plasma
- anisotropic plasma

Holographic thermodynamics  $\langle T^{\mu\nu} \rangle = \begin{pmatrix} -3u_4 & 0 & 0 & -4c_4 \\ 0 & -\frac{B^2}{4} - u_4 - 4w_4 & 0 & 0 \\ 0 & 0 & -\frac{B^2}{4} - u_4 - 4w_4 & 0 \\ -4c_4 & 0 & 0 & 8w_4 - u_4 \end{pmatrix}$   $\langle T^{\mu\nu}_{\rm EFT} \rangle = \begin{pmatrix} \epsilon_0 & 0 & 0 & \xi_V^{(0)}B \\ 0 & P_0 - \chi_{BB}B^2 & 0 & 0 \\ 0 & 0 & P_0 - \chi_{BB}B^2 & 0 \\ \xi_V^{(0)}B & 0 & 0 & P_0 \end{pmatrix} + \mathcal{O}(\partial)$  $\langle J^{\mu} \rangle = (\rho, 0, 0, p_1)$ .

$$\langle J_{\rm EFT}^{\mu} \rangle = \left( n_0, \, 0, \, 0, \, \xi_B^{(0)} B \right) + \mathcal{O}(\partial)$$

with near boundary expansion coefficients  $u_4, w_4, c_4, p_1$ 



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with near boundary expansion coefficients  $u_4, w_4, c_4, p_1$ 

## chiral thermodynamic equilibrium observables (currents) confirmed by holographic model

[Ammon, Grieninger, Kaminski, Koirala, Leiber, Wu; to appear]



Weak B hydrodynamics - poles of 2-point functions  $\langle T^{\mu\nu} T^{\alpha\beta} \rangle$ ,  $\langle T^{\mu\nu} J^{\alpha} \rangle$ ,  $\langle J^{\mu} T^{\alpha\beta} \rangle$ ,  $\langle J^{\mu} J^{\alpha} \rangle$ :

[Ammon, Kaminski et al.; JHEP (2017)] [Abbasi et al.; PLB (2016)] [Kalaydzhyan, Murchikova; NPB (2016)]

spin 1 modes under SO(2) rotations around B

$$\omega = -ik^2 \frac{\eta}{\epsilon_0 + P_0} +$$

former momentum diffusion modes

$$\begin{aligned} \mathbf{\mathfrak{s}}_0 &= s_0/n_0\\ \tilde{c}_P &= T_0 (\partial \mathbf{\mathfrak{s}}/\partial T)_P \end{aligned}$$



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#### spin 1 modes under SO(2) rotations around B

$$\omega = \mp \frac{Bn_0}{\epsilon_0 + P_0} - ik^2 \frac{\eta}{\epsilon_0 + P_0} + k \frac{Bn_0\xi_3}{(\epsilon_0 + P_0)^2} - \frac{iB^2\sigma}{\epsilon_0 + P_0}$$

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#### spin 0 modes under SO(2) rotations around B

$$\omega_{0} = v_{0} k - i D_{0} k^{2} + \mathcal{O}(\partial^{3})$$
 former charge  
diffusion mode

$$\omega_{+} = v_{+}\kappa - i\Gamma_{+}\kappa + O(0)$$
  
former  
$$\omega_{-} = v_{-}k - i\Gamma_{-}k^{2} + O(\partial^{3})$$
  
sound  
modes



Weak B hydrodynamics - poles of 2-point functions  $\langle T^{\mu\nu} T^{\alpha\beta} \rangle$ ,  $\langle T^{\mu\nu} J^{\alpha} \rangle$ ,  $\langle J^{\mu} T^{\alpha\beta} \rangle$ ,  $\langle J^{\mu} J^{\alpha} \rangle$ :

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spin 0 modes under SO(2) rotations around B  $\omega_{0} = v_{0} k - i D_{0} k^{2} + \mathcal{O}(\partial^{3}) \text{ former charge}_{diffusion mode}$   $\omega_{+} = v_{+} k - i \Gamma_{+} k^{2} + \mathcal{O}(\partial^{3})$   $\omega_{-} = v_{-} k - i \Gamma_{-} k^{2} + \mathcal{O}(\partial^{3}) \text{ former}_{modes}$   $\int \mathbf{a \ chiral \ magnetic \ wave}_{[Kharzeev, Yee; PRD (2011)]}$   $v_{0} = \frac{2BT_{0}}{\tilde{c}_{P}n_{0}} \left(\tilde{C} - 3C\mathfrak{s}_{0}^{2}\right)$   $D_{0} = \frac{w_{0}^{2}\sigma}{\tilde{c}_{P}n_{0}^{3}T_{0}}$ 

#### dispersion relations of hydrodynamic modes are heavily modified by anomaly and B



## Holographic result II: hydrodynamic poles

Fluctuations around charged magnetic black branes

[Ammon, Kaminski et al.; JHEP (2017)]

- Weak B: holographic results are in "agreement" with hydrodynamics.
- Strong *B*: holographic result in agreement with thermodynamics, and numerical result indicates that **chiral waves propagate at** ...



confirming conjectures and results in probe brane approach [Kharzeev, Yee; PRD (2011)]



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### **Application: hydrodynamics & neutron star kicks**



observation: neutron stars undergo a large momentum change (a kick)



Gauge/Gravitational Holography - Strong Physics Far From Equilibrium Page

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### **Application: hydrodynamics & neutron star kicks**



hydrodynamics: fluids with lefthanded and right-handed particles produce a current along magnetic field

> e.g. right/left-handed electrons, neutrinos, ...





observation: neutron stars undergo a large momentum change (a kick)



Gauge/Gravitational Holography - Strong Physics Far From Equilibrium Page

### Application: hydrodynamics & neutron star kicks



*hydrodynamics:* fluids with lefthanded and right-handed particles produce a **current** along magnetic field

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Gauge/Gravitational Holography - Strong Physics Far From Equilibrium

## 2. Results

- 2.1 Phase transitions and critical points
- 2.2 Mesons at finite densities
- 2.3 Magnetic field anisotropic hydrodynamics
- 2.4 Chiral transport effects

#### 2.5 Holographic heavy ion collision - correlations



## Holography far-from equilibrium

#### thermalization

*examples:* heavy ion collision



plasma at T>0



## Holography far-from equilibrium



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## Holography far-from equilibrium



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## Holographic heavy ion collision

Non-central collision: [Chesler, Yaffe; JHEP (2015)]



#### **Gravitational dual**

shock wave collision in Anti de Sitter space [Janik; PRD (2006)]

color coding: energy density red = high density blue = low density

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## Holographic plasma "hydrodynamizes" quickly

[Chesler, Yaffe; PRL (2011)]



Suggests existence of a new formulation of "hydrodynamics far from equilibrium". ——> Resurgence in QFT

confirmed by field theory: [Romatschke; (2016), (2017)]



## Holographic plasma thermalization results



Equilibration happens very fast, in agreement with experiment. Hydrodynamics works way before that.

cf. [Romatschke; (2016), (2017)]



## **Background - One Point Functions**

[Fuini, Yaffe; JHEP (2015)]





## **Correlations - finite charge, zero B**

Charged Anistropic Isotropization:

Longitudinal Correlation Comparison



## **Correlations - finite charge, finite B**

Magnetic Charged Anistropic Isotropization:

Longitudinal Correlations Non-equal Time



Assume we have a hard problem that is difficult to solve in a given theory, for example **QCD** 















Solve problems in

effective field



Holography is good at predictions that are qualitative or universal.

- Compare holographic result to hydrodynamics of model theory.
- Compare hydrodynamics of original theory to hydrodynamics of model.
- Understand holography as an effective description.

Solve problems in

effective field



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Solve problems in

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## 3. Discussion & Outlook

Holographic methods allow calculation of observables at high densities in strongly coupled quantum field theories
*phase diagrams, critical points, superconducting phase, ...*

• strong combination: effective theories, e.g. hydrodynamics, together with holography

discovered new transport effect(s): chiral vortical effect

• Effective quantitative results

shear viscosity over entropy density, universal pressure ratio at nonzero B

• Qualitative results

 $\rightarrow$  (chiral) equilibrium currents at nonzero B, $\mu$ , C Kubo transport formulae, ballpark values

• Introduction to holography



## 3. Discussion & Outlook

- Outlook:
  - transport coefficients (Kubo formulae, values) at strong magnetic field (anisotropic hydrodynamics) [Ammon, Grieninger, Kaminski, Koirala, Leiber, Wu; to appear]
  - correlations far from equilibrium at high density and magnetic field with chiral anomaly [Cartwright, Kaminski; to appear] see my talk at HoloQuark2018
  - dynamical electromagnetic fields magnetohydrodynamics
  - comparison to experimental data



### Collaborators

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Dr. Jackson Wu

Roshan Koirala



Strengths of the holographic approach at high baryon densities

Cartwright

### APPENDIX



## Hydrodynamic variables

#### Thermodynamics

 $T, \mu, u^{\nu}$ 

thermodynamic variables: temperature, chemical potential, fluid velocity



#### Hydrodynamics

 $T(t, \vec{x}), \, \mu(t, \vec{x}), \, u^{\nu}(t, \vec{x})$ 





## Hydrodynamic variables

#### Thermodynamics

 $T, \mu, u^{\nu}$ 

thermodynamic variables: temperature, chemical potential, fluid velocity



#### Hydrodynamics

$$T(t, \vec{x}), \mu(t, \vec{x}), u^{\nu}(t, \vec{x})$$

hydrodynamic fields -protected by symmetry




# Hydrodynamic variables

#### Thermodynamics

$$T(\vec{x}), \, \mu(\vec{x}), \, u^{\nu}(\vec{x})$$

thermodynamic variables: temperature, chemical potential, fluid velocity



#### Hydrodynamics

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hydrodynamic fields -protected by symmetry





# Hydrodynamics

effective field theory, expansion in gradients of fields



• constitutive equation (Landau frame)

[Jensen, Kaminski, Kovtun, Meyer, Ritz, Yarom; PRL (2012)]



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# Hydrodynamics

effective field theory, expansion in gradients of fields





 $\nabla_{\nu} j^{\nu} = 0 \quad \begin{array}{c} \text{classical} \\ \text{theory} \end{array}$ 



Derived for any theory with chiral anomaly

 $\nabla_{\mu} j^{\mu} = \underbrace{C \, \epsilon^{\nu \rho \sigma \lambda} F_{\nu \rho} F_{\sigma \lambda}}_{\text{theory}} \text{ quantum theory} \qquad (e.g. the standard model of particle physics)$ 

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Derived for any theory with chiral anomaly

 $\nabla_{\mu} j^{\mu} = \left( C \, \epsilon^{\nu \rho \sigma \lambda} F_{\nu \rho} F_{\sigma \lambda} \right) \begin{array}{c} \text{quantum} \\ \text{theory} \end{array}$ 

(e.g. the standard model of particle physics)

Generalized constitutive equation with external fields:

$$j^{\mu} = nu^{\mu} + \sigma E^{\mu} + \sigma^{B} B^{\mu} + \sigma^{V} \omega^{\mu} + \cdots \text{ Spravity prediction} \\ (non) \quad (ideal) \quad conduc- \\ (non) \quad (ideal) \quad conduc- \\ (non) \quad (ideal) \quad conduc- \\ (non) \quad conduc- \\ (non) \quad chiral \\ (non)$$



Derived for any theory with chiral anomaly

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Generalized constitutive equation with external fields:

theory



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(e.g. the standard model of particle physics)

Generalized constitutive equation with external fields:

theory





#### Dirty details: chiral effects in vector/axial currents

see e.g. [Jensen, Kovtun, Ritz; JHEP (2013)] [Neiman, Oz; JHEP (2010)]

Vector current (e.g. QCD U(1))

$$J_V^{\mu} = \dots + \xi_V \omega^{\mu} + \xi_{VV} B^{\mu} + \xi_{VA} B_A^{\mu}$$

$$\stackrel{\text{chiral}}{\underset{\text{effect}}{\text{magnetic}}}$$

Axial current (e.g. QCD axial U(1))

$$J_A^{\mu} = \dots + \xi \omega^{\mu} + \xi_B B^{\mu} + \xi_{AA} B_A^{\mu}$$

chiral vortical effect

chiral separation effect

Note:

\* hydrodynamic frame choice [Ammon, Kaminski et al.; JHEP (2017)]

\* consistent vs covariant

[Landsteiner; APhysPolC (2016)] [Landsteiner et al; JHEP (2011)]

see e.g. Juan Torres Rincon's talk



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see e.g. Juan Torres Rincon's talk



#### Dirty details: chiral effects in vector/axial currents

see e.g. [Jensen, Kovtun, Ritz; JHEP (2013)] [Neiman, Oz; JHEP (2010)]

Vector current (e.g. QCD U(1))

$$J_{V}^{\mu} = \dots + \xi_{V} \omega^{\mu} + \xi_{VV} B^{\mu} + \xi_{VA} B_{A}^{\mu}$$

$$\stackrel{\text{chiral magnetic effect}}{\text{Axial current (e.g. QCD axial U(1))}}$$

$$J_{A}^{\mu} = \dots + \xi \omega^{\mu} + \xi_{B} B^{\mu}$$

$$\stackrel{\text{chiral chiral separation effect}}{\text{chiral separation effect}} + \xi_{AA} B_{A}^{\mu}$$

Note:

\* hydrodynamic frame choice [Ammon, Kaminski et al.; JHEP (2017)]

#### \* consistent vs covariant

[Landsteiner; APhysPolC (2016) ] [Landsteiner et al; JHEP (2011) ]

see e.g. Juan Torres Rincon's talk



[Kaminski, Uhlemann, Schaffner-Bielich, Bleicher; PLB (2016)]

A bucket full of electrons and electron neutrinos with short mean free path



 $B = 0.1 \,\mathrm{MeV^2}$  $\mu^{\ell} \approx 300 \,\mathrm{MeV}$  $\langle p_{\nu} \rangle \approx \mu^{\ell}.$ 





[Kaminski, Uhlemann, Schaffner-Bielich, Bleicher; PLB (2016)]



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[Kaminski, Uhlemann, Schaffner-Bielich, Bleicher; PLB (2016)]



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[Kaminski, Uhlemann, Schaffner-Bielich, Bleicher; PLB (2016)]



Matthias Kaminski

neutror

star

[Kaminski, Uhlemann, Schaffner-Bielich, Bleicher; PLB (2016)]



Strengths of the holographic approach at high baryon densities

Matthias Kaminski

[Kaminski, Uhlemann, Schaffner-Bielich, Bleicher; PLB (2016)]



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# **Correlations - geodesic approximation**

[Balasubramanian, Ross; PRD(2000)]

Correlator as a sum over geodesics: 
$$\Delta L = L - L_{\text{thermalized}}$$
$$\langle \mathscr{O}(t, \vec{x}_1) \mathscr{O}(t, \vec{x}_2) \rangle = \int \mathcal{DP} e^{i\Delta \mathcal{L}(\mathcal{P})} \approx \sum_{\text{geodesics}} e^{-\Delta L} \approx e^{-\Delta L}$$
Geodesic length (Lagrangian):
$$L = \int d\lambda \sqrt{g_{\mu\nu} \dot{x}^{\mu} \dot{x}^{\nu}} \qquad \Rightarrow \qquad \frac{d^2 x^{\mu}}{d\sigma^2} + \Gamma^{\mu}_{\alpha\beta} \frac{dx^{\alpha}}{d\sigma} \frac{dx^{\beta}}{d\sigma} = 0$$
geodesic equation
$$\left( L = m \int d\lambda \sqrt{g_{\mu\nu} \dot{x}^{\mu} \dot{x}^{\nu}} + \frac{q}{-} A_{\mu} \dot{x}^{\mu} \qquad \text{charged probe particle} \right)$$

$$L = m \int d\lambda \sqrt{g_{\mu\nu} \dot{x}^{\mu} \dot{x}^{\nu}} + \frac{q}{m} A_{\mu} \dot{x}^{\mu} \qquad \text{charged probe particle} \\ Lorentz force term$$

Numerical implementation - relaxation method:

[Ecker, Grumiller, Stricker; JHEP (2015)]

- 1. Generate the dynamic background
- 2. Generate interpolations of the metric functions
- 3. Discretize the geodesic equations using a relaxation scheme
- 4. Approximate the proper length using a Riemann sum

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[Bali, Bruckmann, Endrödi, Katz, Schäfer; JHEP (2014)] [Fuini, Yaffe, JHEP (2015)]

Total action:

 $S = S_{\text{QCD}}(e, B) + S_{\text{EM}}(e, B)$ 

OCD action coupled to external magnetic field (through covariant derivative) (not part of the dynamics)

action for external magnetic field; not included in code



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[Bali, Bruckmann, Endrödi, Katz, Schäfer; JHEP (2014)] [Fuini, Yaffe, JHEP (2015)]

Total action:

$$S = S_{\text{QCD}}(e, B) + S_{\text{EM}}(e, B)$$

QCD action coupled to external magnetic field (through covariant derivative)

action for external magnetic field; not included in code (not part of the dynamics)

# Free energy: $F = -T \log \mathcal{Z}[S]$ = $F_{\text{QCD}}(e, B) + F_{\text{EM}}(e, B)$ $F_{\text{EM}}(e, B) = -V \frac{B^2}{2e^2}$



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[Bali, Bruckmann, Endrödi, Katz, Schäfer; JHEP (2014)] [Fuini, Yaffe, JHEP (2015)]

Total action:

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QCD action coupled to external magnetic field (through covariant derivative) action for external magnetic field; not included in code (not part of the dynamics)

Free energy: 
$$F = -T \log \mathcal{Z}[S]$$
  
=  $F_{\text{QCD}}(e, B) + F_{\text{EM}}(e, B)$   $F_{\text{EM}}(e, B) = -V \frac{B^2}{2e^2}$ 

Transverse 
$$p_{\rm T} = -\frac{L_{\rm T}}{V} \frac{\partial F_{\rm QCD}(e, B)}{\partial L_{\rm T}}$$

Electric charge is renormalization  $e^2(\mu) = Z_e(\mu) e_0^2$ ,  $Z_e(\mu) = 1 + 2b_1 e^2 \log \frac{\mu}{\Lambda}$ ,  $\mu = \sqrt{c_T T^2 + c_L \Lambda_H^2 + c_B |B|}$ scale dependent:

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[Bali, Bruckmann, Endrödi, Katz, Schäfer; JHEP (2014)] [Fuini, Yaffe, JHEP (2015)]

Total action:

$$= S_{\rm QCD}(e,B) + \\ {}_{\rm QCD\ action\ coupled\ to} \\ external\ magnetic\ field \\ (through\ covariant\ derivative) \end{cases}$$

 $S_{\rm EM}(e,B)$ action for external magnetic field; not included in code (not part of the dynamics)

Free energy: 
$$F = -T \log \mathcal{Z}[S]$$

 $=F_{\text{QCD}}(e,B)+F_{\text{EM}}(e,B) \qquad F_{\text{EM}}(e,B)=-V\frac{B^2}{2e^2}$ 

 $\begin{array}{ll} \mbox{Transverse} \\ \mbox{pressure:} \end{array} p_{\rm T} = - \frac{L_{\rm T}}{V} \frac{\partial F_{\rm QCD}(e,B)}{\partial L_{\rm T}} \end{array}$ 

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[Bali, Bruckmann, Endrödi, Katz, Schäfer; JHEP (2014)] [Fuini, Yaffe, JHEP (2015)]

Total action:

$$= S_{\rm QCD}(e,B) + S_{\rm P}$$

$$= S_{\rm QCD}(e,$$

 $S_{\rm EM}(e,B)$ action for external magnetic field; not included in code (not part of the dynamics)

Free energy: 
$$F = -T \log \mathcal{Z}[S]$$
  
= $F_{\text{QCD}}(e, B) + F_{\text{EM}}(e, B)$   $F_{\text{EM}}(e, B) = -V \frac{B^2}{2e^2}$ 

Transverse pressure:  $p_{\rm T} = -\frac{L_{\rm T}}{V} \frac{\partial F_{\rm QCD}(e,B)}{\partial L_{\rm T}}$ 

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[Bali, Bruckmann, Endrödi, Katz, Schäfer; JHEP (2014)] [Fuini, Yaffe, JHEP (2015)]

Total action:

$$= S_{\rm QCD}(e,B) + S_{\rm P}$$
QCD action coupled to
external magnetic field
otherwariant derivative)

EM(e, B)included in code t part of the dynamics)

Free energy: 
$$F = -T \log \mathcal{Z}[S]$$
  
= $F_{\text{QCD}}(e, B) + F_{\text{EM}}(e, B)$   $F_{\text{EM}}(e, B) = -V \frac{B^2}{2e^2}$ 

Transverse 
$$p_{\rm T} = -\frac{L_{\rm T}}{V} \frac{\partial F_{\rm QCD}(e, B)}{\partial L_{\rm T}}$$

this free energy is renormalization scale dependent



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[Bali, Bruckmann, Endrödi, Katz, Schäfer; JHEP (2014)] [Fuini, Yaffe, JHEP (2015)]

Total action:

$$= S_{\rm QCD}(e,B) + S_{\rm E}$$
action coupled to
external magnetic field
(not i)
(not j)

 $S_{\rm EM}(e,B)$ action for external magnetic field; not included in code (not part of the dynamics)

Free energy: 
$$F = -T \log \mathcal{Z}[S]$$
  
= $F_{\text{QCD}}(e, B) + F_{\text{EM}}(e, B)$   $F_{\text{EM}}(e, B) = -V \frac{B^2}{2e^2}$ 

Transverse pressure:

$$p_{\rm T} = -\frac{L_{\rm T}}{V} \frac{\partial F_{\rm QCD}(e, B)}{\partial L_{\rm T}}$$

this free energy is renormalization scale dependent

hence this pressure is renormalization scale dependent



Matthias Kaminski

#### How to compare QCD to Super-Yang-Mills

SYM action: 
$$S = S_{SYM}(e, \mathcal{B}) + S_{EM}(e, \mathcal{B})$$

SYM field content: fermions, scalar particles, vector field

SYM properties: conformal symmetry, supersymmetry, ...

SYM appears to be entirely different from QCD!



#### How to compare QCD to Super-Yang-Mills

SYM action: 
$$S = S_{SYM}(e, \mathcal{B}) + S_{EM}(e, \mathcal{B})$$

SYM field content: fermions, scalar particles, vector field

SYM properties: conformal symmetry, supersymmetry, ...

#### SYM appears to be entirely different from QCD!

#### Strategy:

- compare thermodynamic quantities (macroscopic / effective); e.g. pressure
- match divergencies in the two theories, i.e. match beta functions
- measure magnetic fields in "same units"
- compare two theories at same renormalization scale

SYM magnetic field  ${\cal B}$  vs. QCD magnetic field B: B =



# How to choose a holographic model?

The same way, we choose a hydrodynamic model:

- match symmetries (and anomalies)
- include interesting operators

depends on the physical question

• match magnetic properties



# How to choose a holographic model?

The same way, we choose a hydrodynamic model:

- match symmetries (and anomalies)
- include interesting operators depends on the physical question
- match magnetic properties

Einstein-Maxwell-Chern-Simons gravity has dual with: *cf. talk by K. Landsteiner* 

- chiral anomaly, breaking a U(1) axial symmetry
- axial current and energy momentum tensor *chiral magnetic transport*
- thermodynamics match well (in external B field)

$$S_{grav} = \frac{1}{2\kappa^2} \left[ \int_{\mathcal{M}} \mathrm{d}^5 x \sqrt{-g} \left( R + \frac{12}{L^2} - \frac{1}{4} F_{mn} F^{mn} \right) - \frac{\gamma}{6} \int_{\mathcal{M}} A \wedge F \wedge F \right]$$

dual to N=4 Super-Yang-Mills theory coupled to U(1)



# How to choose a holographic model?

The same way, we choose a hydrodynamic model:

- match symmetries (and anomalies)
- include interesting operators

```
depends on the physical question
```

• match magnetic properties



Successful example: holographic model discovering chiral vortical effect (2008)

Erdmenger, Haack, Kaminski, Yarom; JHEP (2009)]

[Banerjee et al; JHEP (2011)]

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ing a U(1) axial symmetry
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gy momentum tensor *chiral magnetic transport* 

h well (in external B field)

$$S_{grav} = \frac{1}{2\kappa^2} \left[ \int_{\mathcal{M}} \mathrm{d}^5 x \sqrt{-g} \left( R + \frac{12}{L^2} - \frac{1}{4} F_{mn} F^{mn} \right) - \frac{\gamma}{6} \int_{\mathcal{M}} A \wedge F \wedge F \right]$$

dual to N=4 Super-Yang-Mills theory coupled to U(1)



# EFT result III: weak B details

Weak B hydrodynamics - poles of 2-point functions:

[Ammon, Kaminski et al.; JHEP (2017)] [Abbasi et al.; PLB (2016)]

**spin 0 modes under SO(2) rotations around B** [Kalaydzhyan, Murchikova; NPB (2016)]