

# Preliminaries

## Introduction to Nuclear Science

Simon Fraser University  
SPRING 2011

NUCS 342 — January 6, 2011



# Outline

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# Useful links

- For chemistry web site see [this link](#).
- For class website see [this link](#).
- For general information on NuSc 342 see [this link](#).
- For detailed NuSc 342 schedule see [this link](#).
- For useful NuSc 342 numerical constants [this link](#).

# Nuclear Science Minor

Minor in Nuclear Science requires completion of 14 upper division credits from the following courses:

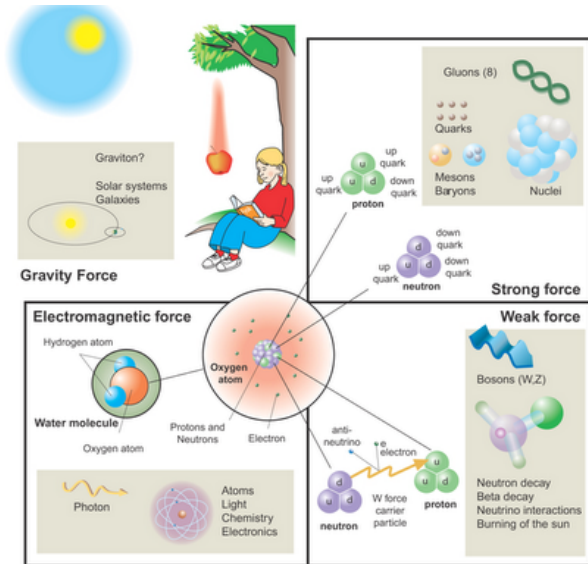
- NuSc 341-3 Introduction to Radiochemistry
- NuSc 342-3 Introduction to Nuclear Science
- NuSc 344-3 Nucleosynthesis and Distribution of the Elements
- NuSc 346-2 Radiochemistry Laboratory ← unique in Canada
- PHYS 385-3 Quantum Physics
- NuSc 444-3 Special Topics in Nuclear Science
- CHEM 482-3 Directed Study in Advanced Topics of Chemistry
- PHYS 485-3 Particle Physics

NuSc courses are currently being taught by Profs C. Andreoiu, J.C. Brodovich, and K. Starosta, PHYS courses by Prof. M. Vetterli.

# Scales in the Universe

For a great applet showing scales in the Universe see [this link](#).

# Fundamental Forces



# Fundamental Forces

| Force           | Range [m]  | Relative strength |                | Impact   |
|-----------------|------------|-------------------|----------------|----------|
|                 |            | within nucleus    | beyond nucleus |          |
| Strong          | $10^{-15}$ | 100               | 0              | Nuclei   |
| Electromagnetic | $\infty$   | 1                 | 1              | Chem/Bio |
| Weak            | $10^{-18}$ | $10^{-5}$         | 0              | Nuclei   |
| Gravity         | $\infty$   | $10^{-43}$        | $10^{-43}$     | Universe |

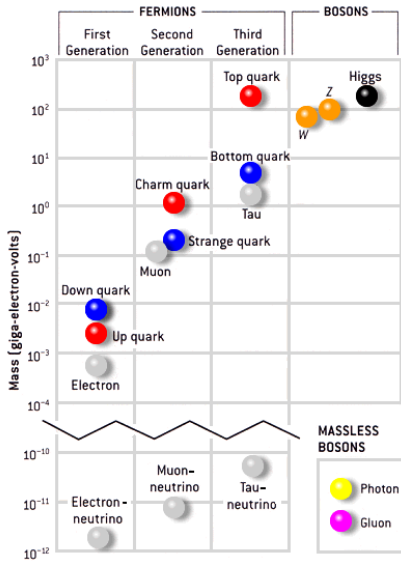
# The Standard Model of Elementary Particles

Three Generations  
of Matter (Fermions)

|                | I  | II   | III  |                                      |
|----------------|--|--|--|--------------------------------------|
| mass →         | 2.4 MeV  | 1.27 GeV                                     | 171.2 GeV                                    | 0                                    |
| charge →       | $\frac{2}{3}$                                  | $\frac{2}{3}$                                | $\frac{2}{3}$                                | 0                                    |
| spin →         | $\frac{1}{2}$                                  | $\frac{1}{2}$                                | $\frac{1}{2}$                                | 1                                    |
| name →         | <b>u</b><br>up                                 | <b>c</b><br>charm                            | <b>t</b><br>top                              | <b><math>\gamma</math></b><br>photon |
|                | 4.8 MeV  | 104 MeV                                      | 4.2 GeV                                      | 0                                    |
|                | $-\frac{1}{3}$                                 | $-\frac{1}{3}$                               | $-\frac{1}{3}$                               | 0                                    |
|                | $\frac{1}{2}$                                  | $\frac{1}{2}$                                | $\frac{1}{2}$                                | 1                                    |
| <b>Quarks</b>  | <b>d</b><br>down                               | <b>s</b><br>strange                          | <b>b</b><br>bottom                           | <b>g</b><br>gluon                    |
|                | <2.2 eV  | <0.17 MeV                                    | <15.5 MeV                                    | 91.2 GeV                             |
|                | 0  | 0  | 0  | 0                                    |
|                | $\frac{1}{2}$                                  | $\frac{1}{2}$                                | $\frac{1}{2}$                                | 1                                    |
|                | <b><math>\nu_e</math></b><br>electron neutrino | <b><math>\nu_\mu</math></b><br>muon neutrino | <b><math>\nu_\tau</math></b><br>tau neutrino | <b>Z<sup>0</sup></b><br>weak force   |
|                | 0.511 MeV                                      | 105.7 MeV                                    | 1.777 GeV                                    | 80.4 GeV                             |
|                | -1   | -1   | -1   | $\pm 1$                              |
|                | $\frac{1}{2}$                                  | $\frac{1}{2}$                                | $\frac{1}{2}$                                | 1                                    |
| <b>Leptons</b> | <b>e</b><br>electron                           | <b><math>\mu</math></b><br>muon              | <b><math>\tau</math></b><br>tau              | <b>W<sup>±</sup></b><br>weak force   |

**Bosons (Forces)**

# Masses of Elementary Particles

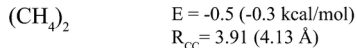
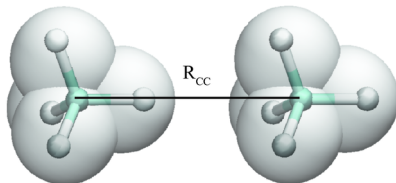
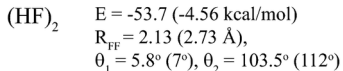
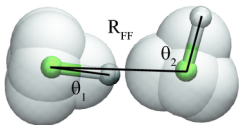
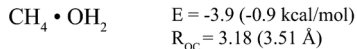
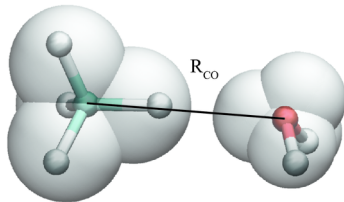
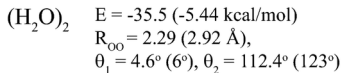
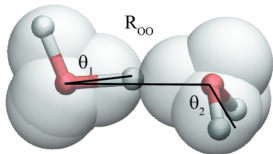


## Effective Interactions between molecules

- In complex systems interactions between fundamental constituents do not always provide the most effective description of the system.
- A known example in Chemistry is provided by the Van der Waals forces between dimer molecules.
- For a description of dimer interactions it is much more efficient to use effective Van der Waals forces between molecules than the fundamental electromagnetic forces between constituent atoms.
- The effective Van der Waals force is a residual force resulting from a superposition of the fundamental electromagnetic (predominantly Coulomb) forces.
- The effective force is weak, since the dimers are neutral. The interactions are complex as they arise due to dipole and higher-order moments either resulting from the molecular structure or induced by the intermolecular interactions.



# Van der Waals forces



# Effective Interactions in nuclei

- In principle nuclei can be viewed as system of quarks bound by the strong force mediated by gluons.
- In practise, this is a very inefficient way to describe low-energy (up to  $\sim 100$  MeV) nuclear excitations.
- In the low-energy range nuclei are much better described as systems composed of protons and neutrons.
- Since protons and neutrons are made up from three quarks the effective nuclear interactions is a (complex) superposition of strong interactions between these constituent quarks.
- One of the goals of nuclear science is to describe in an efficient way effective interactions between nucleons and provide predictive theory for nuclear structure and reactions.

# Degrees of Freedom

- Degrees of freedom in NuSc 342 will be understood as the constituents of a system chosen for the description of system properties.
- For the van der Waals example above one can either choose molecules, atoms, nuclei and electrons or even quarks and leptons as degrees of freedom.
- Choice of degrees of freedom is crucial, it should capture the most important properties of the system without bringing unnecessary computational complications.
- In NuSc 342 the degrees of freedom are: nucleons (protons and neutrons), alpha particles, atomic nuclei as well as electrons, positrons, and neutrinos.

# Lorentz transformation

- Describes a relationship between a position and time in a stationary  $(x, t)$  and moving  $(x', t')$  reference frames.
- Plays a role in the description of nuclear reactions which happen at speeds which are on the order of a few percent of the speed of light ( $c = 3 \times 10^8$  [m/s]) with  $(x, t)$  and  $(x', t')$  representing the target (laboratory) and projectile (beam) reference frames, respectively.

$$t \neq t'$$

$$x = \gamma(x' + vt')$$

$$\beta = \frac{v}{c}$$

$$x' = \gamma(x - vt)$$

$$ct = \gamma(ct' + \beta x')$$

$$\gamma = \frac{1}{\sqrt{1 - \beta^2}}$$

$$ct' = \gamma(ct - \beta x)$$

# The low-speed limit of the Lorentz transformation

$$\begin{array}{ll}
 & t \neq t' \\
 x = \gamma(x' + vt') & ct = \gamma(ct' + \beta x') \\
 \beta = \frac{v}{c} & \gamma = \frac{1}{\sqrt{1 - \beta^2}} \\
 x' = \gamma(x - vt) & ct' = \gamma(ct - \beta x)
 \end{array}$$

but for  $v \ll c$        $\beta \rightarrow 0$ ,  $\gamma \rightarrow 1$ , and

$$x = x' + vt' \quad ct = ct'$$

which is exactly the Galilean transformation used in classical mechanics.

# Conservation of momentum at low speed

At  $v \ll c$  momentum is defined as:

$$\vec{p} = m\vec{v}$$

The second Newton's law can be expressed as

$$\vec{F} = \frac{d\vec{p}}{dt}$$

In the absence of external force  $\vec{F} = 0$

$$\frac{d\vec{p}}{dt} = 0 \quad \Longrightarrow \quad \vec{p} = \text{const.}$$

and momentum is conserved.

# Conservation of momentum at high speed

To maintain the same form of Newton's law and the conservation of momentum up to  $v \sim c$  momentum has to be defined as

$$\vec{p} = \gamma m \vec{v}.$$

At low speed  $v \ll c$ ,  $\beta \sim 0$ ,  $\gamma \sim 1$  momentum becomes

$$\vec{p} = m \vec{v}.$$

# Conservation of energy at low speed

At  $v \ll c$  kinetic energy is defined as:

$$T = \frac{1}{2}mv^2 = \frac{p^2}{2m}$$

In the absence of external force momentum is conserved  $\vec{p}=\text{const.}$  which implies  $p^2=\text{const.}$  and  $T=\text{const.}$ , which means that the kinetic energy is conserved.



# Mass-energy theorem

In 1905 Albert Einstein following his derivation of the Special Theory of Relativity identifies relation between mass and energy of an object at rest:

$$E = mc^2.$$

The corresponding relation for moving object is

$$E = \gamma mc^2.$$

Subsequent experiments confirm conversion between mass and energy in atomic and nuclear processes.

## Energy, momentum and mass at high speed

At high speed the relation between energy, mass and momentum can be derived from

$$\begin{aligned}m^2 c^4 &= \frac{E^2}{\gamma^2} = E^2(1 - \beta^2) = E^2 - \frac{v^2}{c^2} \gamma^2 m^2 c^4 = \\ &= E^2 - \gamma^2 m^2 v^2 c^2 = E^2 - p^2 c^2\end{aligned}$$

or

$$E^2 = p^2 c^2 + m^2 c^4$$

Above equation is known as the relativistic energy-momentum relationship. It defines total energy of an object with mass  $m$  moving with momentum  $p$  in the absence of external force.

## Kinetic energy

Kinetic energy is a difference between energy of an moving object and the energy of the same object at rest.

$$T = E - mc^2 = \gamma mc^2 - mc^2 = (\gamma - 1)mc^2$$

further evaluation of this equation requires the following approximation which works well at low speed  $v \ll c$ ,  $\beta \ll 1$ :

$$\gamma \approx 1 + \frac{1}{2}\beta^2$$

which implies for the kinetic energy at low speed:

$$T = (\gamma - 1)mc^2 \approx (1 + \frac{1}{2}\beta^2 - 1)mc^2 = \frac{1}{2}m\beta c^2 = \frac{1}{2}mv^2.$$