Interaction with matter

accelerated motion:

$$s = \frac{b}{2}t^2$$
 \longrightarrow $t = \sqrt{\frac{2 \cdot s}{b}}$

$$v = v_0 - b \cdot t = v_0 - \sqrt{2 \cdot s \cdot b}$$



$$E = \frac{1}{2}mv^{2} \implies \frac{dE}{ds} = m \cdot \left(\frac{v_{0}}{\sqrt{2 \cdot s \cdot b}} - 1\right) \cdot b$$

$$u = 100^{km} - 27.78^{m} \qquad b = 3.86^{\frac{m}{2}}$$

$$v_0 = 100 \frac{km}{h} = 27.78 \frac{m}{s}$$
 $b = 3.86 \frac{m}{s^2}$



Interaction with matter











Some Nuclear Units

Nuclear energies are very high compared to atomic processes, and need larger units. The most commonly used unit is the MeV.

1 electron Volt = 1 eV = $1.6 \cdot 10^{-19}$ Joules

 $1 \text{ MeV} = 10^6 \text{ eV}; 1 \text{ GeV} = 10^9 \text{ eV}; 1 \text{ TeV} = 10^{12} \text{ eV}$

However, the nuclear size are quite small and need smaller units:

Atomic sizes are on the order of $0.1 \text{ nm} = 1 \text{ Angstrom} = 10^{-10} \text{ m}$. Nuclear sizes are on the order of femtometers which in the nuclear context are usually called fermis:

1 fermi = 1 fm = 10^{-15} m

Atomic masses are measured in terms of atomic mass units with the carbon-12 atom defined as having a mass of exactly 12 amu. It is also common practice to quote the rest mass energy $E=m_0c^2$ as if it were the mass. The conversion to amu is:

 $1 \text{ u} = 1.66054 \cdot 10^{-27} \text{ kg} = 931.494 \text{ MeV/c}^2$

electron mass = 0.511 MeV/c^2 ; proton mass = 938.27 MeV/c^2 ; neutron mass = 939.56 MeV/c^2



Mass data: nucleardata.nuclear.lu.se/database/masses/





Relevant Formulae

The relevant formulae are calculated if A_1 , Z_1 and A_2 , Z_2 are the mass number (amu) and charge number of the projectile and target nucleus, respectively, and T_{lab} is the laboratory energy (MeV)

$$E = T_{lab} + m_0 \cdot c^2$$
$$m \cdot c^2 = T_{lab} + m_0 \cdot c^2$$
$$\frac{m_0 \cdot c^2}{\sqrt{1 - \beta^2}} = T_{lab} + m_0 \cdot c^2$$

beam velocity:

 $\beta = \frac{\sqrt{T_{lab}}^2 + 1863 \cdot A_1 \cdot T_{lab}}{931.5 \cdot A_1 + T_{lab}}$

Lorentz contraction factor:

$$\gamma = (1 - \beta^2)^{-1/2}$$

$$\gamma = \frac{931.5 \cdot A_1 + T_{lab}}{931.5 \cdot A_1}$$

$$\beta \cdot \gamma = \frac{\sqrt{T_{lab}}^2 + 1863 \cdot A_1 \cdot T_{lab}}{931.5 \cdot A_1}$$





Energetic charged particles in matter

$$E_{kin} \cong \frac{1}{2}mv^2 \qquad -\frac{dE}{dx} \propto \frac{mz^2}{E}$$



dominant in the classical limit [40 MeV/A (0.3c) - <1% deviation

$$-\frac{dE}{dx} = \frac{4\pi e^4 z^2}{m_0 v^2} nZ \left[\ln \frac{2m_0 v^2}{I} - \ln \left(1 - \frac{v^2}{c^2} \right) - \frac{v^2}{c^2} \right]$$

Bethe-Block formula

z – projectile atomic number v – projectile velocity m_0 - electron mass e – electron charge

- n target number density
- Z target atomic number
- nZ target electron density
- I average excitation and ionization potential





Energetic charged particles in matter



z – projectile atomic number v – projectile velocity m₀ - electron mass e – electron charge

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Energetic charged particles in matter



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 α -particles are highly ionized and lose their energy very fast by ionization and excitation when passing through matter.



maximum energy transfer T_{max} of a projectile with mass m and velocity β on an electron m_e at rest

$$T_{max} = \frac{2 \cdot m_e c^2 \cdot \beta^2 \cdot \gamma^2 \cdot m^2}{m^2 + m_e^2 + 2 \cdot \gamma \cdot m \cdot m_e}$$

$$T_{max} = 2 \cdot m_e c^2 \cdot \beta^2 \cdot \gamma^2$$

for all heavy primary particles except electrons and positrons

average range $\langle R \rangle$ of α -particles with 5 MeV 2,5cm in air, 2,3cm in Al, 4,3cm in tissue







Interaction of charged particles in matter

Bethe-Bloch formula describes the energy loss of heavy particles passing through matter

$$-\frac{dE}{dx} = \underbrace{4 \cdot \pi \cdot r_e^2 \cdot N_a \cdot m_e c^2}_{\gamma} \cdot \rho \cdot \frac{Z}{A} \cdot \frac{z^2}{\beta^2} \cdot \left[\frac{1}{2} ln \left(\frac{2 \cdot m_e c^2 \cdot \gamma^2 \cdot \beta^2 \cdot T_{max}}{l^2} \right) - \beta^2 - \delta - 2 \cdot \frac{C}{Z} \right] \approx z^2 \cdot \frac{Z}{A} \cdot f(\beta, l)$$

$$= 0.3071 \text{ MeV g}^{-1} \text{cm}^2$$
N_a: Avogadro number 6.02 · 10²³ mol⁻¹
r_e: class. electron radius 2.81 · 10⁻¹³ cm
m_a: electron mass
 ρ : density of abs. matter
Z: element number of abs. matter
A: mass of abs. matter
Z: charge number of incoming particle
W_{max}: max. energy transfer in a single collision
I: average ionization potential
$$F_{\alpha} = \underbrace{4 \cdot \pi \cdot r_e^2 \cdot N_a \cdot m_e c^2}_{12} \cdot \left[\frac{1}{2} ln \left(\frac{2 \cdot m_e c^2 \cdot \gamma^2 \cdot \beta^2 \cdot T_{max}}{l^2} \right) - \beta^2 - \delta - 2 \cdot \frac{C}{Z} \right] \approx z^2 \cdot \frac{Z}{A} \cdot f(\beta, l)$$



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dE/dx has a minimum at βγ~3-4 (minimum ionizing particle)
 for high momenta dE/dx reaches a saturation

for small β the term $1/\beta^2$ is dominant

1.0

10

 $\beta \gamma = p/Mc$

100

1000

0.5 [[]

 0.1



10000

Interaction of charged particles in matter

Bethe-Bloch formula describes the energy loss of heavy particles passing through matter

$$-\frac{dE}{dx} = 4 \cdot \pi \cdot r_e^2 \cdot N_a \cdot m_e c^2 \cdot \rho \cdot \frac{Z}{A} \cdot \frac{z^2}{\beta^2} \cdot \left[\frac{1}{2} ln \left(\frac{2 \cdot m_e c^2 \cdot \gamma^2 \cdot \beta^2 \cdot T_{max}}{l^2}\right) - \beta^2 - \delta - 2 \cdot \frac{C}{Z}\right] \approx z^2 \cdot \frac{Z}{A} \cdot f(\beta, l)$$

$$= 0.3071 \text{ MeV g}^4 \text{ cm}^2$$

$$\Rightarrow \text{ energy loss of a particle is independent of its mass!}$$

$$\Rightarrow \text{ energy loss is an important tool for particle identification}$$

$$\Rightarrow \text{ for minimum ionizing particles m.i.p.} \frac{10}{dE/dx \sim 2 \text{ MeV g}^4 \text{ cm}^2}$$

$$i.e. \text{ for a target density } \rho = 1 \text{ g/cm}^3$$

$$dE/dx \sim 2 \text{ MeV/cm}$$

$$= 0.3071 \text{ MeV g}^4 \text{ cm}^2$$

$$i.e. \text{ for a target density } \rho = 1 \text{ g/cm}^3$$

$$dE/dx \sim 2 \text{ MeV/cm}$$

 \triangleright for small β the term $1/\beta^2$ is dominant

> dE/dx has a minimum at $\beta\gamma$ ~3-4 (minimum ionizing particle)

 \blacktriangleright for high momenta dE/dx reaches a saturation



Energy loss and range of charged particles

$$-\frac{dE}{d\varepsilon} = -\frac{1}{\rho} \cdot \frac{dE}{dx} = z^2 \cdot \frac{Z}{A} \cdot f(\beta, I)$$

 $-dE/d\epsilon$ is independent of the material for equal particles

- the average range for particles with kin. energy T is obtained by integration

$$\bar{R} = \int_{E_0}^0 \left(\frac{dE}{dx}\right)^{-1} dE$$

- 7.7 MeV α -particles in air: $\overline{R}/\rho \approx 7cm$
- range is not exact but there is range straggling, the number of interactions is a statistical process.







Interaction of β -particles with matter

 β -particles are also ionizing, similar to α -particles. Since the mass of the electrons and positrons are very small, the energy transfer per collision is small and the range large. Similar to the X-rays there is first only an attenuation, which finally leads to a maximum range for larger layer thicknesses.





Interaction of β -particles with matter

 β^+ particles behave similarly as β^- particles; they are ionizing and attenuated on their way through matter.

But at the end of the their attenuation one observes a pair annihilation with an electron, which yields high energetic γ -emission. Positrons are hence more dangerous than electrons.







Energy loss for electrons and positrons

 e^{\pm} are exceptional cases due to their low mass. They will be deflected significantly in each collision.

In addition to the energy loss due to **ionization**, the energy loss due to **Bremsstrahlung** is of importance.

$$-\left(\frac{dE}{dx}\right)_{tot} = -\left(\frac{dE}{dx}\right)_{coll} - \left(\frac{dE}{dx}\right)_{rad}$$

For high energies the energy loss due to Bremsstrahlung is given by

$$-\left(\frac{dE}{dx}\right)_{rad} \propto E$$
 and $-\left(\frac{dE}{dx}\right)_{rad} \propto \frac{1}{m^2}$

Other particles like muons also radiate, especially at higher energies.



Bremsstrahlung

Hans-Jürgen Wollersheim - 2017

Bremsstrahlung ('braking radiation') is electromagnetic radiation produced by the deceleration of a charged particle when deflected by another charged particle, typically an electron by an atomic nucleus.

The moving particle loses kinetic energy, which is converted into a photon, thus satisfying the law of conservation of energy.

Bremsstrahlung has a continuous spectrum.



GSÍ



Bremsstrahlung



K = Kramer constant

I = electron current

Z = element number of material

Spectrum of the X-rays emitted by an X-ray tube with a rhodium target, operated at 60 kV. The continuous curve is due to bremsstrahlung, and the spikes are characteristic K lines for rhodium.



Bremsstrahlung

$$-\left(\frac{dE}{dx}\right)_{rad} = \frac{E}{X_0}$$

 X_0 is the radiation length. It is the mean distance over which a high-energy electron loses all but 1/e of its energy by Bremsstrahlung

fit to data:

$$X_0 = \frac{716.4 \cdot A}{Z \cdot (Z+1) \cdot \ln(287/\sqrt{Z})}$$

Usual definition for the critical energy \mathbf{E}_{c} (electron)

$$\left(\frac{dE}{dx}\right)_{ionization} = \left(\frac{dE}{dx}\right)_{bremsstrahlung}$$

$$E_{c}(e^{-}) = \begin{cases} \frac{610 \ MeV}{Z + 1.24} & \text{for solids and liquids} \\ \frac{710 \ MeV}{Z + 0.92} & \text{for gases} \end{cases}$$



example: Pb (Z=82,
$$\rho = 11.34 \text{ [g/cm^3]} \rightarrow \text{E}_c = 7.34 \text{ MeV}$$





Synchrotron radiation

The electromagnetic radiation emitted when charged particles are accelerated radially $(\mathbf{a} \perp \mathbf{v})$ is called **synchrotron radiation.** It is produced, for example, in synchrotrons using bending magnets.

The energy loss of a charged particle $(Z \cdot e)$ due to radiation (during one cycle) is given by

$$\Delta E = \frac{(Ze)^2 \cdot \beta^3 \cdot \gamma^4}{\epsilon_0 \cdot 3R}$$

with Z = element number, ϵ_0 = electric field constant, R = radius of the storage ring, β =v/c and the Lorentz factor

$$\gamma = \frac{1}{\sqrt{1 - \beta^2}} \equiv \frac{E}{m_0 c^2}$$

For relativistic velocities $\beta \approx 1$

$$\Delta E = \frac{(Ze)^2 \cdot E^4}{\epsilon_0 \cdot 3R \cdot (m_0 c^2)^4}$$

It is apparent that one uses light particles to create synchrotron radiation.









Synchrotron radiation



 (\mathfrak{S})



Raja Ramanna Centre for Advanced Technology - Synchrotron radiation







 Applications: condensed matter physics, material science, biology and medicine.



Structure of a ribosome (components of a cell)





Typical range of radioactive radiation in air

range of 5.5 MeV α -particles in air is ~ 4.2 cm

range of 1 MeV β -particles in air is ~ 4 m



range of X-rays, γ -rays and neutrons is very large. shielding or large distances (1/r² law) are the solution



