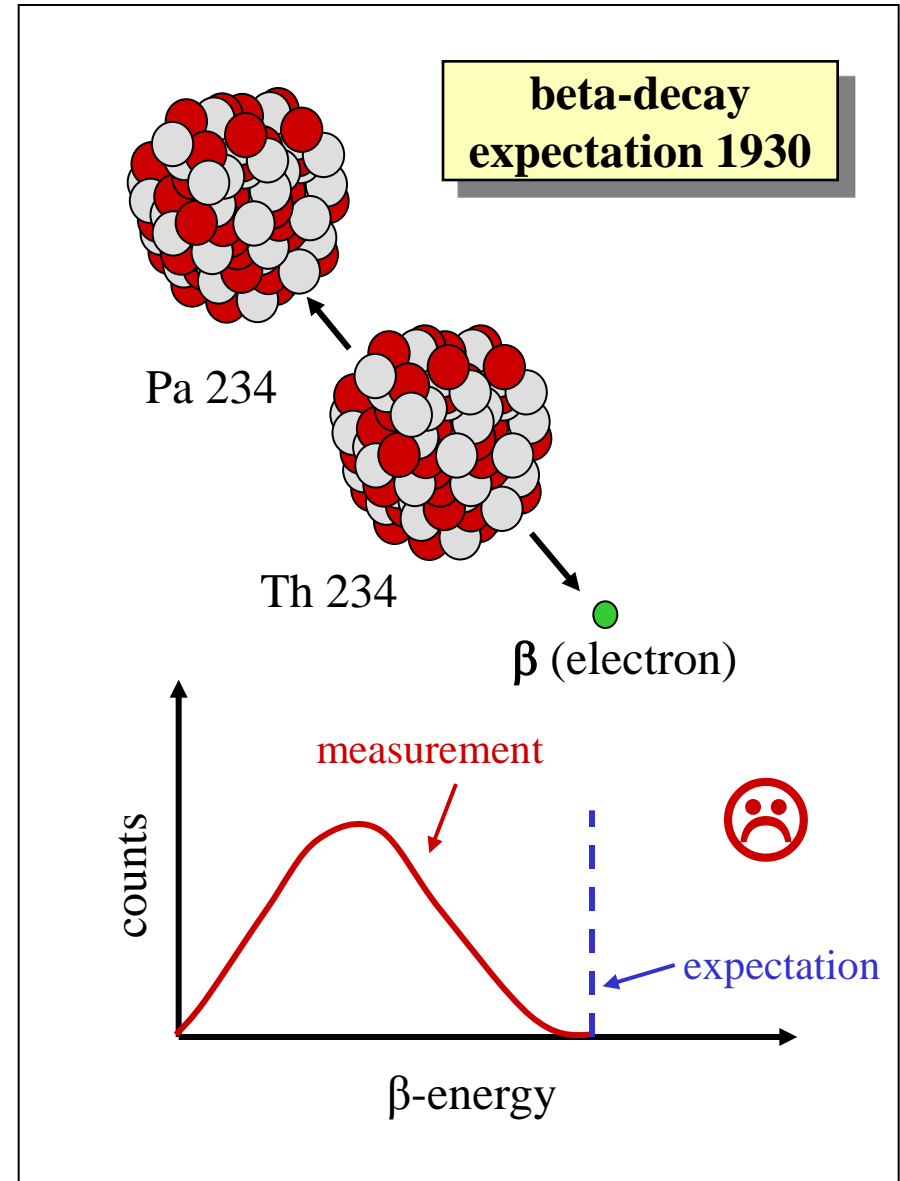
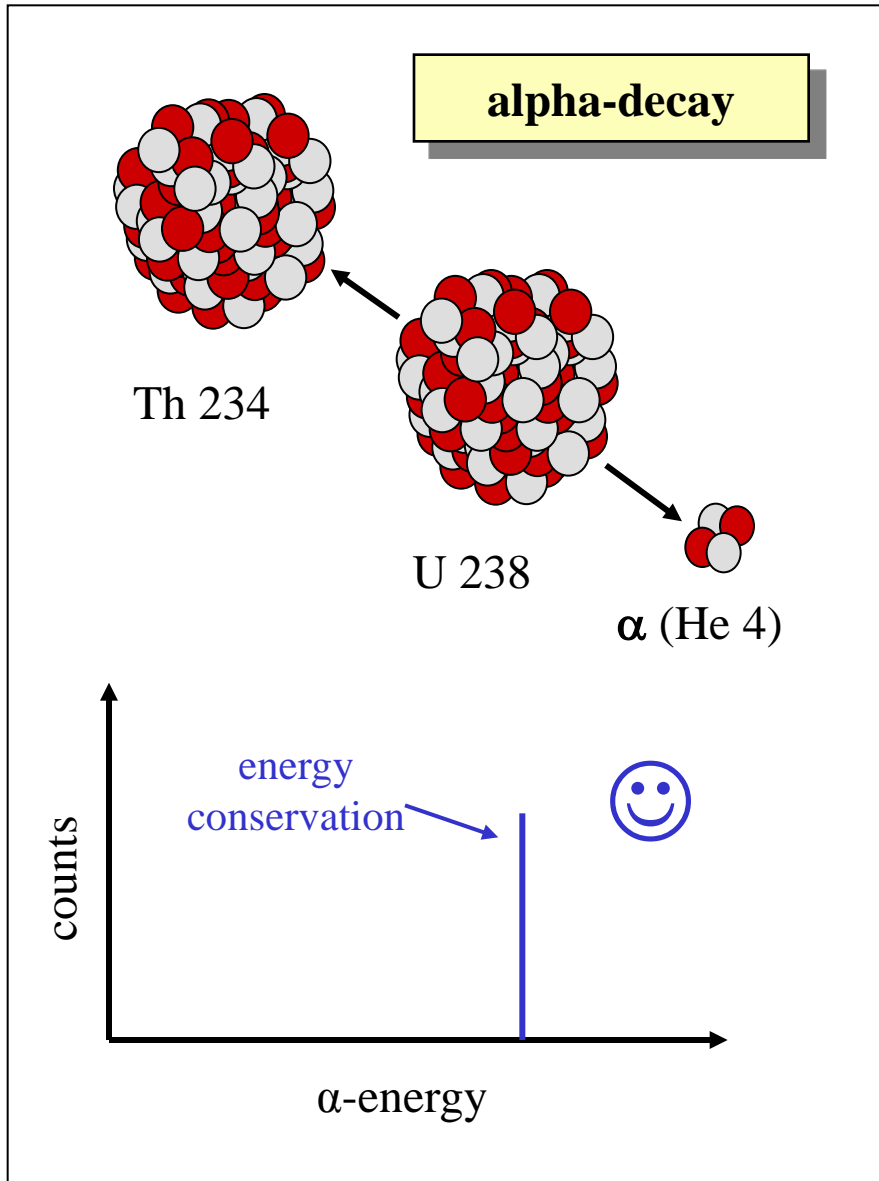


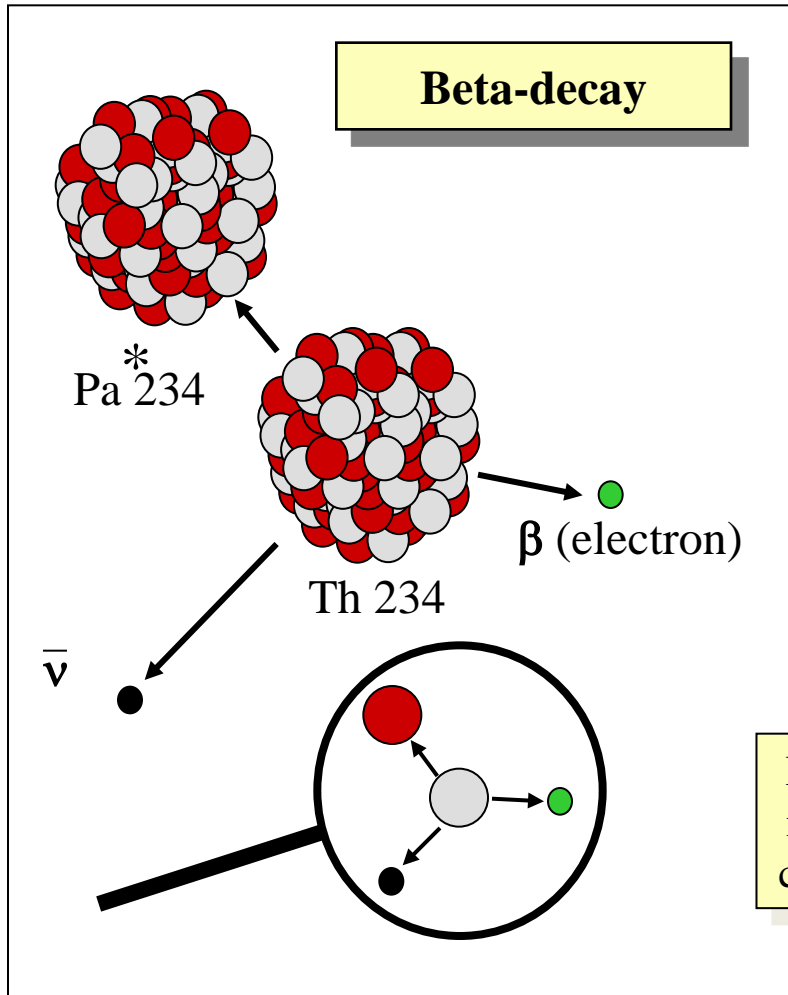
1930: Energy conservation violated in β -decay





Wolfgang Pauli

The new particle proposed by Pauli



| Leptonen | |
|--------------|-----------------------------|
| Charge -1 | Charge 0 |
| Electron e | e-Neutrino ν_e |
| Myon μ | μ -Neutrino ν_μ |
| Tauon τ | τ -Neutrino ν_τ |

In addition to an electron a light neutral particle is created which carries away the missing energy!

„Today I have done something, what one should not do in theoretical physics. I have explained something, what is not understood, by something, which can not be observed!“

Neutrino detection

- Detection of particles:

Interaction of particles with matter (detector)

- Interaction with matter depends strongly on the particle:

Charge particles: Ionization of matter

Photons: Energy transfer to charged particles

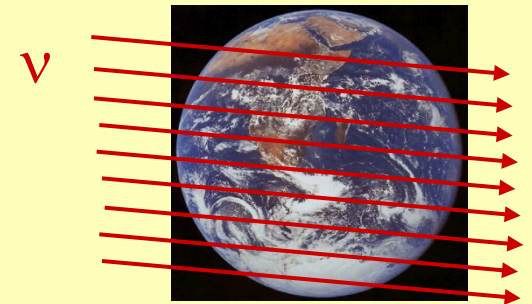
Neutrons: Nuclear reactions yield charge particles

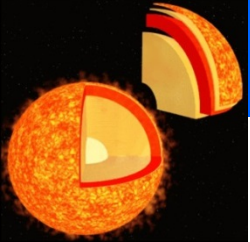


- Neutrinos interact very weakly:

Only **one out of 100 billions neutrinos** from the β -decay will be recognized by the earth.

Calculated 1934: „Hopeless“



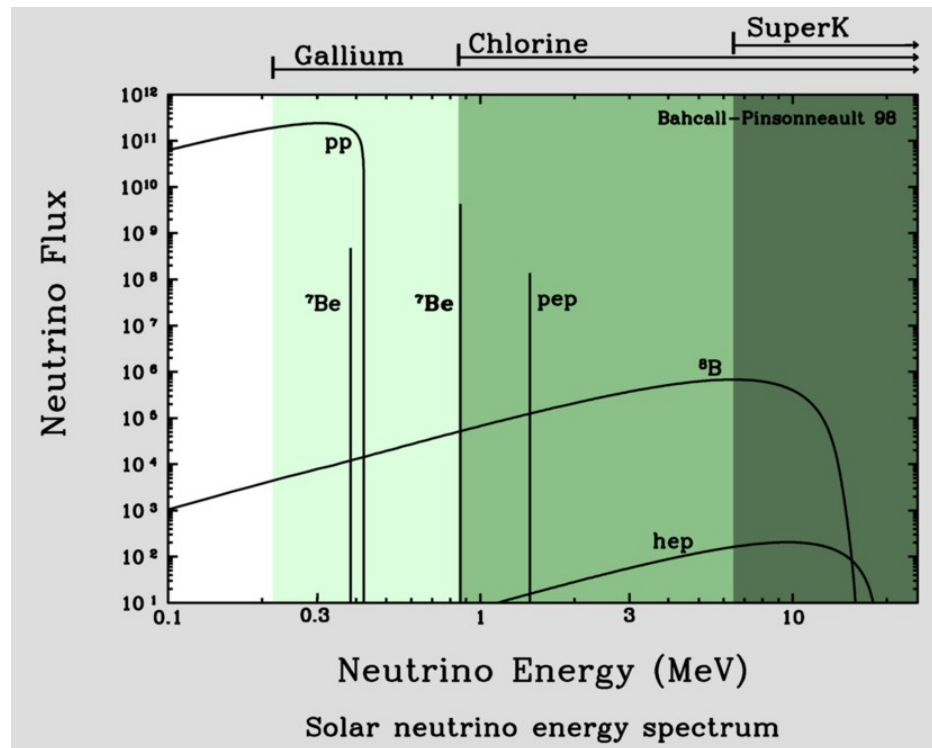
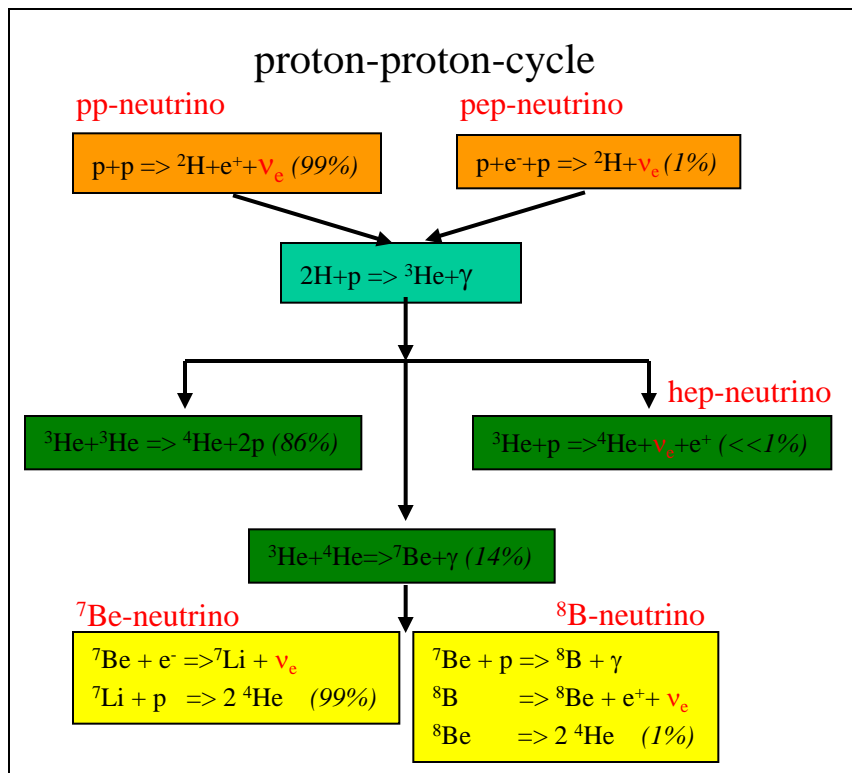


Neutrinos from the sun



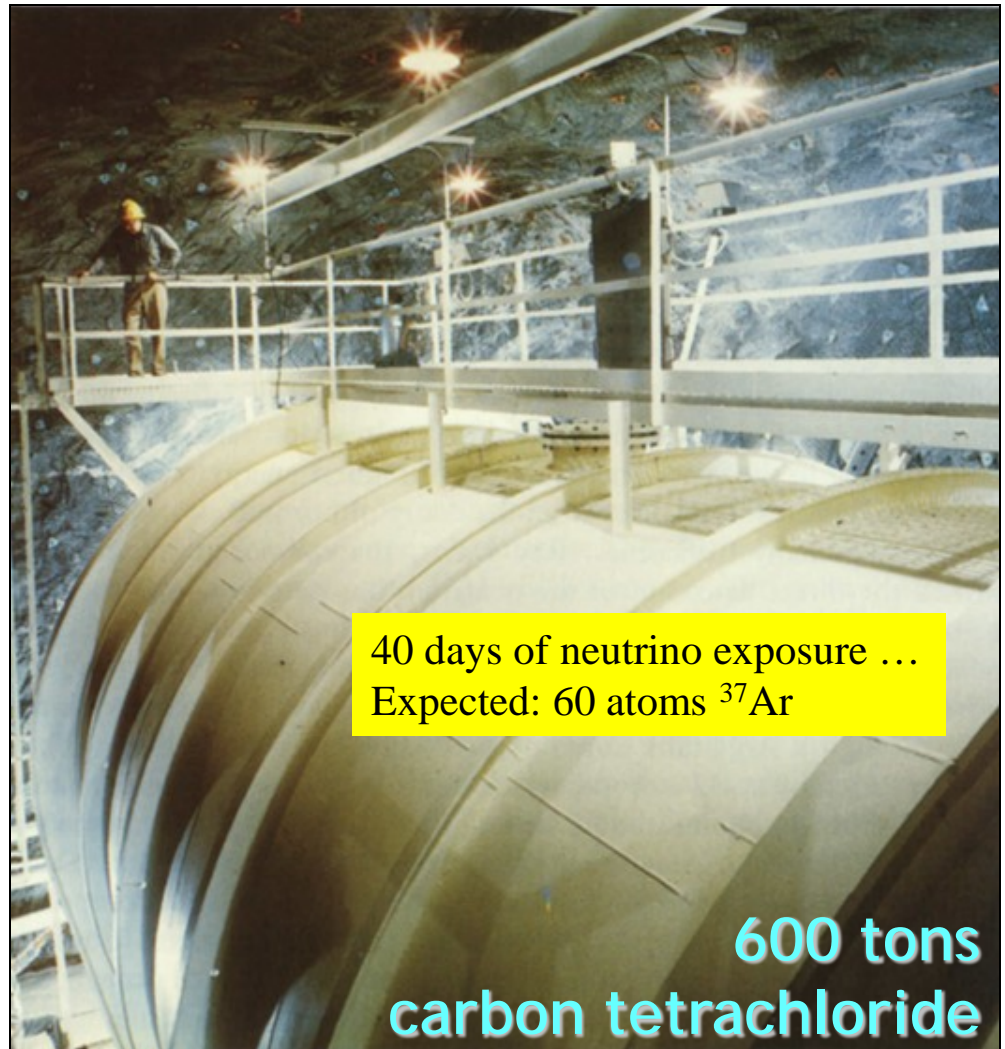
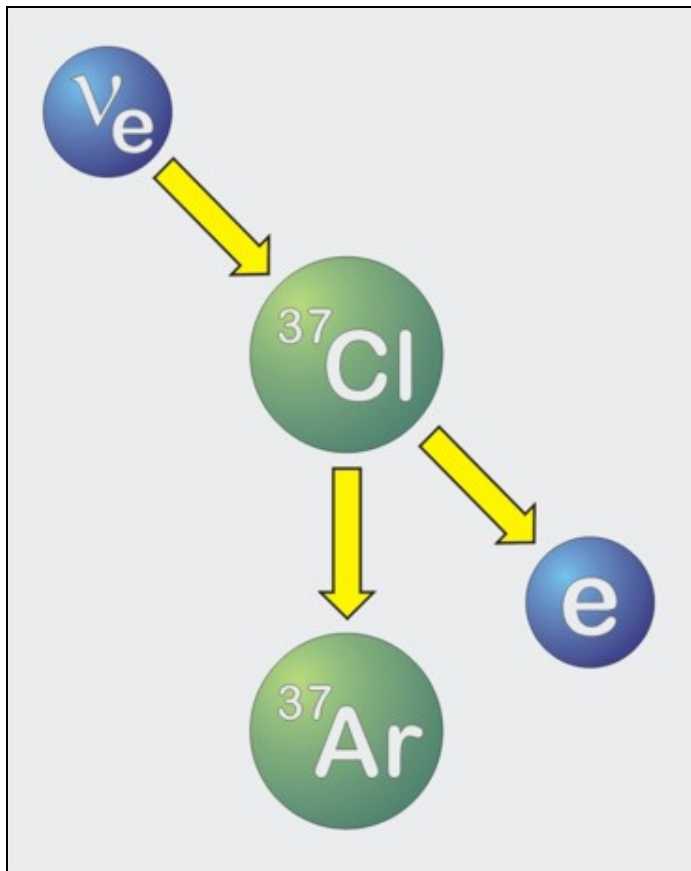
- Known: total emitted energy
- Known: energy per fusion process

➤ number of created neutrinos per sec!
 on earth: 66 billions ν per $(\text{cm}^2 \cdot \text{s}^1)$



First measurements of the solar neutrinos

Inverse beta-decay
(„neutrino-capture“)



Homestake Solar Neutrino
Observatory (1967–2002)

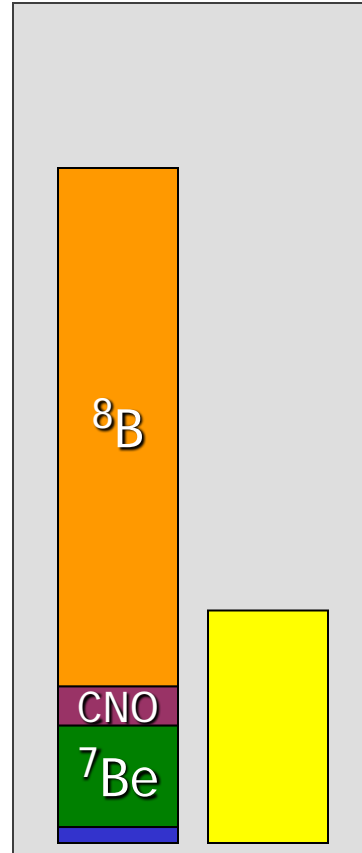
Problem of the “missing” solar neutrinos



John Bahcall
1934 – 2005

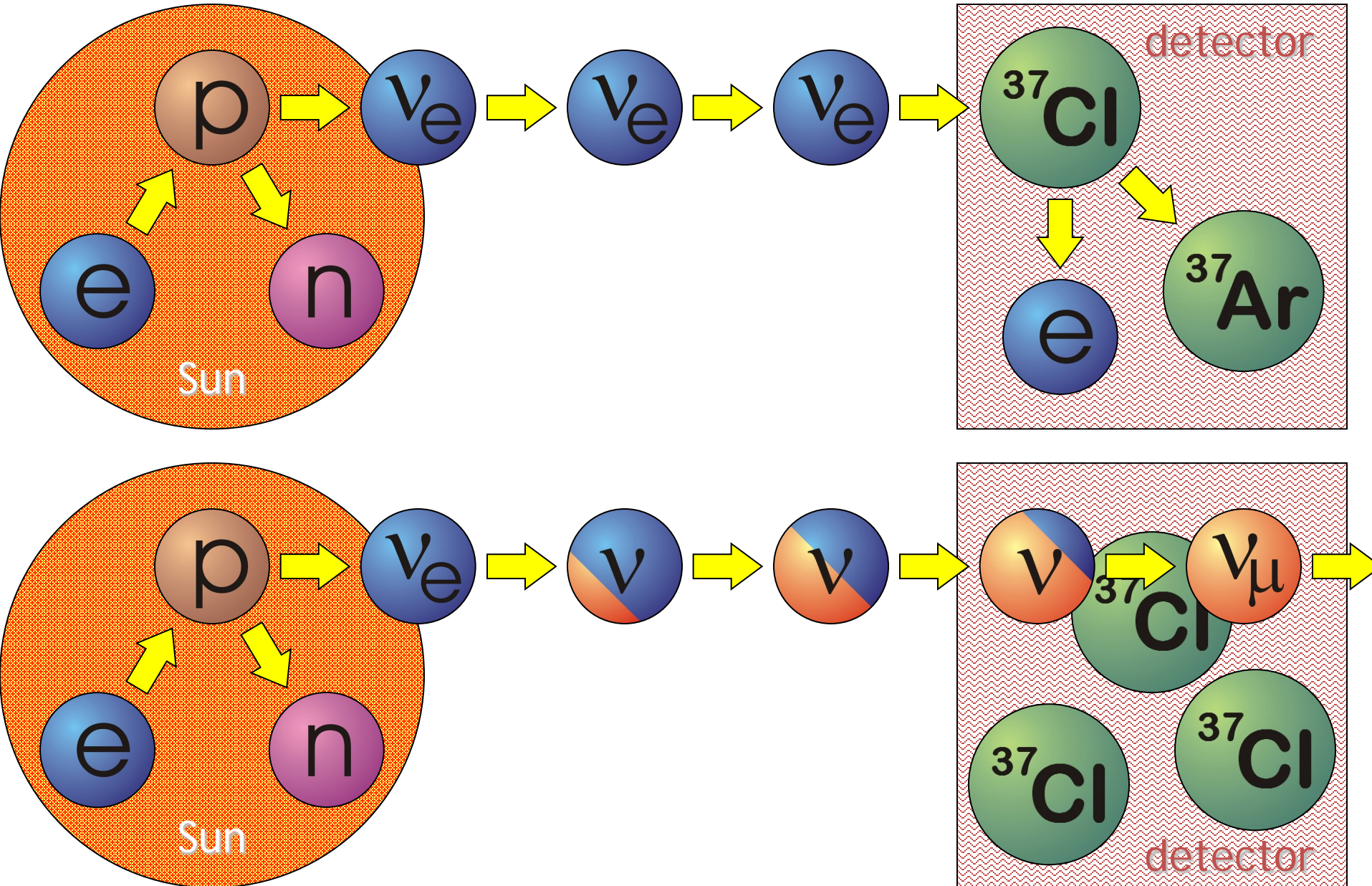
Homestake

Chlorine



Raymond Davis Jr.
1914 – 2006

Neutrino conversion is the solution of the problem



Neutrino oscillations

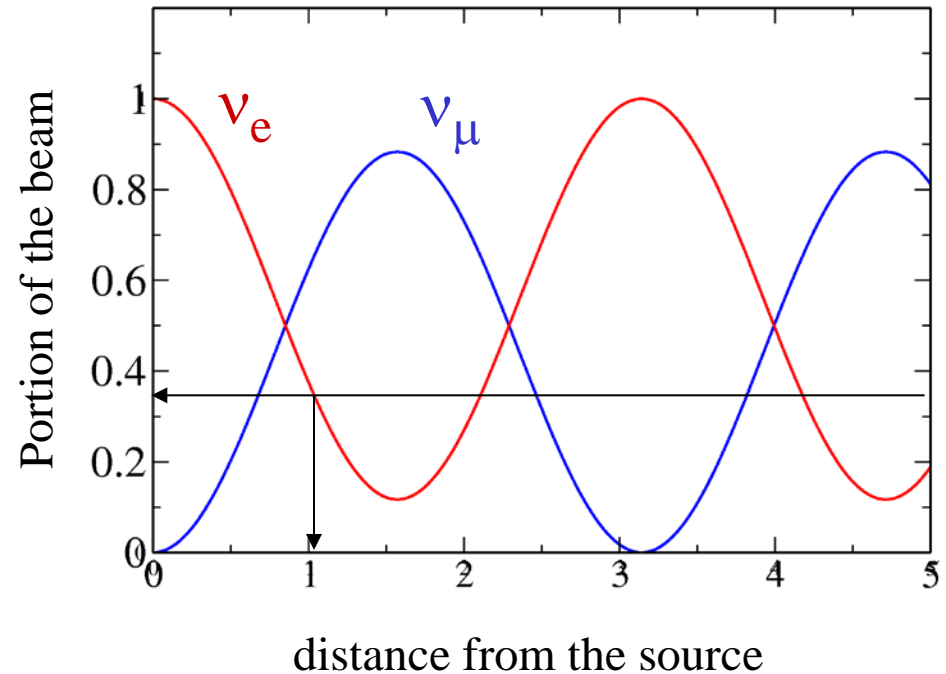
| | | |
|------------|-----------------|------------------|
| Electron e | Myon μ | Tau τ |
| e-Neutrino | μ -Neutrino | τ -Neutrino |

Idea: when neutrinos have a mass, they may convert into each other!

Assumption: Mixture of

ν_e and ν_μ

Conversion of a neutrino beam with the distance from the neutrino source:



1998: Confirmation of the oscillations between Myon- and Tau-neutrinos with methods of Super-Kamiokande (Myon-neutrinos from the atmosphere)

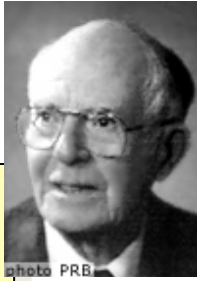
The solar neutrino problem



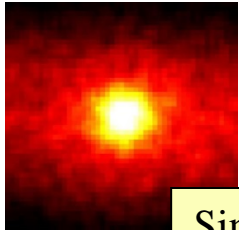
Sun
since 4.5 billion years
fusion

66 billion neutrinos/s/cm²

since 1964: detection with
Homestake-experiment
expected: 1,5 reactions/d
measured: 0.5 reactions/d



solar
neutrino
problem



Since 1986 Kamiokande:
confirms Homestake

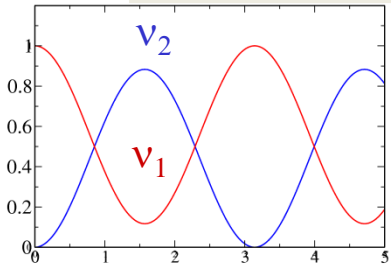
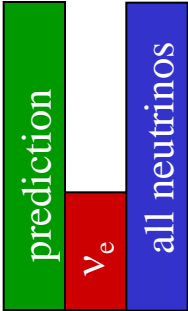


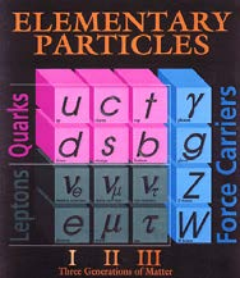
possible explanation:
neutrino-oscillation

2002 SNO-experiment:
examines neutrino-oscillation



solar
neutrino
problem
solved!





Neutrinos as astrophysical messenger



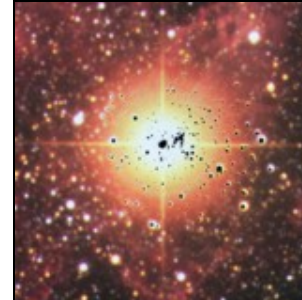
nuclear reactors



Sun



particle accelerator

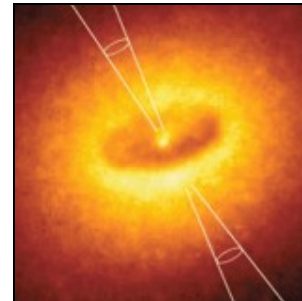
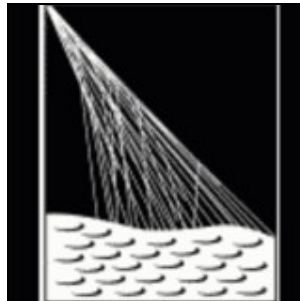


Supernovae
(collapsing stars)

SN 1987A ✓



earth atmosphere
(cosmic radiation)

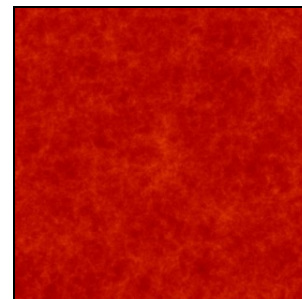


astrophysical
accelerators

soon ?



earth crust
(natural radioactivity)



Big Bang of the Universe
(today $330 \nu/\text{cm}^3$)
indirect evidence