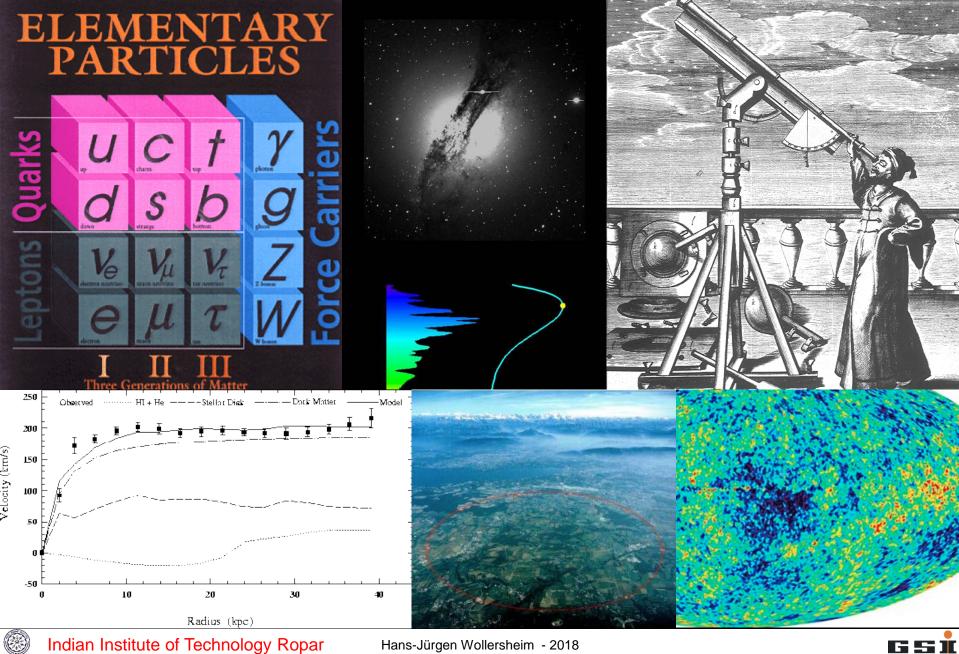
### Beyond the Standard Model



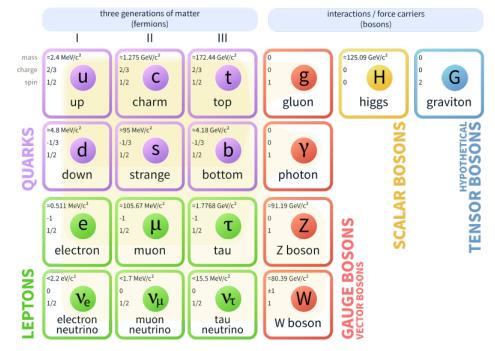
Hans-Jürgen Wollersheim - 2018

The Standard Model is an incomplete theory. It cannot explain why a particle has a certain mass.



Physicists have theorized the existence of the so-called Higgs field, which in theory interacts with other particles to give them mass. The Higgs field requires a particle, the Higgs boson.

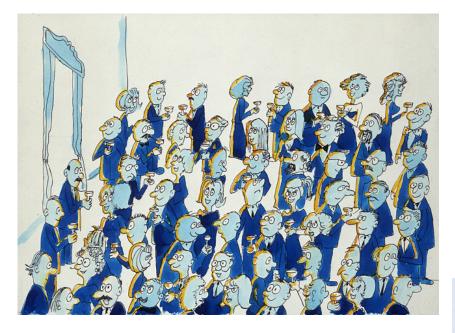
#### Standard Model of Elementary Particles + Gravity







### The Higgs Boson



A massless particle moving in a Higgs field is equivalent to a massive particle ⇒ Higgs field "gives mass to all particles"

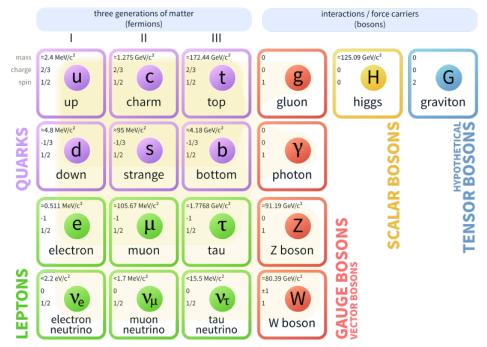




The Standard Model is an incomplete theory. It does not adequately explain:

- Gravity (graviton simply added)
- Dark matter (26%) and dark energy (69%)
- Matter antimatter asymmetry (should be equal)
- Mass of neutrinos (what was their role in the formation of the universe?)
- Superconductivity (how does high T<sub>c</sub> work?)
- Quantum theory of gravity (is there a QTG that can describe the universe we live in?
- Number of dimensions in a fundamental theory of nature.

### Standard Model of Elementary Particles + Gravity

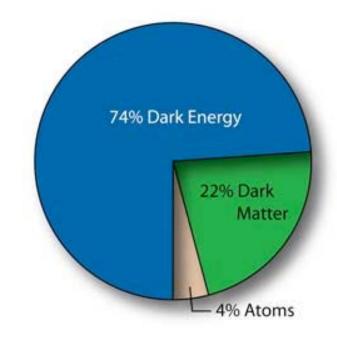






### What is the Universe really made of?

- Particle physicist's answer: stable particles – protons, neutrons, electrons, neutrinos
- (Why not antiprotons, positrons, etc.? another puzzle may be next time?)
- But astronomical observations indicate that the known particles make only about 4% of the stuff in the Universe!

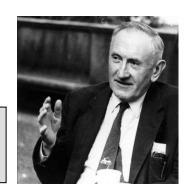








*Coma Cluster*: By measuring the velocities of all these galaxies Fritz Zwicky realized that galaxies toward the edge of the cluster were moving far too fast.

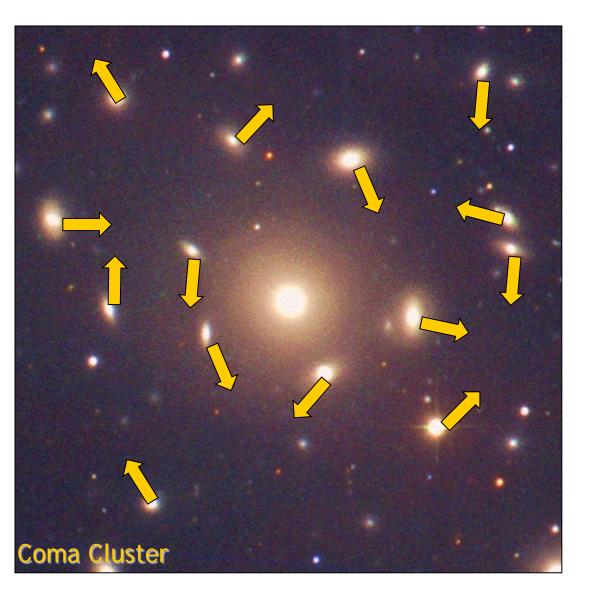


Fritz Zwicky: Die Rotverschiebung von Extragalaktischen Nebeln (The redshift of extragalactic nebulae) Helv. Phys. Acta 6 (1933) 110

In order to obtain, as observed, a medium-sized Doppler effect of 1000 km/s or more, the average density in the Coma system would have to be at least 400 times greater than that derived on the basis of observations of luminous matter. If this should be verified, it would lead to the surprising result that dark matter exists in much greater density than luminous matter.







A gravitational bound system of many 'particles' follows the Virial theorem

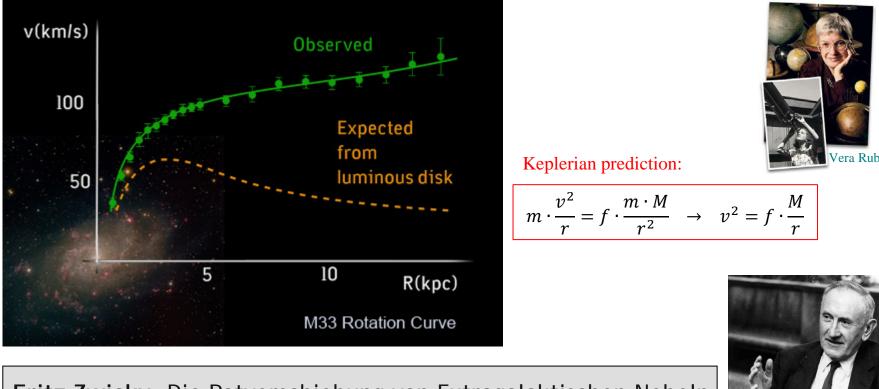
$$2\langle E_{kin} \rangle = -\langle E_{grav} \rangle$$
$$2\langle \frac{mv^2}{2} \rangle = \langle \frac{G_N M_r m}{r} \rangle$$
$$\langle v^2 \rangle \approx G_N M_r \langle r^{-1} \rangle$$

Velocity measurement via Doppler effect of at least three spectral lines

## Mass estimate







Fritz Zwicky: Die Rotverschiebung von Extragalaktischen Nebeln (The redshift of extragalactic nebulae) Helv. Phys. Acta 6 (1933) 110

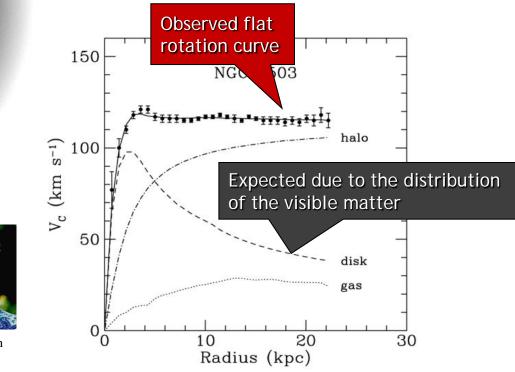
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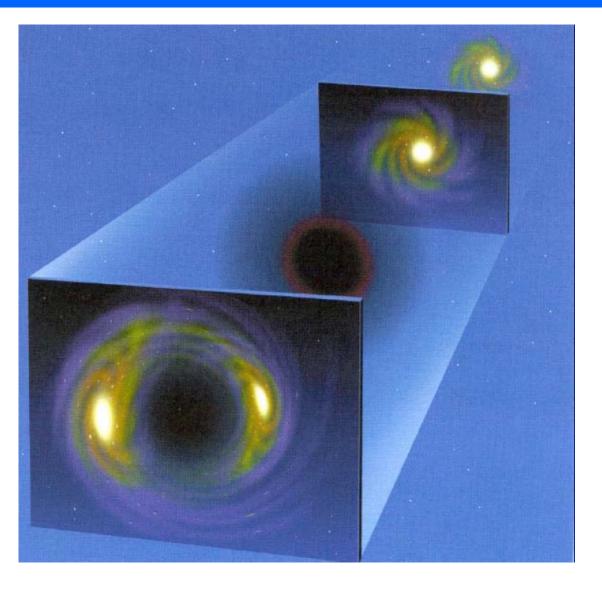
- Dark matter in the Halo around a Galaxy
- Dark matter ~ 10x visible mass







### Einstein ring

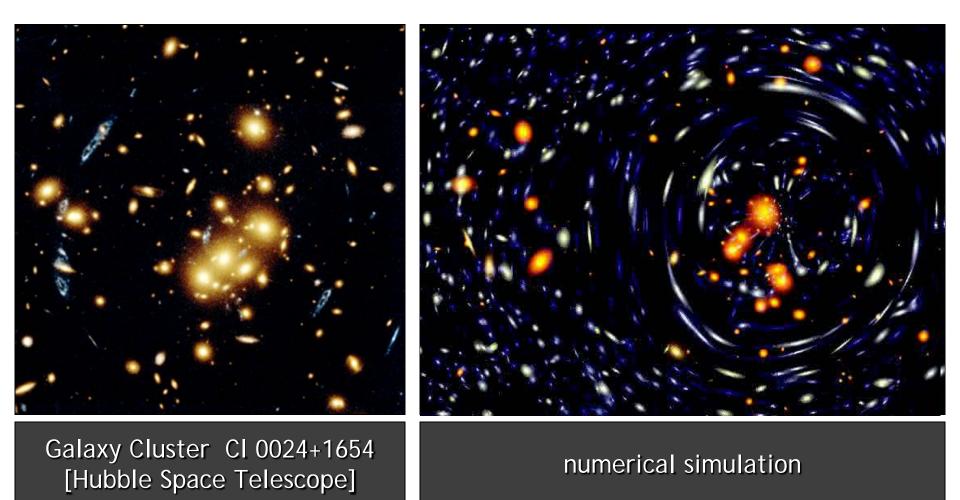


It is the deformation of light from a source (such as a galaxy or star) into a ring through gravitational lensing of the source's light by an object with an extremely large mass.





### Gravitational lensing effect





# Bullet Cluster (1E 0657-56)

Crash of galaxy clusters: Dark matter (blue) creating most of the gravitational potential separated from normal matter (pink).

## What is Dark Energy?

Big Bang Cosmology: Albert Einstein (1879-1955)

**Prediction**: The universe is expanding

**Observation**: Galaxies are moving apart from each other (1929)



Red-shift of the spectral lines in stars (Doppler effect)



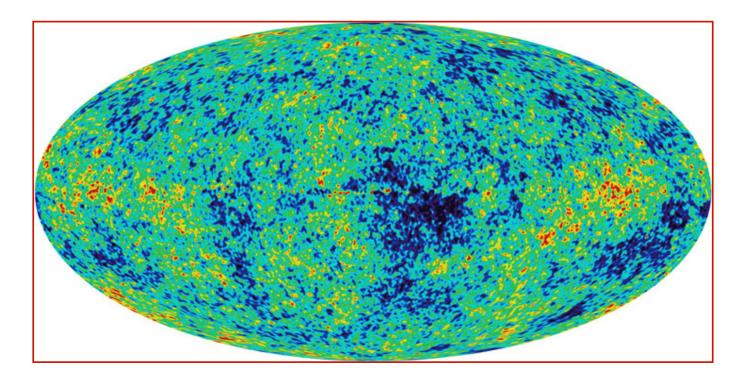


## What is Dark Energy?

### Testing the Big Bang model

**Prediction**: If the universe was denser, hotter, in past, we should see evidence of left-over heat from early universe

**Observation**: Left-over heat from the early universe (Penzias and Wilson, 1965)





## What is Dark Energy?

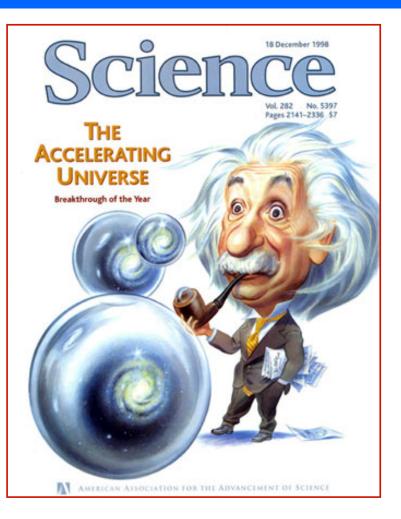
### Testing the Big Bang model

**Observation**: Expansion is accelerating.

Refine: Extra energy content.

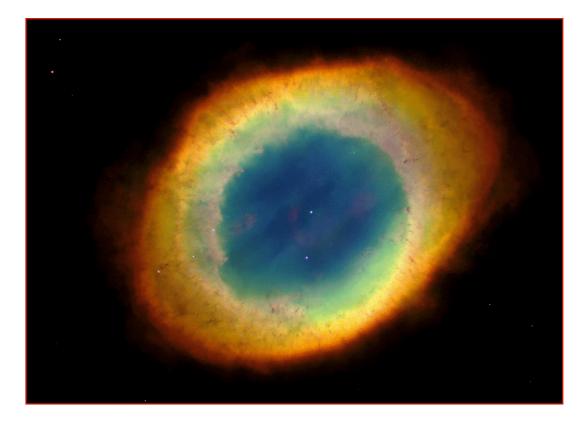
A recent discovery and of unknown origin, the concept of Dark Energy is actually an integral part of Einstein's theory of gravity.

theory of relativity lies nearest at hand; whether, from the standpoint of present astronomical knowledge, it is tenable, will not here be discussed. In order to arrive at this consistent view, we admittedly had to introduce an extension of the field equations of gravitation which is not justified by our actual knowledge of gravitation. It is to be emphasized, however, that a positive curvature of space is given by our results, even if the supplementary term is not introduced. That term is necessary only for the purpose of making possible a quasi-static distribution of matter, as required by the fact of the small velocities of the stars.





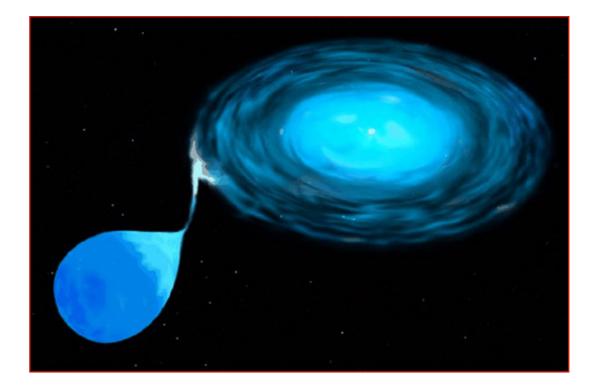




A dying star becomes a white dwarf.



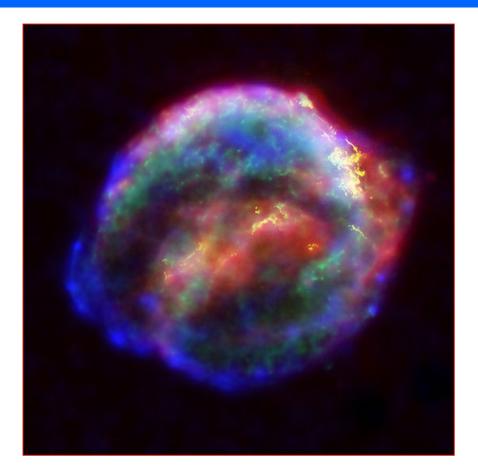




The white dwarf strips gas from its stellar companion ...







... and uses it to become a hydrogen bomb. Bang!







The explosion (Supernova Type Ia) is as bright as an entire galaxy of stars





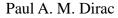
Why is there more matter than antimatter?

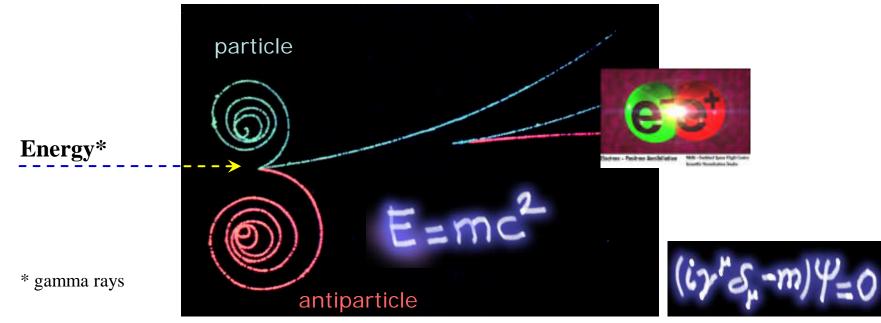
**Dirac: Key Discovery** 

Relativity + Quantum Theory

# ⇒ 'Antiparticles'







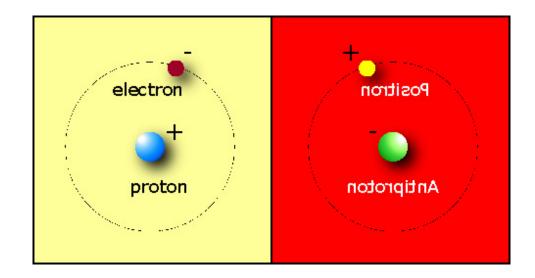




## Is the Universe symmetric?

### 1933 Dirac (from his Nobel lecture)

"If we accept the view of complete symmetry between positive and negative electric charge so far as concerns the fundamental laws of Nature, we must regard it rather as an accident that the Earth (and presumably the whole solar system), *contains a preponderance of negative electrons and positive protons*.







## The Antimatter Mystery



Without asymmetry - we would not be here!!

# What kind of asymmetry??



## Sakharov's Idea



A.D.Sakharov

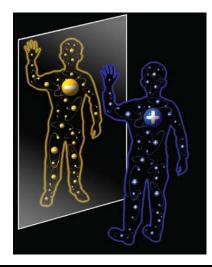
Particles decay (a little) faster than antiparticles\*

Small imbalance (1,000,000,001:1,000,000,000) Occurs during cool-down of Universe

Most particle-antiparticle pairs annihilate to radiation

Galaxies, stars, planets, us = 'left-over'

### \*For experts: this is called 'CP violation'

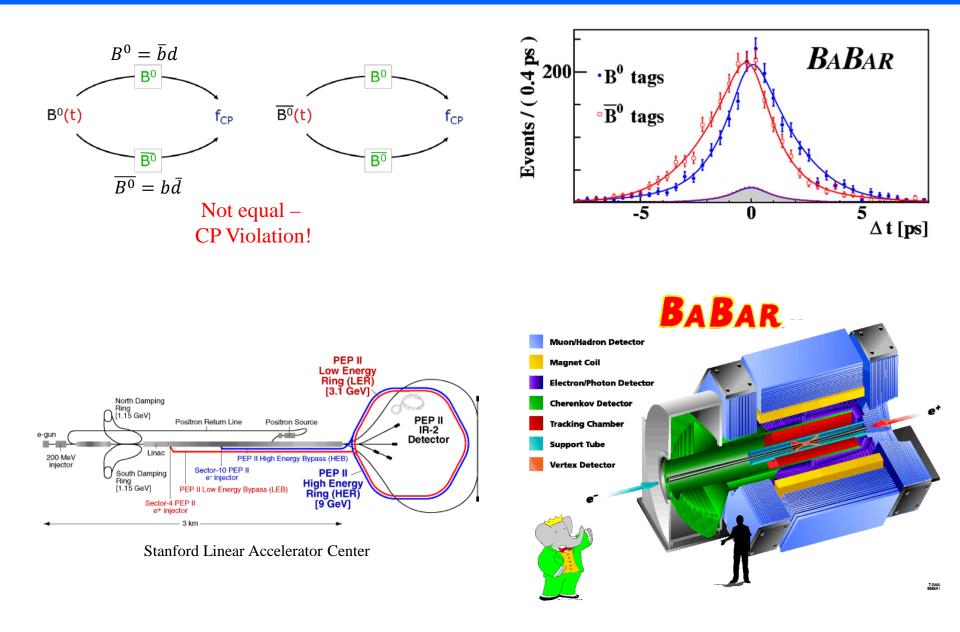


When we look at our image in a standard mirror, we are looking at the parity transformation of ourselves



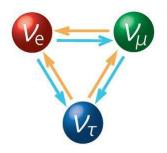


### Measuring CP Violation with B<sup>0</sup>s





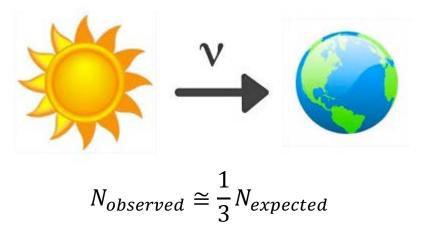
### Neutrino masses



The Standard Model was built with the assumption of massless neutrinos

• No right-handed neutrinos, and then no Dirac mass

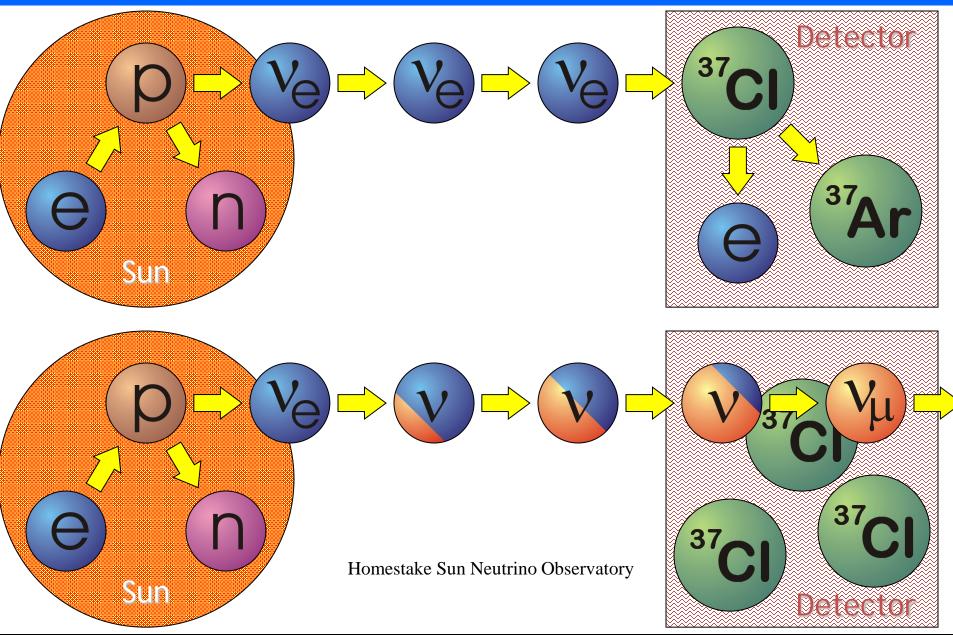
### Solar neutrino problem:







### Solar Neutrino problem

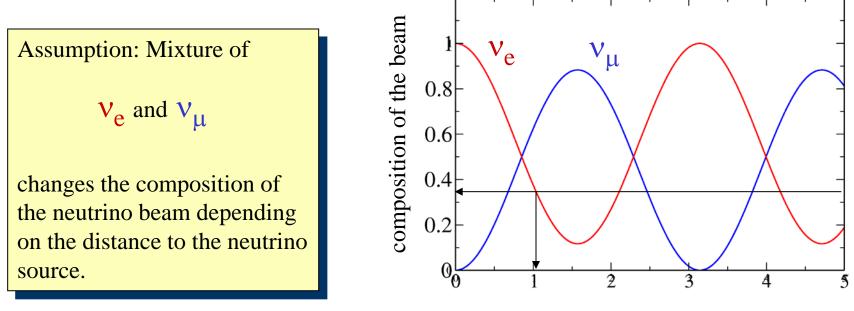




## **Neutrino Oscillations**

Electron e	Myon µ	Tau τ
e-Neutrino	µ-Neutrino	τ-Neutrino

Idea: if the neutrinos have a non-zero mass, they can interact with each other!



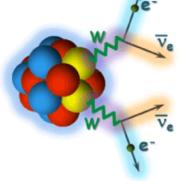
distance from the source

1998: Detection of oscillation between Myon- und Tau-neutrinos with methods of Super-Kamiokande (Myon-Neutrinos from the atmosphere)



### Neutrinoless Double Beta Decay

[Double beta decay]

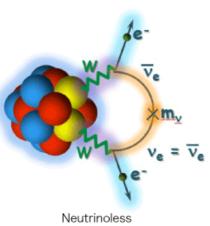


Double beta decay which emits anti-neutrinos



Paul Dirac

#### Fermions particle-antiparticle pair



double beta decay

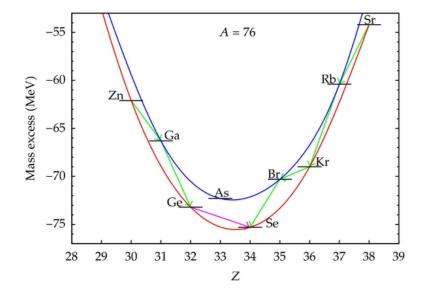


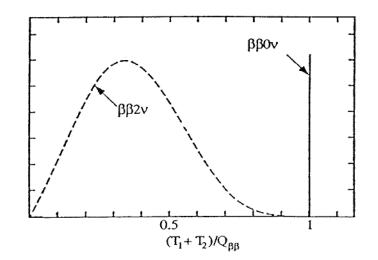
Ettore Majorana

### **Neutral Particles**

they are their own antiparticles works for massless neutrinos

$$\nu = \overline{\nu}$$







Indian Institute of Technology Ropar



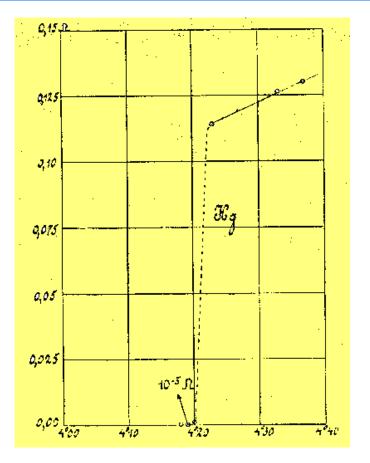
## Superconductivity - discovery



- Liquid Helium (4K) (1908) *Boiling point* 4.22K
- Superconductivity in Hg  $T_C = 4.2K$  (1911)

"Mercury has passed into a new state, which on account of its extraordinary electrical properties may be called the superconducting state"

H. Kamerlingh Onnes 1913 (Nobel price 1913)



Resistivity R=0 below T<sub>C</sub>; (R<10<sup>-23</sup>  $\Omega$ ·cm, 10<sup>18</sup> times smaller than for Cu)



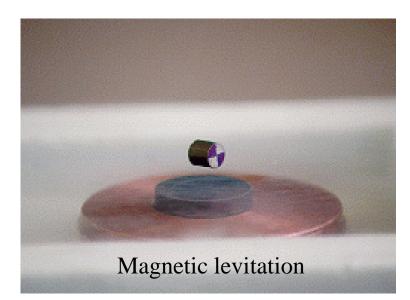


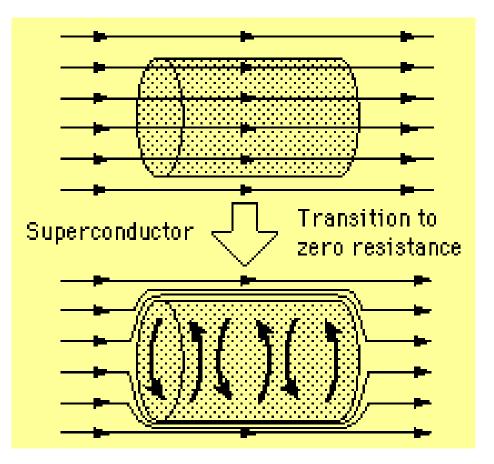
### Meissner - Ochsenfeld - effect

A superconductor is a perfect diamagnet. Superconducting material expels magnetic flux from the interior.

W. Meissner, R. Ochsenfeld (1933)

On the surface of a superconductor  $(T < T_C)$  superconducting current will be induced. This creates a magnetic field compensating the outside one.

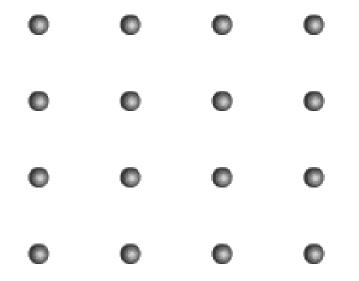






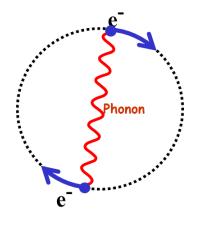


An electron on the way through the lattice interacts with lattice sites (cations). The electron produces **phonon**.



The lattice deformation creates a region of relative positive charge which can attract another electron.

During one phonon oscillation an electron can cover a distance of  $\sim 10^4$ Å. The second electron will be attracted without experiencing the repulsing electrostatic force .



Size of Cooper pair 100 nm Lattice spacing 0.1 - 0.4 nm



### Further discoveries

1911-1986: "Low temperature superconductors" Highest  $T_C=23K$  for Nb<sub>3</sub>Ge

1986 (January): High Temperature Superconductivity  $(LaBa)_2 CuO_4 T_C = 35K$ 

K.A. Müller und G. Bednorz (IBM Rüschlikon) (Nobel price 1987)

1987 (January):  $YBa_2Cu_3O_{7-x}T_C = 93K$ 

1987 (December): Bi-Sr-Ca-Cu-O  $T_C = 110K$ ,

1988 (January): Tl-Ba-Ca-Cu-O  $T_C = 125K$ 

1993: Hg-Ba-Ca-Cu-O  $T_C = 133K$ 

(A. Schilling, H. Ott, ETH Zürich)



Professor Dr. Dr. h. c. mult. Karl Alex Müller (links) und Dr. Johannes Georg Bednorz

Z. Phys. B - Condensed Matter 64, 189-193 (1986)



Possible High  $T_c$  Superconductivity in the Ba-La-Cu-O System

> J.G. Bednorz and K.A. Müller IBM Zürich Research Laboratory, Rüschlikon, Switzerland

Received April 17, 1986

Metallic, oxygen-deficient compounds in the Ba – La – Cu – O system, with the composition Ba Lu<sub>5-x</sub>Cu<sub>5-0</sub> have been prepared in polycrystalline form. Samples with x=1 and 0.75, y>0, annealed below 900 °C under reducing conditions, consist of three phases, one of them a perovskite-like mixed-valent copper compound. Upon cooling, the samples show a linear decrease in resistivity, then an approximately logarithmic increase, interpreted as a beginning of localization. Pinally an abrupt decrease by up to three orders of magnitude occurs, reminiscent of the onset of percolative superconductivity. The highest onset temperature is observed in the 30 K range. It is markedly reduced by high entremt densities. Thus, it results partially from the percolative nature, bute possibly also from 2D superconducting fluctuations of double perovskite layers of one of the phases present.





## How does high T<sub>c</sub> superconductivity work

▶ In 1986, Müller and Bednorz found a ceramic compound that superconducted at 30 K, 13 degrees above the highest previously known superconductor. What was so odd about this is that the material (a compound of lanthanum, barium, copper, and oxygen) was a ceramic – normally an insulator – so people *never expected* such materials might be high  $T_c$  superconductors.



Müller and Bednorz

➢ By substituting yttrium for lanthanum, such compounds exceeded 77 K, the temperature of LN2 (a major milestone since

LN2 is far cheaper than methods of cooling below this temperature). For the first time, concepts such as Maglev trains, superconducting magnets for accelerators, lossless power transmission, etc. became possibilities.



The current record holder is 138 K (for Hg<sub>0.8</sub>Tl<sub>0.2</sub>Ba<sub>2</sub>Ca<sub>2</sub>Cu<sub>3</sub>O<sub>8.33</sub>). How high will transition temperatures reach? Are there yetundiscovered materials that might exceed room temperature (which would completely revolutionize the entire

electronics and power industries)?



Will it ever be possible to understand high  $T_c$  materials on a quantitative level?

