Building a level scheme: γ-decay

✤ Gamma-ray emission is usually the dominant decay mode

Measurements of γ-rays let us deduce:
Energy, Spin (angular distr. / correl.), Parity (polarization), magnetic moment, lifetime (recoil distance, Doppler shift), ...
of the involved nuclear levels.

¹³⁷Cs detected in red: NaI scintillator











γ-decay in a Nutshell

- The photon emission of the nucleus essentially results from a re-ordering of nucleons within the shells.
- * This re-ordering often follows α or β decay, and moves the system into a more energetically favorable state.







γ-decay

 γ -spectroscopy yields some of the most precise knowledge of nuclear structure, as spin, parity and ΔE are all measurable.

Transition rates between initial Ψ_N^* and final Ψ_N' nuclear states, resulting from electromagnetic decay producing a photon with energy E_{γ} can be described by Fermi's Golden rule:

$$\lambda = \frac{2\pi}{\hbar} \left| \left\langle \Psi_{N}^{'} \psi_{\gamma} \right| \mathcal{M}_{em} \left| \Psi_{N}^{*} \right\rangle \right|^{2} \frac{dn_{\gamma}}{dE_{\gamma}}$$

where \mathcal{M}_{em} is the electromagnetic transition operator and dn_{γ}/dE_{γ} is the density of final states. The photon wave function ψ_{γ} and \mathcal{M}_{em} are well known, therefore measurements of λ provide detailed knowledge of nuclear structure.

A γ -decay lifetime is typically 10^{-12} [s] and sometimes even as short as 10^{-19} [s]. However, this time span is an eternity in the life of an excited nucleon. It takes about $4 \cdot 10^{-22}$ [s] for a nucleon to cross the nucleus.



- Level and γ properties
 - •
 - Measured E_{γ} 's \rightarrow level energies Efficiency-corrected γ -peak areas \rightarrow intensities, branching ratios, level populations
 - γ ADs/ACs, γ polarizations, internal conversion \rightarrow multipolarity \rightarrow level spin, parity •

Intensity of EM field (emitted photon) given by Poynting vector, which depends on spherical harmonics $Y_{lm}(\theta, \phi)$, where θ is relative to quantization axis



AD: angular distribution, AC: angular correlation



• Angular distributions

- Quantization axis = spin direction
- May be known event by event

Reaction plane defines spin vector





- Angular distributions
 - Quantization axis = spin direction
 - May be known event by event

Reaction plane defines spin vector



(ϕ is the difference between *lab* angles ϕ_{γ} and ϕ_{p})





(ϕ is the difference between *lab* angles ϕ_{γ} and ϕ_{p})

Measure intensity as a function of angle



- Angular distributions
 - Quantization axis = spin direction
 - May be known event by event
 - ... or it may not! What to do then?

$$W(\theta) = \sum_{k=0}^{L} A_{2k} P_{2k}(\cos \theta)$$

Reaction plane still defines spin vector, even if not determined experimentally





- Angular distributions
 - Quantization axis = spin direction
 - May be known event by event
 - ...or it may not! What to do then?

$$W(\theta) = \sum_{k=0}^{L} A_{2k} P_{2k}(\cos \theta)$$

Pure dipole only has P_2 term, quadrupole adds P_4 , octupole has P_6 , etc.

With M1/E2 mixing, $A_4 \neq 0$; mixing ratio δ^* can be deduced, providing insight into underlying structure

Strictly speaking, AD in singles; "gated AD": $\gamma\gamma$ coincidence with no angle condition on gating γ (clean-up only) \rightarrow if 4π coverage, same AD as in singles

AD: angular distribution, AC: angular correlation



* $\delta^2 = (E2 \text{ intensity})/(M1 \text{ intensity})$



Electric and magnetic dipole fields have opposite parity: Magnetic dipoles have even parity and electric dipole fields have odd parity.

 $\Rightarrow \pi(M\ell) = (-1)^{\ell+1}$ and $\pi(E\ell) = (-1)^{\ell}$







It is possible to describe the angular distribution of the radiation field as a function of the *multipole order* using Legendre polynomials.

- ℓ : The index of radiation
 - 2^{ℓ} : The multipole order of the radiation
- $\ell = 1 \rightarrow Dipole$ $\ell = 2 \rightarrow Quadrupole$ $\ell = 3 \rightarrow Octupole$
- The associated Legendre polynomials $P_{2\ell}(cos(\theta))$ are: For $\ell = 1$: $P_2 = \frac{1}{2}(3 \cdot cos^2(\theta) - 1)$ For $\ell = 2$: $P_4 = \frac{1}{8}(35cos^4(\theta) - 30cos^2(\theta) + 3)$



- The photon is a spin-1 boson
- * Like α-decay and β-decay the emitted γ-ray can carry away units of *angular momentum* ℓ , which has given us different multipolarities for transitions.
- ✤ For orbital angular momentum, we can have values $\ell = 0, 1, 2, 3, \cdots$ that correspond to our multipolarity.
- Therefore, our selection rule is:

$$\left|J_i - J_f\right| \le \ell \le \left|J_i + J_f\right|$$



Characteristics of multipolarity

L	multipolarity	$\pi(\mathrm{E}\ell) / \pi(\mathrm{M}\ell)$	angular distribution	$\ell = 1$
1	dipole	-1 / +1		X10
2	quadrupole	+1 / -1		X _{1,±1}
3	octupole	-1 / +1		
4	hexadecapole	+1 / -1		
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$$E_{\gamma} = E_i - E_f$$
$$\left|I_i - I_f\right| \le \ell \le I_i + I_f$$
$$\Delta \pi(E\ell) = (-1)^{\ell}$$
$$\Delta \pi(M\ell) = (-1)^{\ell+1}$$







 $|2-0| \le \ell \le 2+0$

Here $\Delta I = 2$ and $\ell = 2$ this is a stretched transition







 $|3-2| \leq \ell \leq 3+2$

Here $\Delta I = 1$ but $\ell = 1,2,3,4,5$ and the transition can be a mix of 5 multipolarities







Electromagnetic transitions:

 $\Delta \pi (electric) = (-1)^{\ell}$ $\Delta \pi (magnetic) = (-1)^{\ell+1}$

Λπ	yes	E1	M2	E3	M 4
$\Delta \pi$	no	M 1	E2	M3	E4







 $|2-0| \leq \ell \leq 2+0$

 $\ell = 2$ and no change in parity

Λπ					
	no	M 1	E2	M3	E4







 $|3-2| \leq \ell \leq 3+2$

Here $\Delta I = 1$ but $\ell = 1, 2, 3, 4, 5$

Δπ	yes	E1	M2	E3	M 4

mixed E1,M2,E3,M4,E5





The basics of the situation



 $|3-2| \leq \ell \leq 3+2$

Here $\Delta I = 1$ but $\ell = 1, 2, 3, 4, 5$



mixed M1,E2,M3,E4,M5





 $3^+ \rightarrow 2^+$: mixed M1,E2,M3,E4,M5

 $3^+ \rightarrow 2^-$: mixed E1,M2,E3,M4,E5

In general only the lowest 2 multipoles compete

and (for reasons we will see later)

 ℓ + 1 multipole generally only competes if it is electric:

 $3^+ \rightarrow 2^+$: mixed M1/E2 $3^+ \rightarrow 2^-$: almost pure E1 (very little M2 admixture)





Characteristics of multipolarity

L	multipolarity	$\pi(\mathrm{E}\ell) / \pi(\mathrm{M}\ell)$	angular distribution	$\ell = 1$	<i>ℓ</i> =2
1	dipole	-1 / +1		X10	X_{2,0}
2	quadrupole	+1 / -1			$X_{2,\pm 1}$ $X_{2,\pm 2}$
3	octupole	-1 / +1			
4	hexadecapole	+1 / -1			
÷				and the second s	

parity: electric multipoles $\pi(E\ell) = (-1)^{\ell}$, magnetic multipoles $\pi(M\ell) = (-1)^{\ell+1}$

The power radiated is proportional to:

$$P(\sigma\ell) \propto \frac{2(\ell+1) \cdot c}{\varepsilon_0 \cdot \ell \cdot [(2\ell+1)!!]^2} \left(\frac{\omega}{c}\right)^{2\ell+2} |\mathcal{M}(\sigma\ell)|^2$$

where σ means either E or M and $\mathcal{M}(\sigma \ell)$ is the E or M multipole moment of the appropriate kind.





$$T(E1; I_i \to I_f) = 1.590 \ 10^{17} \ E_{\gamma}^3 \ B(E1; I_i \to I_f)$$

$$T(E2; I_i \to I_f) = 1.225 \ 10^{13} \ E_{\gamma}^5 \ B(E2; I_i \to I_f)$$

$$T(E3; I_i \to I_f) = 5.709 \ 10^8 \ E_{\gamma}^7 \ B(E3; I_i \to I_f)$$

$$T(E4; I_i \to I_f) = 1.697 \ 10^4 \ E_{\gamma}^9 \ B(E4; I_i \to I_f)$$

$$T(M1; I_i \to I_f) = 1.758 \ 10^{13} \ E_{\gamma}^3 \ B(M1; I_i \to I_f)$$

$$T(M2; I_i \to I_f) = 1.355 \ 10^7 \ E_{\gamma}^5 \ B(M2; I_i \to I_f)$$

$$T(M3; I_i \to I_f) = 6.313 \ 10^0 \ E_{\gamma}^7 \ B(M3; I_i \to I_f)$$

where $E_{\gamma} = E_i - E_f$ is the energy of the emitted γ quantum in MeV (E_i , E_f are the nuclear level energies, respectively), and the reduced transition probabilities B(E ℓ) in units of $e^2(barn)^{\ell}$ and B(M ℓ) in units of $\mu_N^2 = (e\hbar/2m_Nc)^2 (fm)^{2\ell-2}$





$$B(E\lambda;I_i \to I_{gs}) = \frac{(1.2)^{2\lambda}}{4\pi} (\frac{3}{\lambda+3})^2 A^{2\lambda/3} e^2 (fm)^{2\lambda}$$

$$B(M\lambda; I_i \to I_{gs}) = \frac{10}{\pi} (1.2)^{2\lambda - 2} (\frac{3}{\lambda + 3})^2 A^{(2\lambda - 2)/3} \ \mu_N^2 (fm)^{2\lambda - 2}$$

For the first few values of λ , the Weisskopf estimates are

$$\begin{split} B(E1;I_i \to I_{gs}) &= 6.446 \ 10^{-4} \ A^{2/3} \ e^2(barn) \\ B(E2;I_i \to I_{gs}) &= 5.940 \ 10^{-6} \ A^{4/3} \ e^2(barn)^2 \\ B(E3;I_i \to I_{gs}) &= 5.940 \ 10^{-8} \ A^2 \ e^2(barn)^3 \\ B(E4;I_i \to I_{gs}) &= 6.285 \ 10^{-10} \ A^{8/3} \ e^2(barn)^4 \\ B(M1;I_i \to I_{gs}) &= 1.790 \ (\frac{e\hbar}{2Mc})^2 \end{split}$$



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Measuring g-factors





Recoil-in-vacuum (RIV) technique

- Nucleus produced in reaction in oriented state
- Recoil exits target into vacuum \rightarrow spin is deoriented by hyperfine interactions
- Deorientation dependent on g-factor





Measuring g-factors

- **R**ecoil **In V**acuum in ¹³⁴Te
 - Coulex of ¹³⁴Te RIB (and also ¹³⁰Te SIB)
 - Simultaneously measure B(E2) and g-factor for 2^+
 - Observe attenuated ACs \rightarrow relate to $G_k \rightarrow g$ -factor









GSĬ



Measuring g-factors

During the collision, the projectile can be firstly excited reaching a certain excited state (I_f).

Before a γ -decay, the scattered projectile and the target recoil nucleus exit the target into the vacuum highly excited and ionized.



At that time, the scattered projectile decays rapidly by gamma emissions. However, the strong fluctuations of the hyperfine fields can lead to a de-orientation of the nuclear state.



♦ $g(2^+) = 0$ → unperturbed p-γ angular correlation

$$G_{k} = [1 + k \cdot (k + 1) \cdot \lambda \cdot \tau(I)]^{-1}$$
$$\lambda = 1/3 \cdot \tau_{c} \cdot \omega^{2}$$
$$\omega = g \cdot \frac{H \cdot \mu_{N}}{\hbar}$$

A. Abragam & R.V. Pound, Phys. Rev. 92 (1953) 943 $G_4 = \frac{3 \cdot G_2}{(10 - 7 \cdot G_2)}$



• Angular distributions

- Quantization axis = spin direction
- May be known event by event
- ... or it may not! What to do then?
- Spins not aligned indefinitely \rightarrow AD attenuated through hyperfine interactions

If there is no spin orientation, all substates contribute equally and $W(\theta) = \text{constant}$

What to do if alignment is lost (or was never there), e.g. for γ 's below isomers?

AD: angular distribution, AC: angular correlation



• Angular correlations

- Coincidence technique
- Detect first $\gamma \rightarrow$ defines quantization axis
- Detect second $\gamma \rightarrow$ determine relative angle

$$W(\psi) = \sum_{k=0}^{L} A_{2k} P_{2k}(\cos \psi)$$

Same form as for AD, different coefficients. *L* limit determined by lowest multipole.

GSÍ













Imagine the situation of an M1 decay between two states, the initial one has J^{π} value of 1⁺ and the final one a J^{π} of 0⁺

The initial $J^{\pi}=1^+$ state has 3 degenerate magnetic substates which differ by the magnetic quantum numbers m of ± 1 and 0.

The final $J^{\pi}=0^+$ state has a single magnetic substate with m=0.

When the substates of $J^{\pi}=1^+$ state decay, the γ -rays emitted have different angular patterns.





The basics of the situation



So the total distribution is
$$W_{M1} = \frac{1}{3}W_{M1,\Delta m=1} + \frac{1}{3}W_{M1,\Delta m=0} + \frac{1}{3}W_{M1,\Delta m=-1}$$

= $\frac{1}{8\pi}(1 + \cos^2\theta + \sin^2\theta) = \frac{1}{4\pi}$

no angular dependence







Let's imagine we have two γ -rays which follow immediately after each other in the level scheme.

If we measure γ_1 or γ_2 in singles, then the distribution will be isotropic (same intensity at all angles) ... there is no preferred direction of emission

Now imagine that we measure γ_1 and γ_2 in coincidence. We say that measuring γ_1 causes the intermediate state to be aligned. We define the z-direction as the direction of γ_1

The angular distribution of the emission of γ_2 then depends on the spin/parities of the states involved and on the multipolarity of the transition.





A simple example:



Hence, for γ_2 we only see the m=±1 to m=0 part of the distribution i.e. we see that the intensity measured as a function of angle (relative to γ_1) follows a 1 + $cos^2\theta$ distribution.





General formula



In general, the γ -ray intensity varies as:

$$W(\theta) = \sum_{k_{even}} A_k(\gamma_1) A_k(\gamma_2) Q_k(\gamma_1) Q_k(\gamma_2) P_k(\cos\theta)$$

where

 θ is the relative angle between the two γ -rays

 Q_k accounts for the fact that we do not have point detectors

 A_k depends on the details of the transition and the spins of the level

$$P_0 = 1 \quad P_2 = \frac{1}{2}(3 \cdot \cos^2(\theta) - 1) \quad P_4 = \frac{1}{8}(35\cos^4(\theta) - 30\cos^2(\theta) + 3)$$

$$W(\theta) = 1 + a_2 \cos^2 \theta + a_4 \cos^4 \theta$$





R.D. Evans, The Atomic Nucleus



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$I_1(\ell_1)$	$I_2(\ell_2)$	I ₃	a ₂	a ₄
0 (1)	1 (1)	0	1	0
1 (1)	1 (1)	0	-1/3	0
1 (2)	1 (1)	0	-1/3	0
2 (1)	1 (1)	0	1/13	0
3 (2)	1 (1)	0	-3/29	0
0 (2)	2 (2)	0	-3	4
1 (1)	2 (2)	0	-1/3	0
2 (1)	2 (2)	0	3/7	0
2 (2)	2 (2)	0	-15/13	16/13
3 (2)	2 (2)	0	-3/29	0
4 (2)	2 (2)	0	1/8	1/24



General formula



In general, the γ -ray intensity varies as:

$$W(\theta) = \sum_{k_{even}} A_k(\gamma_1) A_k(\gamma_2) Q_k(\gamma_1) Q_k(\gamma_2) P_k(\cos\theta)$$

where θ is the relative angle between the two γ -rays Q_k accounts for the fact that we do not have point detectors A_k depends on the details of the transition and the spins of the level

$$P_0 = 1 \quad P_2 = \frac{1}{2}(3 \cdot \cos^2(\theta) - 1) \quad P_4 = \frac{1}{8}(35\cos^4(\theta) - 30\cos^2(\theta) + 3)$$

$$A_{k}(\gamma_{1}) = \frac{F_{k}(J_{2}J_{1}\ell,\ell) - 2 \cdot \delta \cdot F_{k}(J_{2}J_{1}\ell,\ell+1) + \delta^{2} \cdot F_{k}(J_{2}J_{1}\ell+1,\ell+1)}{1 + \delta^{2}}$$
$$A_{k}(\gamma_{2}) = \frac{F_{k}(J_{2}J_{3}L,L) - 2 \cdot \delta \cdot F_{k}(J_{2}J_{3}L,L+1) + \delta^{2} \cdot F_{k}(J_{2}J_{3}L+1,L+1)}{1 + \delta^{2}}$$

Ferentz-Rosenzweig coefficients

$$F_k(LL'I_1I_2) = (-1)^{I_1+I_2+1}\sqrt{2k+1}\sqrt{2L+1}\sqrt{2L'+1}\sqrt{2I_2+1} \begin{pmatrix} L & L' & k \\ 1 & -1 & 0 \end{pmatrix} \begin{cases} L & L' & k \\ I_1 & I_1 & I_2 \end{cases}$$

https://griffincollaboration.github.io/AngularCorrelationUtility/





A special case:

 $^{195}_{~78} Pt (n,\gamma) ^{196}_{~78} Pt$








Many arrays are designed symmetrically, so the range of possible angles is reduced.

Therefore one measures a Directional Correlation from Oriented Nuclei (DCO ratio) In the simplest case, if you have an array with detectors at 35^0 and 90^0 . Gate on 90^0 detector, measure coincident intensities in

- other 90⁰ detectors
- 35⁰ detectors

Take the ratio and compare with calculations ... can usually separate quadrupoles from dipoles but cannot measure mixing ratios















Angular distribution

In heavy-ion fusion-evaporation reactions, the compound nuclei have their spin aligned in a plane perpendicular to the beam axis:

 $\vec{\ell} = \vec{r} \times \vec{p}$

Depending on the number and type of particles 'boiled off' before a γ -ray is emitted, transitions are emitted from oriented nuclei and therefore their intensity shows an angular dependence.



$$W(\theta) = A_0 \left(1 + \frac{A_2}{A_0} \cdot B_2 \cdot Q_2 \cdot P_2 \left(\cos\theta \right) + \frac{A_4}{A_0} \cdot B_4 \cdot Q_4 \cdot P_4 \left(\cos\theta \right) \right)$$

where A_k , Q_k and P_k are as before and B_k contains information about the alignment of the state





Angular distribution



Measure: the γ -ray yield as a function of θ



Building a level scheme

• ADs/ACs → multipole order L, but do not directly distinguish E vs M

- Pure $\Delta I=1$ E1 looks the same as pure $\Delta I=1$ M1, e.g.
- Nonzero mixing ratio may change that, as M1/E2 far more likely than E1/M2
- For more direct determination of E or M: measure **polarization**

Measure Compton scattering within and perpendicular to reaction plane

$$P \propto \frac{aN_{\perp} - N_{\parallel}}{aN_{\perp} + N_{\parallel}}$$

(a corrects for different intrinsic count rates)





Linear polarization





A segmented detector can be used to measure the linear polarization which can be used to distinguish between magnetic (M) and electric (E) character of radiation of the same multipolarity.



The Compton scattering cross section is larger in the direction perpendicular to the electrical field vector of the radiation. Define experimental asymmetry as: $A = \frac{N_{90} - N_0}{N_{90} + N_0}$

where N_{90} and N_0 are the intensities of scattered photons perpendicular and parallel to the reaction plane.

The experimental linear polarization P=A/Q where Q is the polarization sensitivity of the detector

Q~13% at 1 MeV





Linear polarization



Klein-Nishina formula:

$$\frac{d\sigma_c}{d\Omega} = \frac{r_0^2}{2} \left(\frac{E_{\gamma\prime}}{E_{\gamma}}\right)^2 \cdot \left\{\frac{E_{\gamma}}{E_{\gamma\prime}} + \frac{E_{\gamma\prime}}{E_{\gamma}} - 2sin^2\theta \cdot cos^2\varphi\right\}$$

Maximum polarization at $\theta = 90^{\circ}$



Proof of Principle





Linear polarization



Plot P against the angular distribution information to uniquely define the multipolarity.

Data from Eurogam



Building a level scheme

• ADs/ACs → multipole order L, but do not directly distinguish E vs M

- Pure $\Delta I=1$ E1, e.g., looks the same as pure $\Delta I=1$ M1
- Nonzero mixing ratio may change that, as M1/E2 far more likely than E1/M2
- For more direct determination of E or M: measure polarization (e.g. Clover detectors)
- ...or internal-conversion electrons



Instead of γ emission, energy transferred to atomic electron, which is ejected

Nuclear level decay via internal conversion (IC)

- first observed over 100 y ago [van Baeyer and Hahn, Physik. Z. 11, 488 (1 910)] *
- ubiquitous process
- usable as analysis tool
 - (cf. BrIcc, http://bricc.anu.edu.au/)

* nucleus discovered in 1911!

AD: angular distribution, AC: angular correlation





Internal conversion



Energetics of CE-decay (i=K, L, M,....) $E_i = E_f + E_{ce,i} + E_{BE,i}$

 γ - and CE-decays are independent; transition probability ($\lambda \sim$ Intensity)

$$\lambda_{\mathrm{T}} = \lambda_{\gamma} + \lambda_{\mathrm{CE}} = \lambda_{\gamma} + \lambda_{\mathrm{K}} + \lambda_{\mathrm{L}} + \lambda_{\mathrm{M}} \dots$$

Conversion coefficient

$$\alpha_i = \frac{\lambda_{CE,i}}{\lambda_{\gamma}}$$





Internal conversion



• For an electromagnetic transition internal conversion can occur instead of emission of gamma radiation. In this case the transition energy $Q = E_{\gamma}$ will be transferred to an electron of the atomic shell.

 $T_e = E_{\gamma} - B_e$

 T_e : kinetic energy of the electron

 B_e : binding energy of the electron

internal conversion is important for:

- heavy nuclei ~ Z^3
- high multipolarities $E\ell$ or $M\ell$
- small transition energies

$$\alpha_k(El) \propto Z^3 \left(\frac{L}{L+1}\right) \left(\frac{2m_e c^2}{E}\right)^{L+5/2}$$



Electron spectroscopy





Doppler shift correction for projectile:

$$T_e^* = \gamma \cdot T_e \cdot \left\{ 1 - \beta_1 \cdot \sqrt{1 + 2m_e c^2 / T_e} \cdot \cos\theta_{e1} \right\} + m_e c^2 \cdot (\gamma - 1)$$

 $cos\theta_{e1} = cos\vartheta_1 cos\vartheta_e + sin\vartheta_1 sin\vartheta_e cos(\varphi_e - \varphi_1)$

resolution of the spectrometer		$\left(\frac{\Delta p}{p}\right)_{e}/\%$	
as calculated for a point source		0.4	
scattering in the target	(i)	0.004	
beam optics	(ii)	0.11	
evaporation of neutrons	(iii)	0.09	
energy loss in the target	(iv)	0.31	*
energy straggling of the projectiles	(v)	0.006	
quadratic sum experimental resolution		0.53 0.56	%

Mini Orange setup for conversion electron spectroscopy









Coulomb excitation experiment





Surface oscillations in deformed nuclei



Building a level scheme

- E0 decays
 - Most transitions have γ vs e⁻ competition, but one case that "never" has a γ : $0^+ \rightarrow 0^+$
 - Transition would be E0 (monopole moment = charge), which cannot radiate to points external to nucleus





E0/E2 branching ratio

$$\frac{Y_e(E0)}{Y_{\gamma}(E2)} = \frac{\Omega_K [s^{-1}]}{2.56 \cdot 10^9 \cdot A^{4/3} \cdot E_{\gamma}^{-5} [MeV]} \cdot \frac{B(E0; I \to I')}{B(E2; I \to I')}$$

= 14
$$\beta^2$$
 for $2_\beta \rightarrow 2$



 Ω_{K} : conversion probability electronic factor

D.A. Bell et al.; Can. J. of Phys. 48 (1970),2542



Building a level scheme

• E0 decays

- Most transitions have γ vs e⁻ competition, but one case that "never" has a γ : $0^+ \rightarrow 0^+$
- Transition would be E0 (monopole moment = charge), which cannot radiate to points external to nucleus

Why "never" and not *never*?

If $E > 2m_ec^2 = 1.022$ MeV, internal pair (e⁺e⁻) production is possible.

e⁺ subsequently annihilates with another e⁻, yielding 511-keV γ 's.



• Level and γ properties

- Measured E_{γ} 's \rightarrow level energies Efficiency-corrected γ peak areas \rightarrow intensities, branching ratios, level populations
- γ ADs/ACs, γ polarizations, internal conversion \rightarrow multipolarity \rightarrow level spin, parity
- Relative γ times \rightarrow level half-lives, B(XL)'s

Reduced transition rates
$$B(XL) \propto E_{\gamma}^{-(2L+1)} T_{1/2,\gamma}^{-1}$$

can be a sensitive probe of matrix elements connecting initial and final states:

$$B(XL; I_i \to I_f) = \frac{1}{2I_i + 1} \langle \psi_f || XL || \psi_i \rangle^2$$



Excited-state half-lives range from ~ 10^{-23} s (>10-MeV resonances) to > 4.5x10¹⁶ y (^{180m}Ta). That's over 47 orders of magnitude!

Most excited states have sub-ns half-lives; why are some exceptionally long?

- Low transition energy
- Large spin difference
- Large difference in underlying configurations
- *K*-forbiddenness (change in spin projection on symmetry axis $\Delta K > L$)

No single approach to measuring half-lives will apply to all cases.



- Observe exponential decay for half-lives of ~tens of ps or longer
 - Measure γ-ray time relative to a fast START signal
 - START with preceding β decay in e.g. fast plastic
 - ... or START with RF (clock tick marking reaction)
 - ... or START with preceding γ in e.g. Ge ($T_{1/2} > ns$) or LaBr₃ (sub-ns)
 - Get half-life from slope of exponential



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- Observe exponential decay for half-lives of ~tens of ps or longer
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 - ... or START with RF (clock tick marking reaction)
 - ... or START with preceding γ in e.g. Ge (T_{1/2} > ns) or LaBr₃ (sub-ns)
 - Get half-life from slope of exponential...or from centroid shift





- For ~ps half-lives, instrumentation not fast enough
 - Exploit physics of in-beam measurement: Doppler shifts

Brief Doppler diversion...



Doppler effect



Ions lose energy while traversing a medium \rightarrow v/c decreases with time.

By modelling the energy loss, time-dependent Doppler shift determined \rightarrow Doppler-shift attenuation and (differential) plunger techniques



- For ~ps half-lives, instrumentation not fast enough
 - Exploit physics of in-beam measurement: Doppler shifts
 - Recoil-distance method (plunger)
- Produce nucleus of interest with initial velocity v₀ via Coulex or KO reactions
- Observe Doppler-shifted γ's emitted between target and degrader
- Ion slows in the degrader to velocity v_d < v₀
- Observe Doppler-shifted γ's emitted after degrader
- Increase distance between target and degrader; change in relative peak intensities depends on $T_{1/2}$
- $T_{1/2} \sim 10-100s$ of ps can be measured



Figure: adapted from K. Starosta

Energy



Recoil distance method





• Double-degrader plunger



Iwasaki et al., PRL112, 142502 (2014).



Doppler Shift Attenuation Method



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- For ~ps half-lives, instrumentation not fast enough
 - Exploit physics of in-beam measurement: Doppler shifts
 - Recoil-distance method (plunger)
 - Doppler-broadened line shapes



• Sensitive to sub-ps $T_{1/2}$'s



Chiara *et al.*, PRC**61**, 034318 (2000)



Lorentz transformation



Efficiency versus resolution



With a source at rest, the intrinsic resolution of the detector can be reached;

efficiency decreases with the increasing detector-source distance.



Energy resolution

The major factors affecting the final energy resolution (FWHM) at a particular energy are as follows:

$$\Delta E_{\gamma}^{final} = \left(\Delta E_{Int}^2 + \Delta \theta_{det}^2 + \Delta \theta_N^2 + \Delta v^2\right)^{1/2}$$

- ΔE_{Int} The intrinsic resolution of the detector system. It includes contributions from the detector itself and the electronic components used to process the signal.
- $\Delta \theta_{det}$ The Doppler broadening arising from the opening angle of the detectors
- $\Delta \theta_N$ The Doppler broadening arising from the angular spread of the recoils in the target
- Δv The Doppler broadening arising from the velocity (energy) variation of the excited nucleus



Special relativity

Lorentz transformation:



□Consider the space-time point

- in a given frame S: (t, x, y, z)
- and in a (moving) frame S': (t', x', y', z')
- 1) S' moves with a constant velocity v along z-axis

Space-time Lorentz transformation $S \leftarrow \rightarrow S'$:

 $\frac{S \Rightarrow S'}{x' = x} \qquad \frac{S' \Rightarrow S}{x = x'}$ $\frac{y' = y}{\sqrt{1 - v^2}} \qquad \frac{y' = \gamma(z - vt)}{t' = \gamma(t - vz)} \qquad z = \gamma(z' + vt)$

Note: units c=1

□ Consider the 4-momentum: • in a given frame S: $p \equiv (E, p) = (E, p_x, p_y, p_z)$ • in the (moving) frame S': $p' \equiv (E', \vec{p}') = (E', p'_x, p'_y, p'_z)$ Lorentz transformation for 4-momentum S←→S': $p'_x = p_x, p'_y = p_y$ $p'_z = \gamma(p_z - \nu E)$ $E' = \gamma(E - \nu p_z)$

$$t' = \gamma(t - \frac{v}{c^2}z)$$
$$t = \gamma(t' + \frac{v}{c^2}z)$$
$$v_z = |\vec{v}| = v$$



Lorentz transformation



total energy:

$$E^* = \gamma \cdot E - \gamma \cdot \nu \cdot P \cdot \cos\theta$$

with

$$E = \sqrt{(mc^2)^2 + (Pc)^2}$$

 E^* , P^* total energy and momentum in the rest system E, P total energy and momentum in the laboratory system

Doppler formula for zero-mass particle (photon): E=Pc $E^* = \gamma \cdot E - \gamma \cdot \beta \cdot E \cdot cos\theta$ $E^* = \gamma \cdot E(1 - \beta \cdot cos\theta)$



Hendrik Lorentz

E. Byckling, K. Kajantie J. Wiley & Sons London



Doppler effect



180
Doppler broadening and position resolution

$$E_{\gamma 0} = E_{\gamma} \frac{1 - \beta \cdot \cos \vartheta_{\gamma}}{\sqrt{1 - \beta^2}} \quad (\beta, \vartheta_p = 0^0, \vartheta_{\gamma} \text{ and } E_{\gamma} \text{ in lab - frame})$$

$$\left(\frac{\Delta E_{\gamma 0}}{E_{\gamma 0}}\right)^{2} = \left(\frac{\beta \cdot \sin \theta_{\gamma}}{1 - \beta \cdot \cos \theta_{\gamma}}\right)^{2} (\Delta \theta_{\gamma})^{2} + \left(\frac{\beta - \cos \theta_{\gamma}}{(1 - \beta^{2}) \cdot (1 - \beta \cdot \cos \theta_{\gamma})}\right)^{2} (\Delta \beta)^{2} + \left(\frac{1}{E_{\gamma}}\right)^{2} (\Delta E_{\gamma})^{2}$$
Position resolution
Angular resolution
$$A \theta$$

$$\theta$$
beam
projectile



12 10 12 2 Days

Doppler broadening (opening angle of detector)



Doppler broadening (velocity variation)





Experimental arrangement



experimental problem:

Doppler broadening due to finite size of Ge-detector $\frac{\Delta E}{E} \sim 1\% \quad \text{for} \quad \Delta \vartheta_{\gamma} = 20^0 \quad \beta_1 \cong 10\%$

For projectile excitation:

$$E^* = \gamma \cdot E \cdot (1 - \beta_1 \cdot \cos \theta_{\gamma 1})$$
 Doppler shift with

$$\cos\theta_{\gamma 1} = \cos\vartheta_1 \cos\vartheta_\gamma + \sin\vartheta_1 \sin\vartheta_\gamma \cos(\varphi_\gamma - \varphi_1)$$

 $\Delta E \cong E^* \cdot \beta_1 \cdot \sin \theta_{\gamma 1} \cdot \Delta \theta_{\gamma 1} \qquad \text{Doppler broadening}$

Inelastic heavy-ion scattering







Lorentz transformation



Contraction of the solid angle element in the laboratory system

$$\frac{d\Omega}{d\Omega^*} = \left\{\frac{E^*}{E}\right\}^2$$

with

$$E^* = \gamma \cdot E \cdot (1 - \beta \cdot \cos\theta)$$
 Doppler formula



Experimental arrangement (electron detection)



Doppler broadening

 $\Delta \vartheta_e = 20^0$ target – Mini-Orange: 19 cm Mini-Orange – Si detector: 6 cm

For projectile excitation:

$$T_e^* = \gamma \cdot T_e \cdot \left\{ 1 - \beta_1 \cdot \sqrt{1 + 2m_e \, c^2 / T_e} \cdot \cos\theta_{e1} \right\} + m_e c^2 \cdot (\gamma - 1)$$

with

$$cos\theta_{e1} = cos\vartheta_1 cos\vartheta_e + sin\vartheta_1 sin\vartheta_e cos(\varphi_e - \varphi_1)$$



Lorentz transformation





Segmented detectors







