# Gamma-ray spectroscopy II

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# Outline

First lecture

- $\succ$  Properties of  $\gamma$ -ray transitions
- Fusion-evaporation reactions
- Germanium detector arrays
- Coincidence technique
- > Nuclear deformations
- Rotation of deformed nuclei
- > Pair alignment
- Superdeformed nuclei
- > Hyperdeformed nuclei
- Triaxiality and wobbling

#### Second lecture

- Angular distribution
- Linear polarization
- Jacobi shape transition
- Charged-particle detectors
- Neutron detectors
- Prompt proton decay
- Recoil-decay tagging
- Rotation and deformation alignment

#### Third lecture

- Spectroscopy of transfermium nuclei
- Conversion-electron spectroscopy
- Quadrupole moments and transition rates
- Recoil-distance method
- Doppler shift attenuation method
- Fractional Doppler shift method
- Magnetic moments
- Perturbed angular distribution
- Magnetic Rotation
- Shears Effect

#### Fourth lecture

- Fast fragmentation beams
- Isomer spectroscopy after fragmentation
- ➢ E0 transitions
- > Shape coexistence
- ➤ Two-level mixing
- Coulomb excitation
- Reorientation effect
- ISOL technique
- Low-energy Coulomb excitation of <sup>74</sup>Kr
- Relativistic Coulomb excitation of 58Cr
- ➢ Gamma-ray tracking
- > AGATA





 $\Rightarrow$  compare experimental  $R_{DCO}$  with values for dipole-dipole, dipole-quadrupole, quadrupole-quadrupole cascades

 $\Rightarrow$  often sufficient for spin assignments

$$R_{DCO} = \frac{I(\gamma_1, \theta_1; \gamma_2, \theta_2)}{I(\gamma_1, \theta_2; \gamma_2, \theta_1)}$$

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#### **Angular distribution**



Alignment of angular momentum after fusion-evaporation reaction:

$$W(\vartheta) = 1 + \sum_{k} A_{k} P_{k}(\cos \vartheta)$$



The coefficients  $A_k$  depend on

the multipolarity L
 the mixing parameter δ
 the population width σ

$$A_{k}(L,L',I_{f},I_{i}) = \rho_{k}(I_{i})\frac{1}{1+\delta^{2}} \Big[F_{k}(L,L,I_{f},I_{i}) + 2\delta F_{k}(L,L',I_{f},I_{i}) + 2\delta^{2}F_{k}(L',L',I_{f},I_{i})\Big]$$

$$\rho_k(I_i) = \sqrt{2I_i + 1} \sum_{m=-I}^{+I} (-1)^{I_i - m} \langle I_i m I_i - m | k 0 \rangle P(m)$$

Ferentz-Rosenzweig coefficients

$$F_{k}(L,L',I_{f},I_{i}) = (-1)^{I_{f}+I_{i}-1} \sqrt{(2L+1)(2L'+1)(2I_{i}+1)(2k+1)} \begin{pmatrix} L & L' & k \\ 1 & -1 & 0 \end{pmatrix} \begin{cases} L & L' & k \\ I_{i} & I_{i} & I_{f} \end{cases}$$
Clebsch-
Bacah

 $\sigma/I$  is approximately constant (for a given reaction). Normalize to transition with known multipolarity, e.g. 2<sup>+</sup> $\rightarrow$ 0<sup>+</sup>

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# Example: Angular distribution with EUROBALL



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#### Measuring the mixing parameter $\delta$



We know  $\sigma/I$  and have assigned I<sup> $\pi$ </sup> For wobbling bands, we expect  $\Delta I$ =1 E2 inter-band transitions.  $\Rightarrow$  L=1, L'=2, large  $\delta$ 

Two possible solutions



Angular distribution cannot distinguish between the two.  $\Rightarrow$  measure the linear polarization to establish electric or magnetic character.



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# Linear polarization



linear polarization: fixed direction of electric field vector E



Compton scattering is sensitive to linear polarization: Klein-Nishina formula

$$\frac{d\sigma}{d\Omega} = \frac{r_0^2}{2} \frac{\omega'^2}{\omega^2} \left( \frac{\omega'}{\omega} + \frac{\omega}{\omega'} - 2\sin^2\theta \cos^2\zeta \right)$$



electric transitions appear positive, magnetic transitions negative







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#### Polarization measurement in <sup>163</sup>Lu



	Eγ	$A = \frac{N(90^{\circ}) - N(0^{\circ})}{N(90^{\circ}) + N(0^{\circ})}$		
E2	579	0.10 ± 0.03		
	697	0.13 ± 0.03	positive	
	386	0.06 ± 0.05		
	534	0.05 ± 0.04		
M1	349	-0.11 ± 0.05	negative	
inter-band	607	0.05 ± 0.05		
	626	0.12 ± 0.05		
	643	0.11 ± 0.05	positive	
	659	0.17 ± 0.09	$\Rightarrow$ electric	
	673	0.18 ± 0.09		

Confirmation of the wobbling mode in <sup>163</sup>Lu through combined angular distribution and linear polarization measurement.



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MacLaurin shapes

What happens if we spin a liquid drop ?

It becomes oblate !

Jupiter: ➤ T = 9 h 50 min ➤ polar / equatorial axis ~ 15/16

MacLaurin shape after C. MacLaurin (1698-1746)

But what if we spin really fast ?

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# Jacobi shapes

The equilibrium shape changes abruptly to a very elongated triaxial shape rotating about its shortest axis.

piece of moon rock from Apollo mission



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#### The Jacobi shape transition in nuclei



Carl Gustav Jacob Jacobi (1804 - 1851) discovered transition from oblate to triaxial shapes in the context of rotating, idealized, incompressible gravitating masses in 1834.

In 1961 Beringer and Knox suggested a similar transition in the case of atomic nuclei, idealized as incompressible, uniformly charged, liquid drops endowed with surface tension.



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What is the signature of a Jacobi transition in nuclei?

- sharp decrease of frequency with increasing angular momentum (giant backbend of the moment of inertia)
- Frequency of collective rotation is related to the E2 γ-ray energy:  $\hbar ω = \frac{1}{2} E_γ$
- many rotational bands at high spin quasi-continuous transitions



- measure the energy of the quasi-continuous 'E2 bump' as a function of angular momentum
- series of experiments with Gammasphere

<sup>48</sup>Ca + <sup>50</sup>Ti @ 200 MeV
<sup>48</sup>Ca + <sup>64</sup>Ni @ 207 MeV
<sup>48</sup>Ca + <sup>96</sup>Zr @ 207 MeV
<sup>48</sup>Ca + <sup>124</sup>Sn @ 215 MeV

- as neutron rich as possible:
- $\Rightarrow$  higher fission barrier



# Measuring angular momentum with Gammasphere







108 Ge detectors 6 x 108 = 648 BGO detectors

increase in false veto signals reduced Ge efficiency but very high granularity

K = number of hits = fold  $M = \gamma$  rays emitted = multiplicity (from response function) J = initial angular momentum (from angular distribution)



# The E2 bump





# K measures the angular momentum E2 bump measures rotational frequency

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# **Comparison to liquid drop calculations**



D. Ward et al., Phys. Rev. C 66, 024317 (2002)

two modifications:

- Iower effective moment of inertia at low spin due to pairing
- no collective rotation about axially symmetric (MacLaurin) shapes in nuclei, instead, collective rotations are associated with (mostly) prolate shapes
  - $\rightarrow$  no sharp transition caused by breaking of axial symmetry, but smooth transition

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## <sup>40</sup>Ca+<sup>40</sup>Ca @ 167 MeV

The nucleus of interest is often only weakly populated compared to a large background of other nuclei.

Additional sensitivity from:➤ charged-particle detectors➤ neutron detectors

- recoil detectors
- tagging techniques

neutrons are deeply				761/	<sup>78</sup> Zr 2n	<sup>79</sup> Zr 1n	<sup>80</sup> Zr
particle evaporation favored despite Coulomb barrier			p3n	// Y p2n 0.18	pn	<sup>79</sup> Ү 1р	
			<sup>75</sup> Sr αn	<sup>76</sup> Sr 2p2n 4.06	<sup>77</sup> Sr 2pn 2.95	<sup>78</sup> Sr 2p	
		<sup>72</sup> Rb αp3n	<sup>73</sup> Rb αp2n 0.01	<sup>74</sup> Rb αpn 2.49	<sup>75</sup> Rb αp 5.35	<sup>76</sup> Rb 3pn 101	<sup>77</sup> Rb 3p 2.31
	<sup>70</sup> Kr 2α2n	<sup>71</sup> Kr 2αn 0.18	<sup>72</sup> Kr 2α 0.37	<sup>73</sup> Kr α2pn 52	<sup>74</sup> Kr α2p 20.3	<sup>75</sup> Kr 4pn 132	<sup>76</sup> Kr 4p 74.2
	<sup>69</sup> Br 2αp2n	<sup>70</sup> Br 2αpn 6.82	<sup>71</sup> Br 2αp 11.4	<sup>72</sup> Br α3pn 38.2	<sup>73</sup> Br α3p 128	<sup>74</sup> Br 5pn 3.23	<sup>75</sup> Br 5p 68.2
<sup>67</sup> Se 3αn	<sup>68</sup> Se 3α 5.07	<sup>69</sup> Se 2α2pn 1.57	<sup>70</sup> Se 2α2p 94	<sup>71</sup> Se α4pn 0.46	<sup>72</sup> Se α4p 102	<sup>73</sup> Se 6pn	<sup>74</sup> Se 6p 1.57
<sup>66</sup> As 3αpn	<sup>67</sup> As 3αp 15	<sup>68</sup> As 2α3pn	<sup>69</sup> As 2α3p 35.6	<sup>70</sup> As α5pn	<sup>71</sup> As α5p 2.03	<sup>72</sup> As 7pn	<sup>73</sup> As 7p
<sup>65</sup> Ge 3α2pn	<sup>66</sup> Ge 3α2p 8.3	<sup>67</sup> Ge 2α4pn	<sup>68</sup> Ge 2α4p 0.37	<sup>69</sup> Ge α6pn	cross sections in mb		



<sup>64</sup>Ge

4α 1.48



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# **Neutron detection**



Gammasphere with Microball and Neutron shell Eu (Washington University, St. Louis)







- used to study nuclei near N=Z line
  - isospin symmetry
  - proton-neutron pairing
  - > shape coexistence
  - astrophysical rapid-proton capture process
- neutrons are separated from γ rays by time of flight and pulse shapes (zero-crossing time)
- difficult to distinguish two-neutron hit from scattering



# Prompt proton decay in <sup>58</sup>Cu



≥ <sup>28</sup>Si(<sup>36</sup>Ar,αpn)<sup>58</sup>Cu

D. Rudolph et al., Phys. Rev. Lett. 80, 3018 (1998) Eur. Phys. J. A 14, 137 (2002)

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# **Recoil decay tagging**





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#### <sup>189</sup>Bi level schemes



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#### Systematics of the neutron-deficient Bi isotopes

